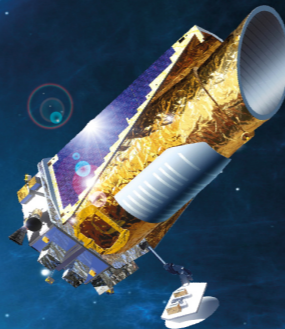


# 3<sup>rd</sup> COROT Symposium 7<sup>th</sup> KASC Meeting



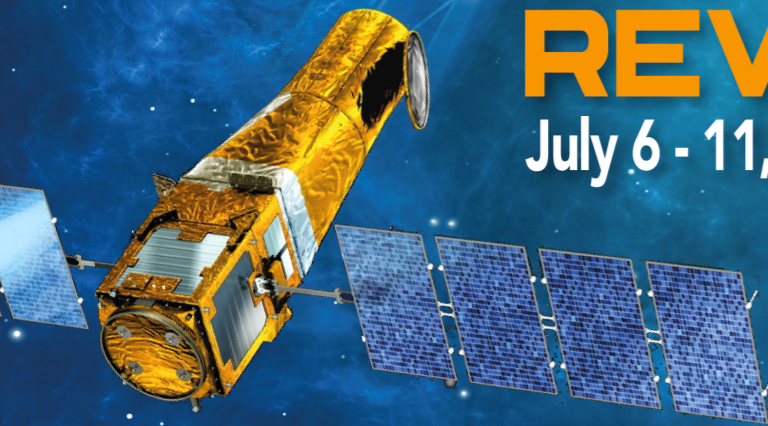
EXTRASOLAR PLANETS  
STELLAR STRUCTURE



ACTIVITY

# THE SPACE PHOTOMETRY REVOLUTION

July 6 - 11, 2014 - Toulouse, FRANCE



### Scientific Organizing Committee

C. Aerts - A. Baglin  
J. Ballot (Deputy Chair) - N. M. Batalha  
J. Christensen-Dalsgaard - P. De Cat  
H. Deeg - M. Deleuil  
R.A. García (Chair) - E. Janot-Pacheco  
L. Kiss - H. Kjeldsen (Co-Chair)  
D.W. Kurtz - D.W. Latham  
T.S. Metcalfe - E. Michel (Co-Chair)  
M.J.P.F.G. Monteiro - E. Poretti  
H. Rauer - D. Stello  
G. Vauclair - W.W. Weiss

### Local Organizing Committee

M.A. Albouy - J. Ballot (Chair)  
S. Charpinet - S. Deheuvels (co-chair)  
R. A. García - D. Granat  
N. Le Gal - F. Lignières  
G. Mirouh - P. Vert

ASTEROSEISMOLOGY  
ROTATION

[corot3-kasc7.sciencesconf.org](http://corot3-kasc7.sciencesconf.org)

# concluding remarks

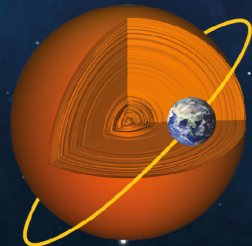
Andrea Miglio

School of Physics and Astronomy  
University of Birmingham

and



STELLAR ASTROPHYSICS CENTRE  
University of Aarhus





# the space photometry revolution

## Definitions and Characteristics of Revolutions:

### Dictionary:

- a forcible overthrow of a government or social order in favor of a new system
- a dramatic and wide-ranging change in the way something works or is organized or in people's ideas about it

### Vladimir Lenin:

- A revolution is impossible without a revolutionary situation.
- It is impossible to predict the time and progress of revolution. It is governed by its own more or less mysterious laws.

Surely the workings of ESA and NASA qualify as mysterious, at best — in an alternate Universe . . .

# the space photometry revolution

## Definitions and Characteristics of Revolutions:

### Dictionary:

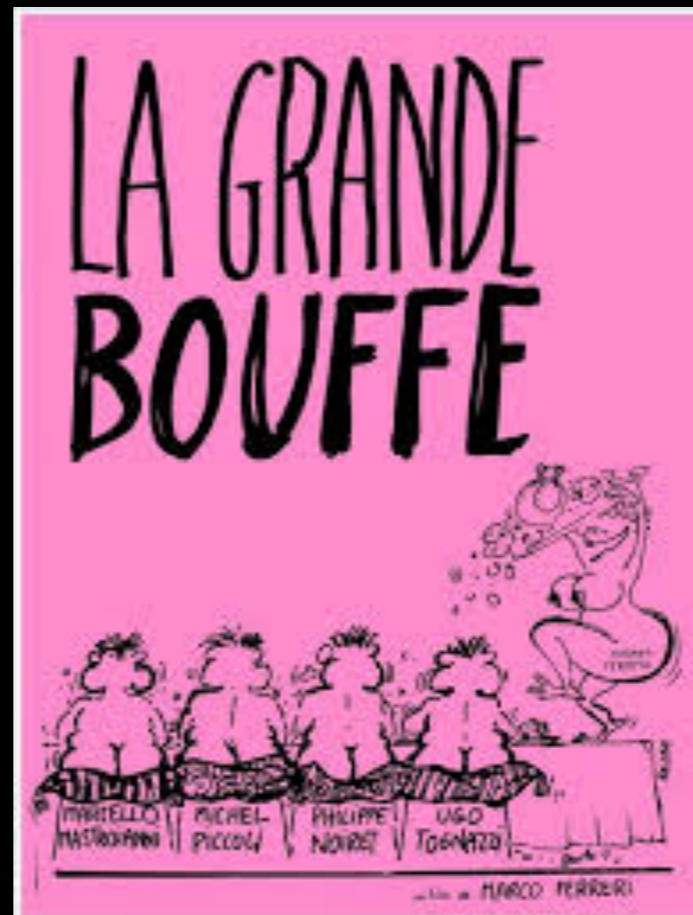
- a forcible overthrow of a government or social order in favor of a new system
- a dramatic and wide-ranging change in the way something works or is organized or in people's ideas about it

### Vladimir Lenin:

- A revolution is impossible without a revolutionary situation.
- It is impossible to predict the time and progress of revolution. It is governed by its own more or less mysterious laws.

Surely the workings of ESA and NASA qualify as mysterious, at best — in an alternate Universe . . .

solar-like oscillators:



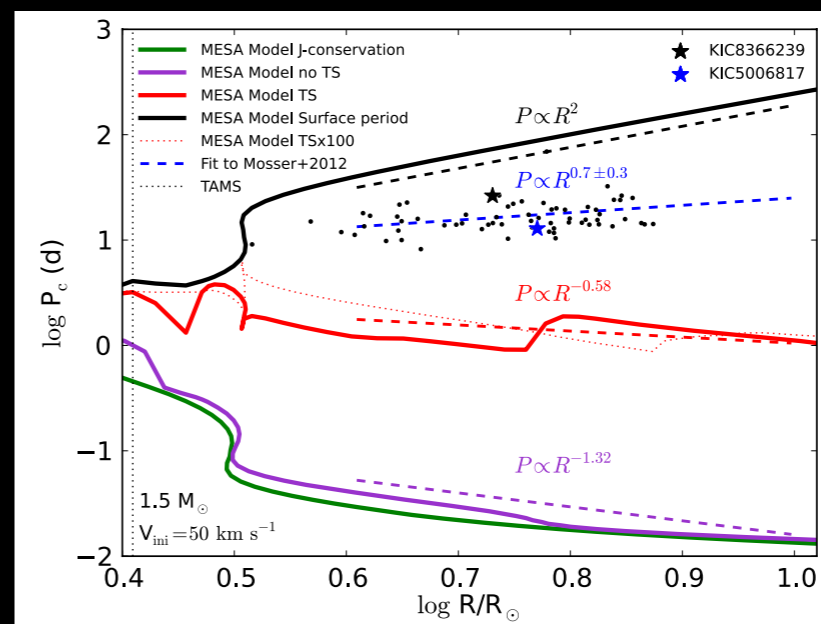
is it over?

# solar-like oscillators:

starting to stress-test  
models

transport of angular  
momentum

Matteo Cantiello



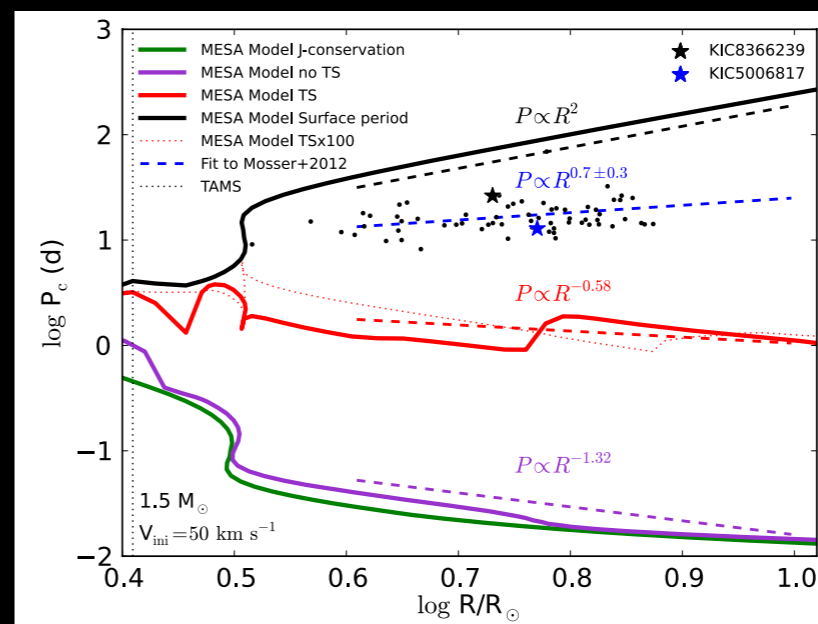
# solar-like oscillators:

starting to stress-test  
models

transport of angular  
momentum

mixing

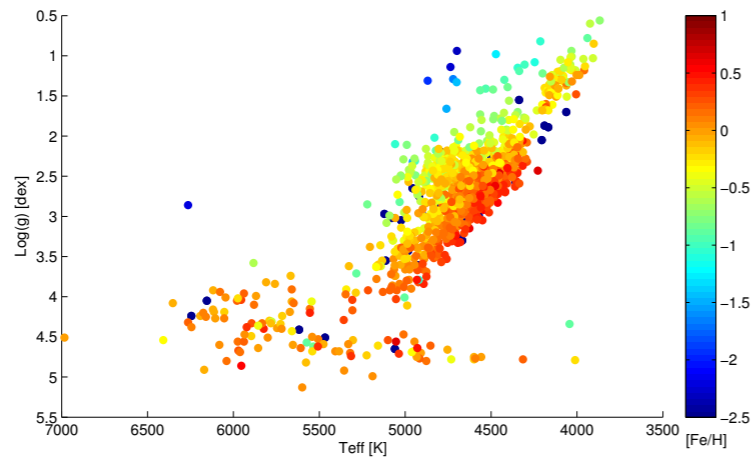
Matteo Cantiello



Sebastien Deheuvels

## CoRoT stars in GES-DR2

First results for the CoRoT-GES DR2



Marica Valentini

The GES-CoRoT collaboration

Marica Valentini

Josefina Montalbán

Marc Pinsonneault

Peter De Cat

## *Strömgen survey for Asteroseismology and Galactic Archaeology*

**Luca Casagrande**

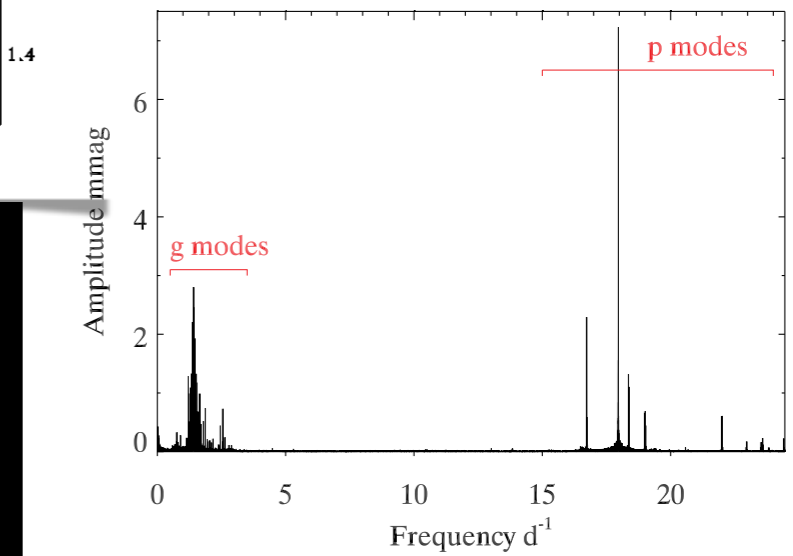
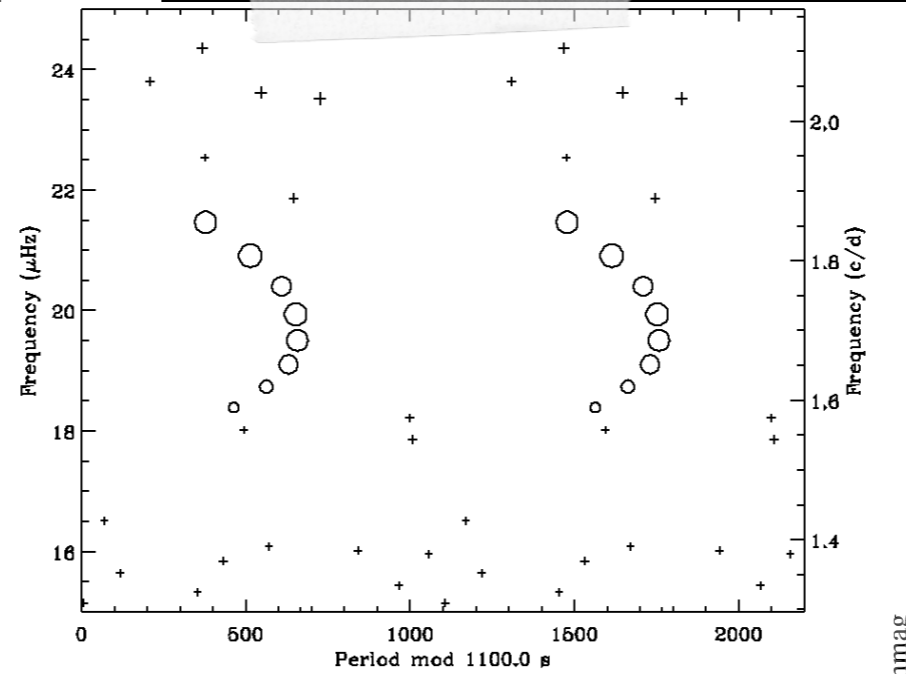
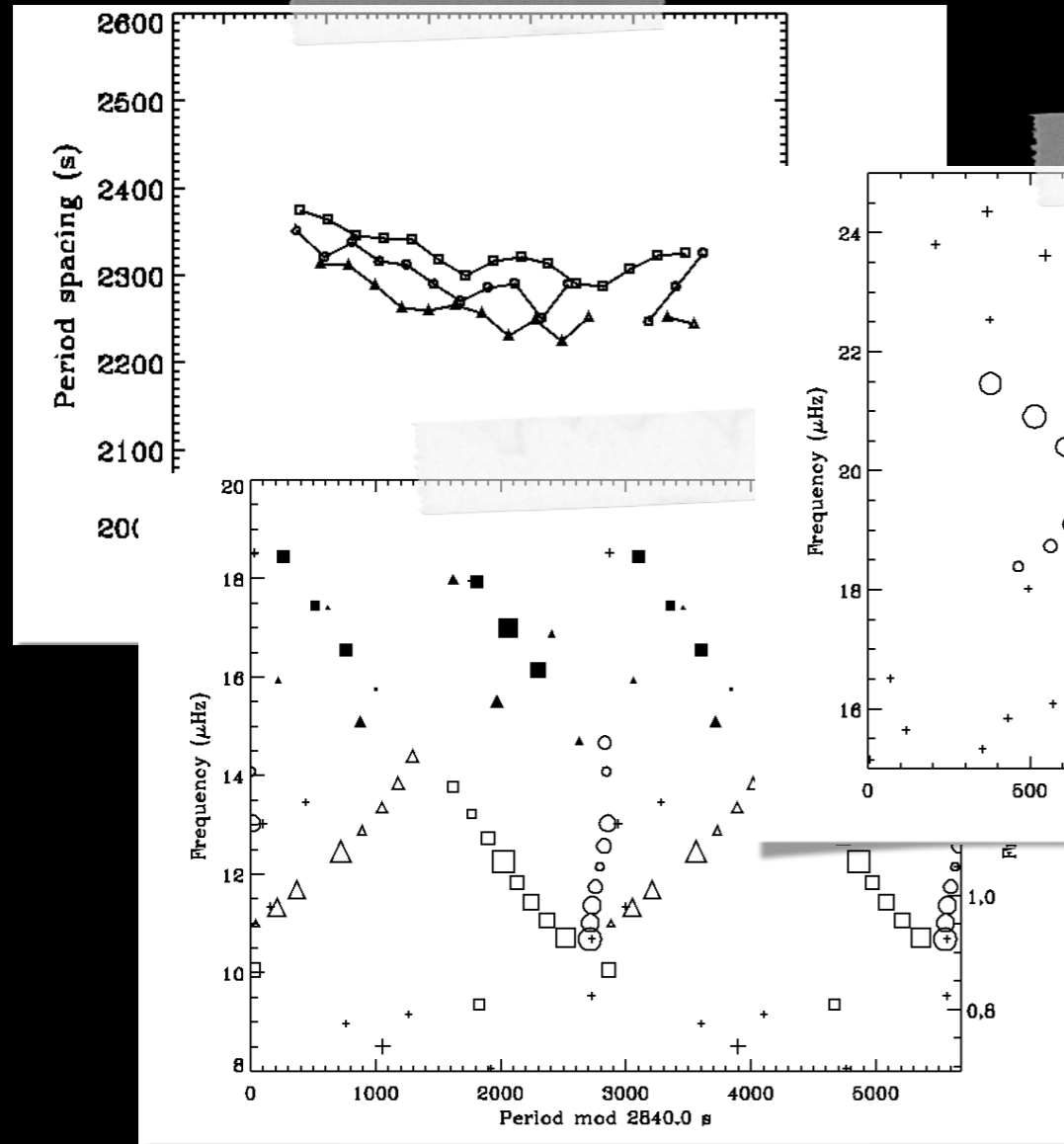
*V. Silva Aguirre, D. Stello, D. Huber, A. Serenelli, M.N. Lund, S. Cassisi, A. Dotter, A.P. Milone, S. Hodgkin, A.F. Marino, A. Pietrinferni, M. Asplund, S. Feltzing, C. Flynn, F. Grundahl, P.E. Nissen, R. Schönrich, K.J. Schlesinger, W. Wang*



Luca Casagrande



# classical pulsators strike back?



Tim Bedding, Don Kurtz

# classical pulsators strike back?

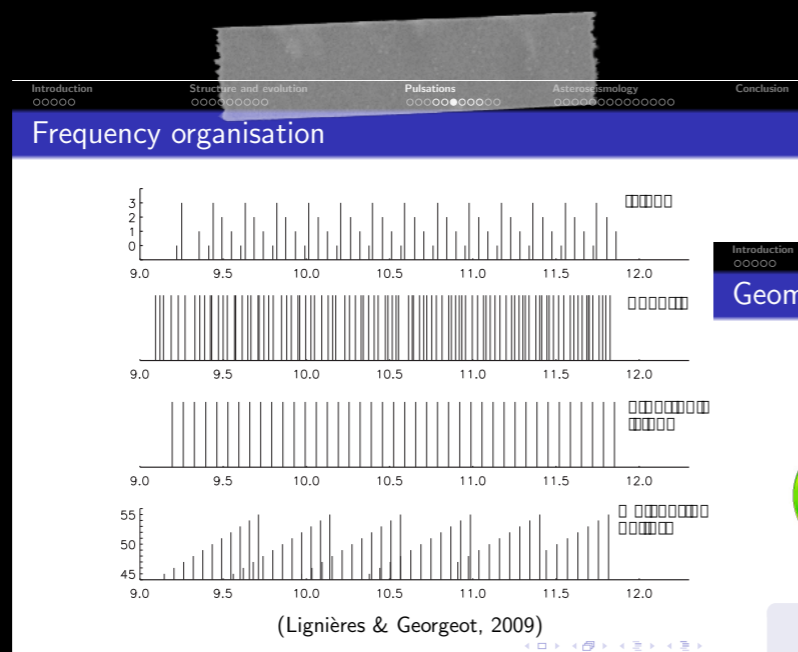
$\delta$  Scuti stars? encouraging results Konstanze Zwintz

# classical pulsators strike back?

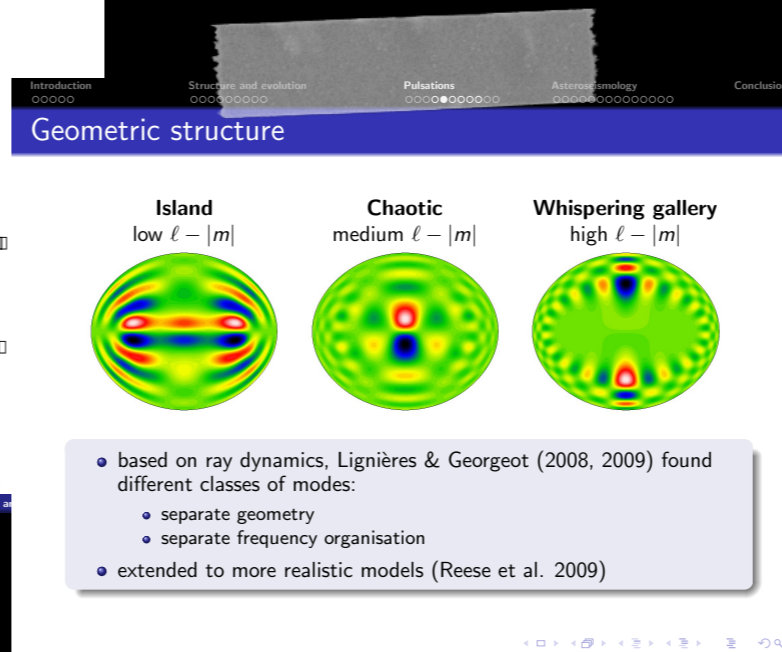
$\delta$  Scuti stars? encouraging results Konstanze Zwintz

fast rotators  
(rule not the exception)

... more sophisticated  
interpretation needed



(Lignières & Georgot, 2009)



asteroseismology and exoplanets:  
from revolution to symbiosis

# asteroseismology and exoplanets: from revolution to symbiosis

## *Symbiosis*

Interaction between two different organisms living in close physical association

# asteroseismology and exoplanets: from revolution to symbiosis

## *Symbiosis*

Interaction between two different organisms living in close physical association

## *A Parasitism*

A parasitic relationship is one in which one member of the association benefits while the other is harmed.

# astroseismology and exoplanets: from revolution to symbiosis

## *Symbiosis*

Interaction between two different organisms living in close physical association

### *A Parasitism*

A parasitic relationship is one in which one member of the association benefits while the other is harmed.

### *B Commensalism*

An association between two organisms in which one benefits and the other derives neither benefit nor harm.

# astroseismology and exoplanets: from revolution to symbiosis

## *Symbiosis*

Interaction between two different organisms living in close physical association

### *A Parasitism*

A parasitic relationship is one in which one member of the association benefits while the other is harmed.

### *B Commensalism*

An association between two organisms in which one benefits and the other derives neither benefit nor harm.

### *C Mutualism*

Mutualism is any relationship between individuals of different species where both individuals benefit.

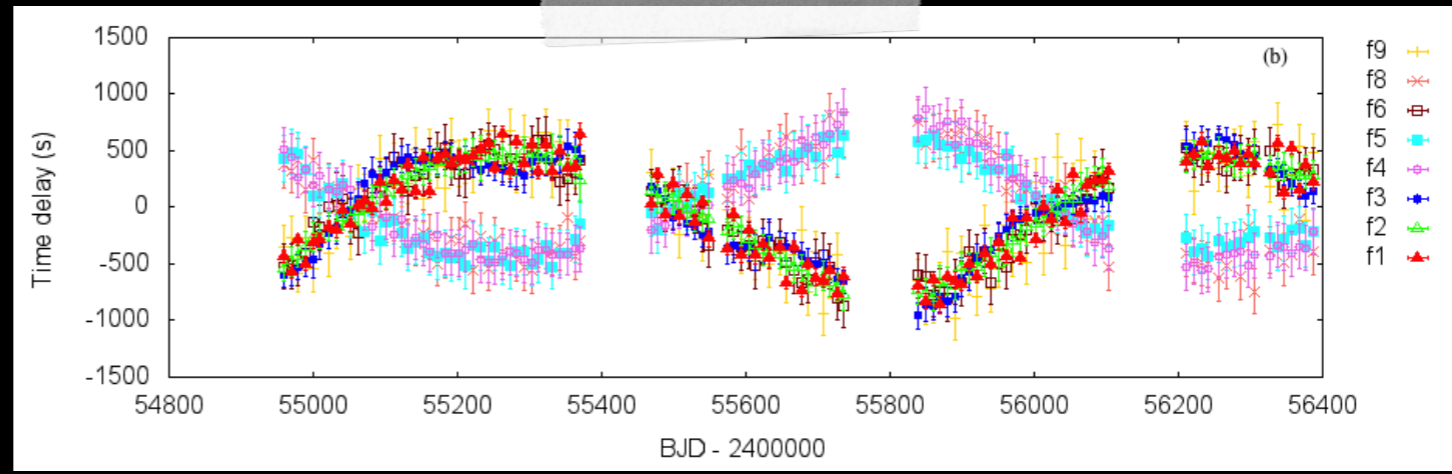




# from revolution to symbiosis

## light-time effects

pulsation phase modulation



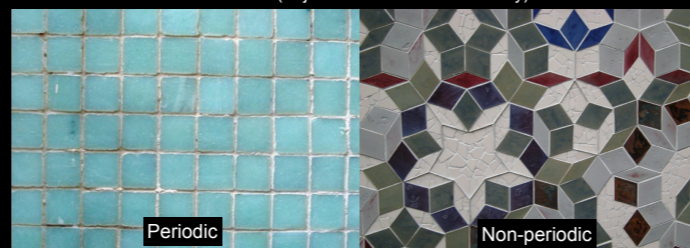
Simon Murphy PB2



### Transit timing variations in Kepler's data

Aviv Ofir

And  
Stefan Dreizler (IAG), Carolina von-Essen (Aarhus University), Tamás Borkovits (Baja Astronomical observatory)



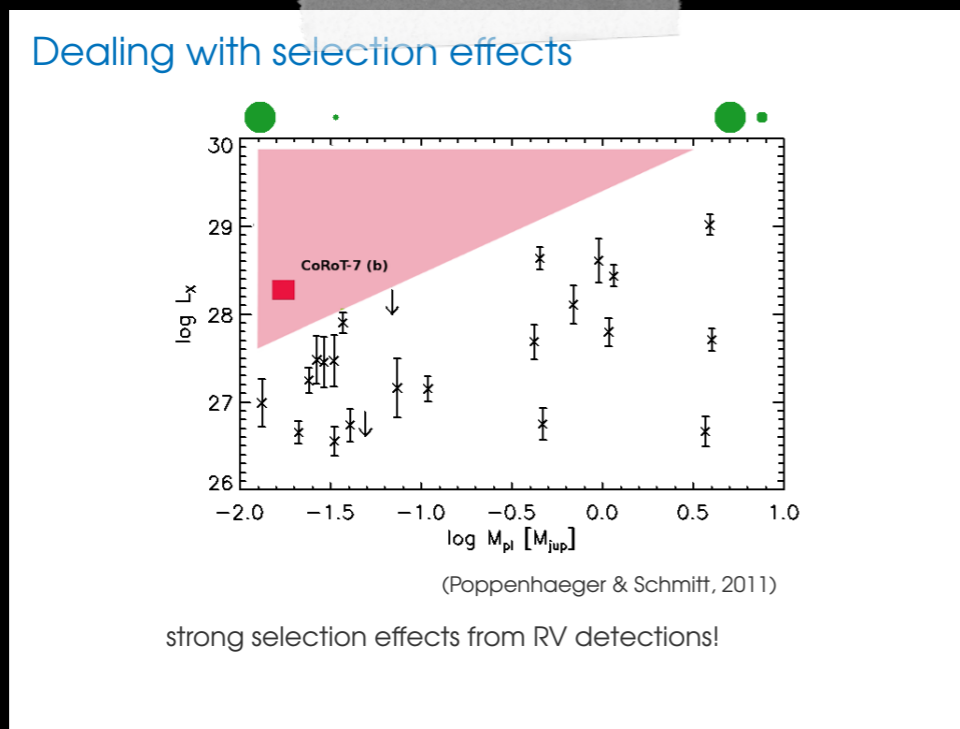
TTVs

Aviv Ofir

# from revolution to symbiosis

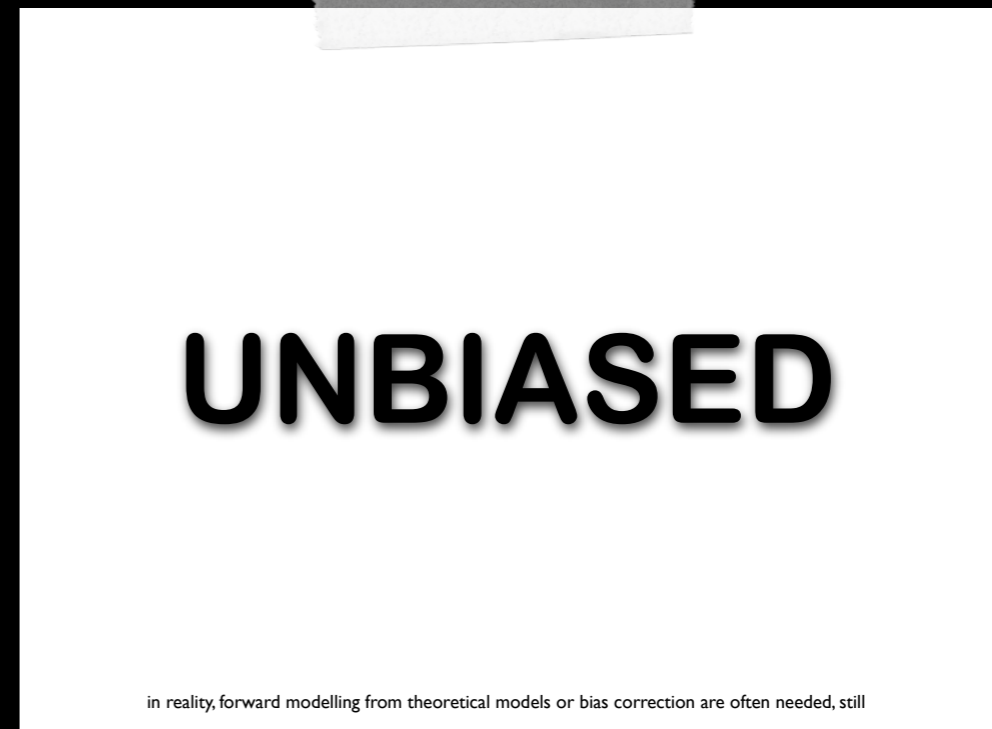
large samples: selection effects and biases

populations of exoplanets



Katja Poppenhaeger

stellar populations



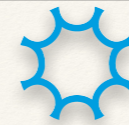
Luca Casagrande

# from revolution to symbiosis

asteroseismic constraints on the host star

Victor Silva Aguirre's review talk

## Summary



- ❖ Virtuous circle: seismology and exoplanets
- ❖ Characterise ~100 KOIs
- ❖ Accurate masses, radii, and ages
- ❖ Facilitate studies of planetary formation and evolution
- ❖ Future looks bright: K2, TESS, PLATO...

# planets

# stars

## Adiabatic interior (fully convective): revisiting the standard picture?

### Reduced heat transport in planetary interiors:

(Stevenson & Salpeter 1977; Stevenson 1979; Chabrier & Baraffe 2007)

- **Idea:** reduced heat transport in planetary interior due to **molecular weight gradient**

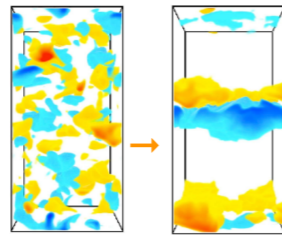
Presence of  $\nabla_{\mu}$  ---> Stabilizing effect against convection

$$\nabla_{ad} > \nabla_T + \nabla_{\mu} \chi_{\mu} / \chi_T \quad (\text{Ledoux criterion})$$

⇒ « layered convection » : system of convective layers + thin diffusive layers

(double diffusive convection or semiconvection)

Layers formation are **observed** in oceans (Pr = 7) and laboratory experiments



### 3D numerical simulations:

⇒ Layers can form in low-Pr (< 1) double diffusive convection (Rosenblum et al. 2011)

Isabelle Baraffe

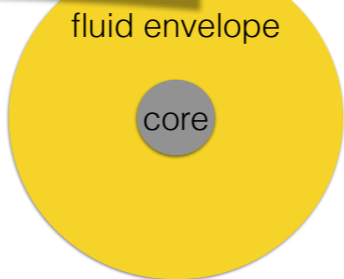
## Evolution with CEPAM

$$\begin{cases} \frac{\partial P}{\partial m} = -\frac{Gm}{4\pi r^4} \\ \frac{\partial T}{\partial m} = \left(\frac{\partial P}{\partial m}\right) \frac{T}{P} \nabla_T, \\ \frac{\partial r}{\partial m} = \frac{1}{4\pi r^2 \rho}, \\ \frac{\partial L}{\partial m} = \epsilon - T \frac{\partial S}{\partial t}, \end{cases}$$

$$\rho = \rho(P, T, \{X_i\}); \quad S = S(P, T, \{X_i\})$$

$$\begin{aligned} m = 0 &\rightarrow r = L = 0 \\ m = M &\rightarrow P = P_{\text{phot}}(g, L) \\ &\quad T = T_{\text{phot}}(g, L) \end{aligned}$$

Guillot & Morel (1995)



Mass

Radius

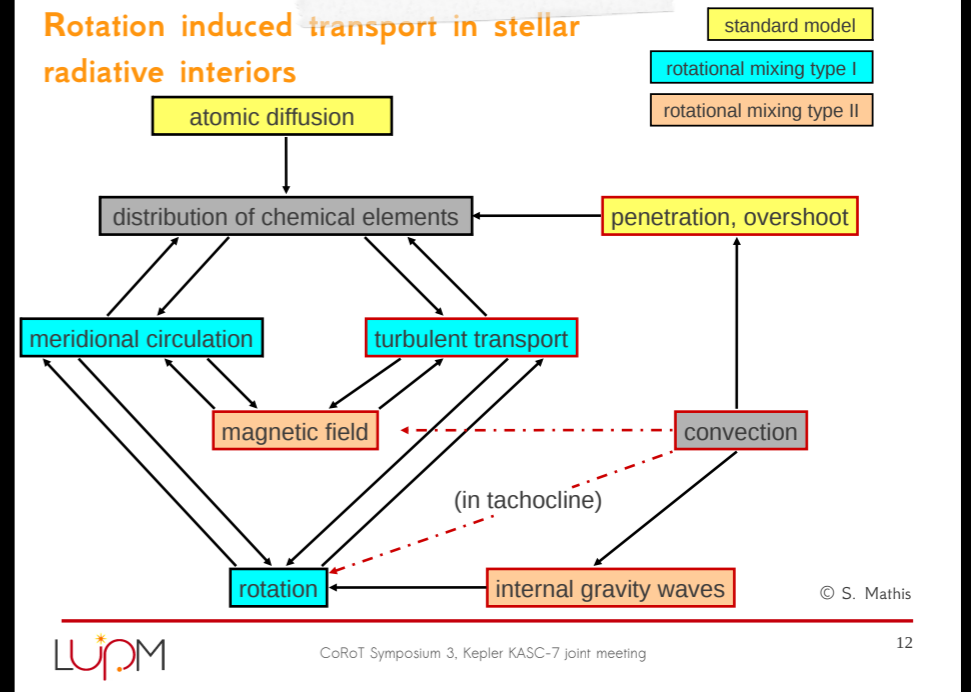
Luminosity

Atmospheric  
T-P profile

Atmospheric  
composition

Rotation rate,  
gravity field

## Rotation induced transport in stellar radiative interiors



Ana Palacios

Mathieu Havel

# planets

# stars

## Adiabatic interior (fully convective): revisiting the standard picture?

### Reduced heat transport in planetary interiors:

(Stevenson & Salpeter 1977; Stevenson 1979; Chabrier & Baraffe 2007)

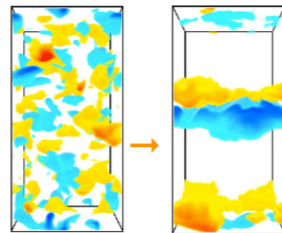
- **Idea:** reduced heat transport in planetary interior due to **molecular weight gradient**

Presence of  $\nabla_{\mu}$  ---> Stabilizing effect against convection

$$\nabla_{ad} > \nabla_T + \nabla_{\mu} \chi_{\mu} / \chi_T \quad (\text{Ledoux criterion})$$

⇒ « layered convection » : system of convective layers + thin diffusive layers  
(double diffusive convection or semiconvection)

Layers formation are **observed** in oceans (Pr = 7) and laboratory experiments



### 3D numerical simulations:

⇒ Layers can form in low-Pr (< 1) double diffusive convection (Rosenblum et al. 2011)

Isabelle Baraffe

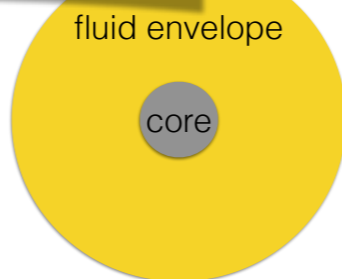
## Evolution with CEPAM

$$\begin{cases} \frac{\partial P}{\partial m} = -\frac{Gm}{4\pi r^4} \\ \frac{\partial T}{\partial m} = \left(\frac{\partial P}{\partial m}\right) \frac{T}{P} \nabla_T, \\ \frac{\partial r}{\partial m} = \frac{1}{4\pi r^2 \rho}, \\ \frac{\partial L}{\partial m} = \epsilon - T \frac{\partial S}{\partial t}, \end{cases}$$

$$\rho = \rho(P, T, \{X_i\}); \quad S = S(P, T, \{X_i\})$$

$$\begin{aligned} m = 0 &\rightarrow r = L = 0 \\ m = M &\rightarrow P = P_{\text{phot}}(g, L) \\ &\quad T = T_{\text{phot}}(g, L) \end{aligned}$$

Guillot & Morel (1995)



Mass

Radius

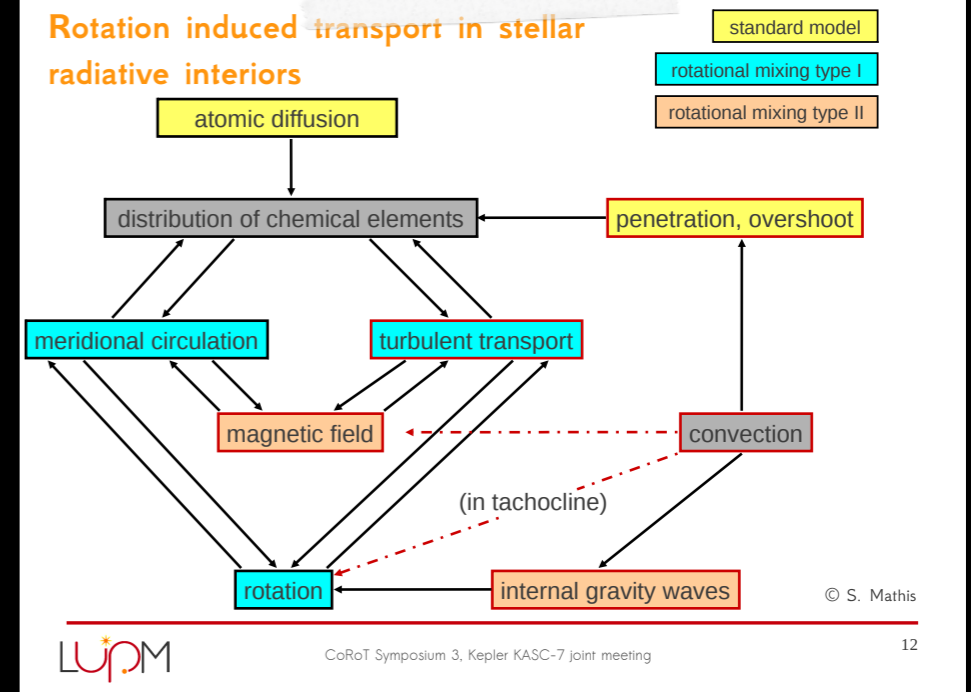
Luminosity

Atmospheric  
T-P profile

Atmospheric  
composition

Rotation rate,  
gravity field

## Rotation induced transport in stellar radiative interiors



Ana Palacios

Mathieu Havel

+ tidal and magnetic interactions

Stephane Mathis

# stellar ages accurate to 10%

“need better stellar physics”  
challenge for the next decade(s)



more realistic stellar models

*Twinkle twinkle little star how I wonder  
what you are:*  
Towards Age/Rotation/Magnetic  
activity relation with seismology

Savita Mathur  
Space Science Institute  
Boulder (USA)

10/07/14 CoRoT3-KASC7 conference 1

Savita Mathur

Jennifer Van Saders



PLATO 2.0 science conference

3-5 December 2014  
Catania, Italy

while historically, symbiosis has received less attention than other interactions such as predation or competition, it is increasingly recognised as an important selective force behind evolution.

*wikipedia*