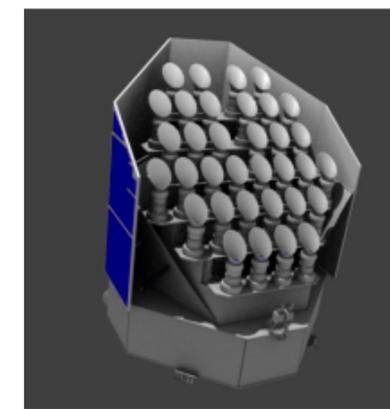
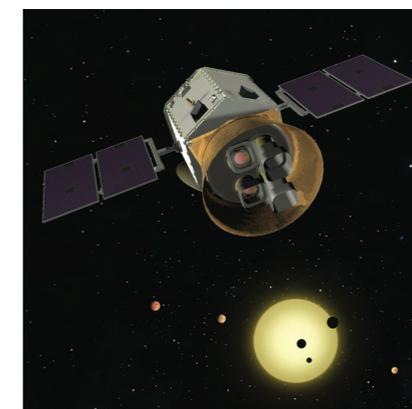
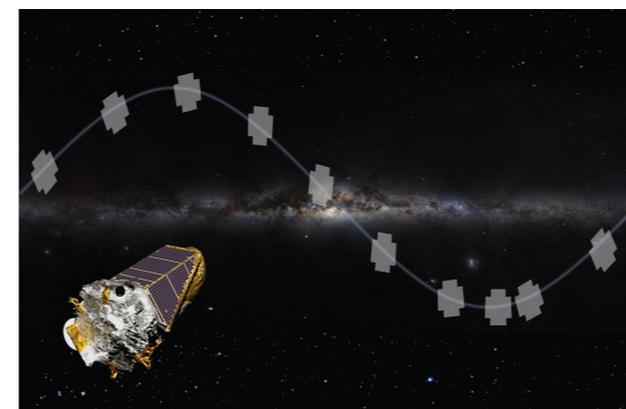


The space photometry revolution and our understanding of RR Lyrae stars

Róbert Szabó
Konkoly Observatory, MTA CSFK
Budapest, Hungary



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- **W. W. Weiss** Institute of Astronomy, University of Vienna, Vienna, Austria
- **KASC RR Lyrae + Cepheid Working Group**

• KASC RR Lyrae + Cepheid Working Group

• W. W. Weiss Institute of Astronomy, University of Vienna, Vienna, Austria



Introduction



In 2006, a colleague of mine asked:

'What are you going to do with RR Lyrae light curves observed from space?

Those stars are soooooooooooooo

'

In 2006, a colleague of mine asked:

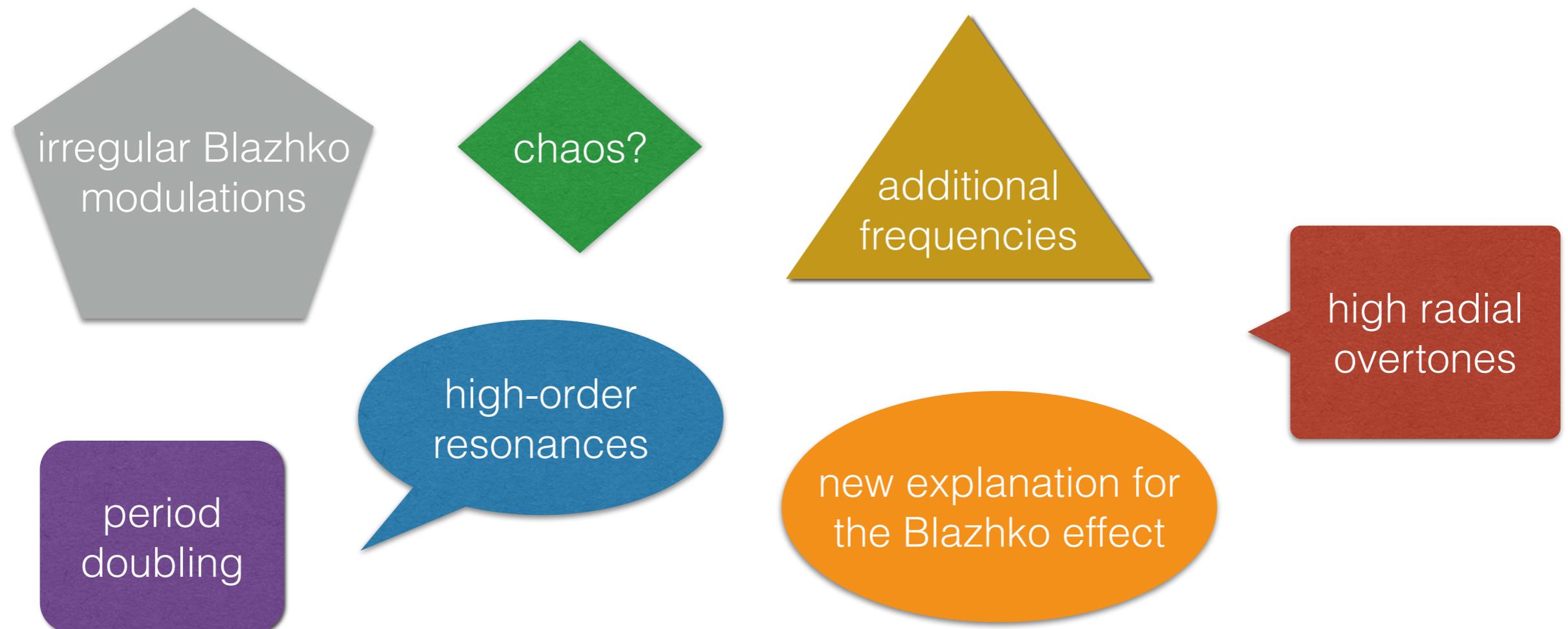
‘What are you going to do with RR Lyrae light curves observed from space?’

Those stars are sooooooooooooo **BORING** *!’*

In 2006, a colleague of mine asked:

‘What are you going to do with RR Lyrae light curves observed from space?’

Those stars are sooooooooooooo **BORING** !’



In 2006, a colleague of mine asked:

‘What are you going to do with RR Lyrae light curves observed from space?’

Those stars are sooooooooooooo **BORING** !’

irregular Blazhko modulations

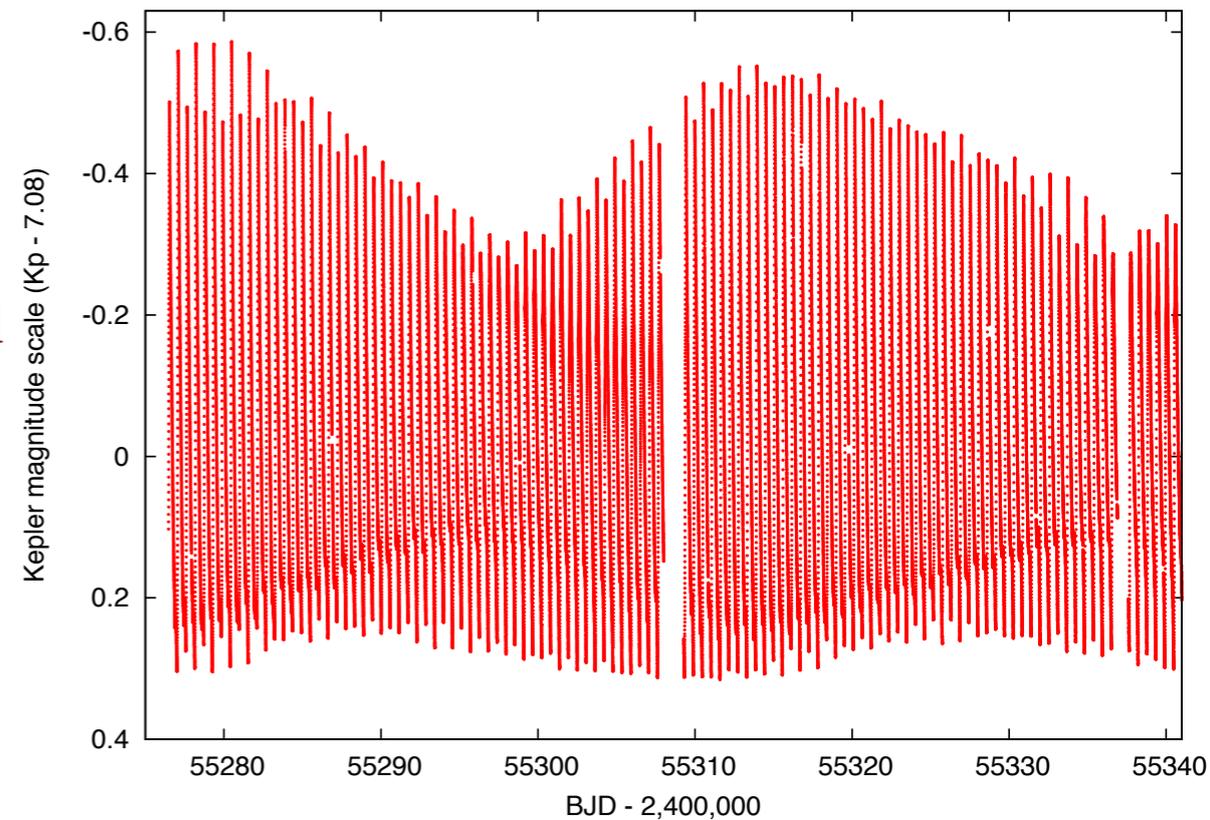
period doubling



on for
effect

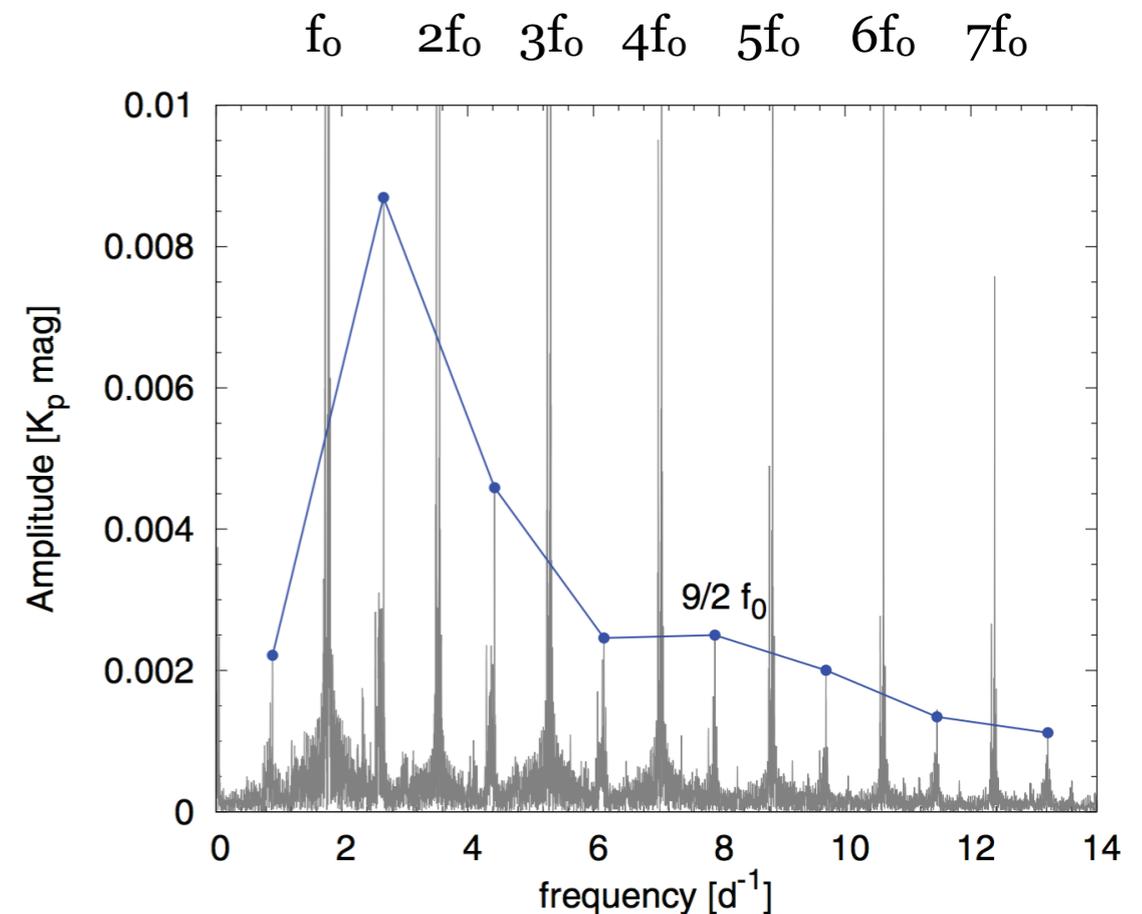
high radial overtones

- Alternating maxima/cycles
- Half-integer frequencies (HIF)
- PD bifurcations - route to chaos
- Only in Blazhko RR Lyrae, no PD in non-modulated RRab stars
- PD is caused by a 9:2 resonance between the F and the 9th radial overtone (Szabó et al. 2010, Kolláth et al. 2011)
- Resonance paradigm: the same resonance is able to cause the Blazhko effect itself (Buchler & Kolláth 2011)
- PD occurrence statistics:
Kepler: 10/16 (Benkó et al. 2014, accepted)
CoRoT: 4/6 (Szabó et al. 2014, submitted)



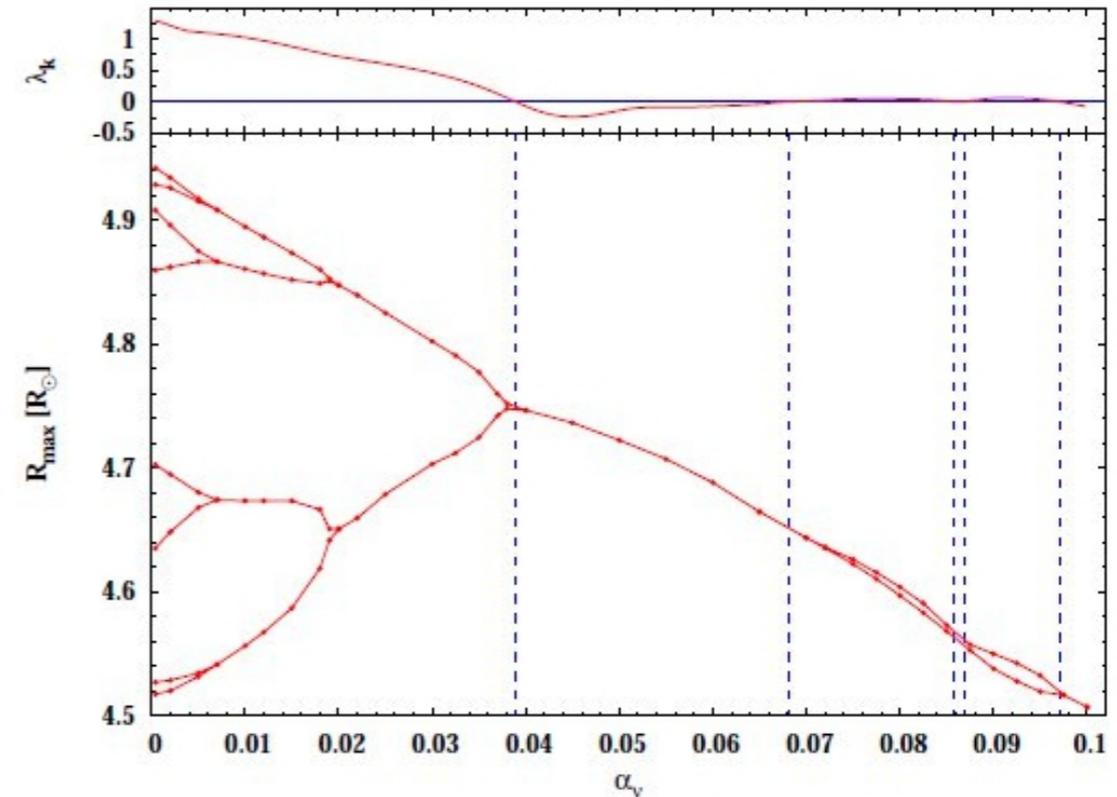
Period doubling
 RR Lyrae, the eponym
 Kepler, Q5 SC

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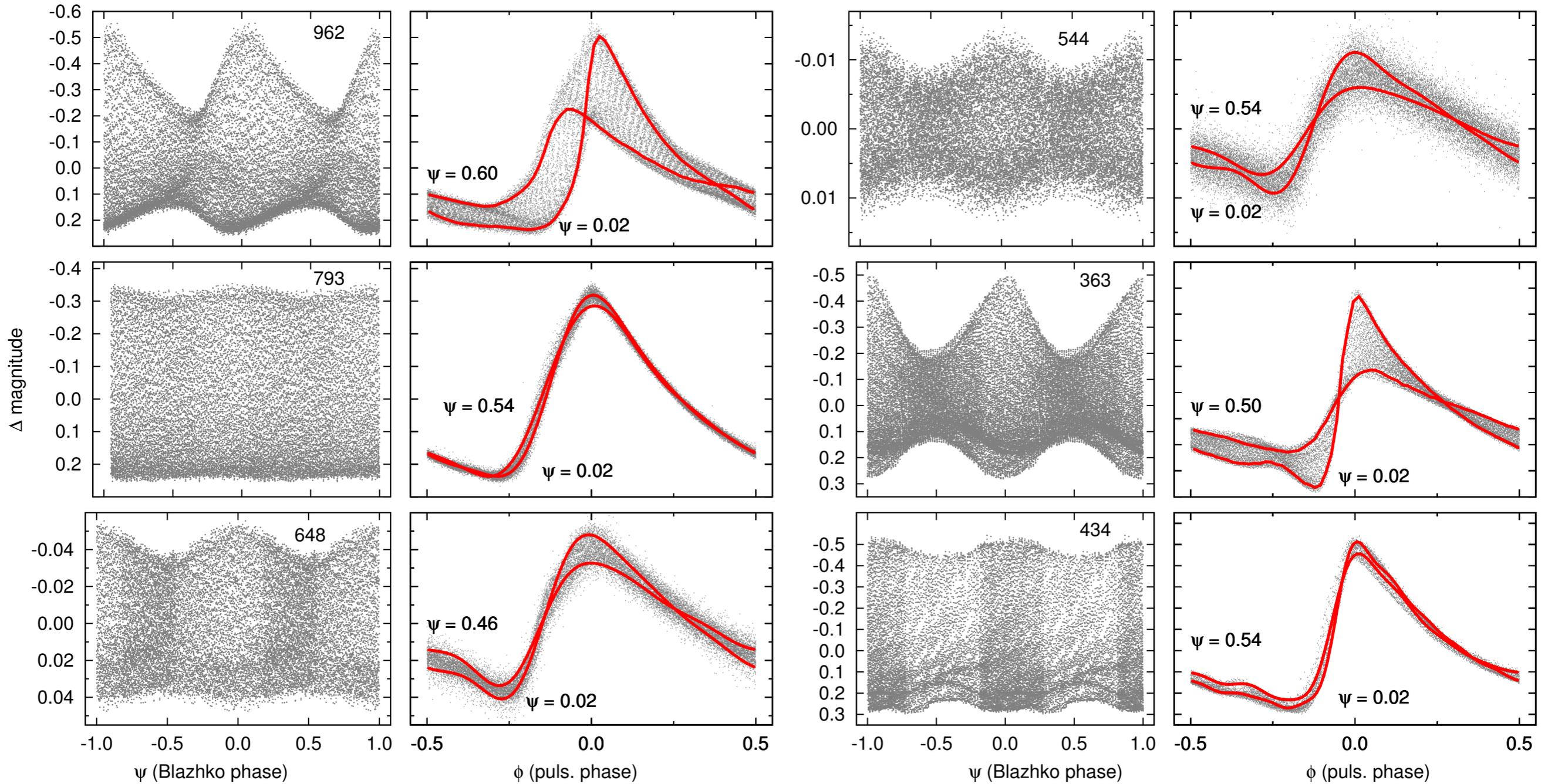


Half-integer frequencies
 RR Lyrae, the prototype
 Molnár et al. 2012

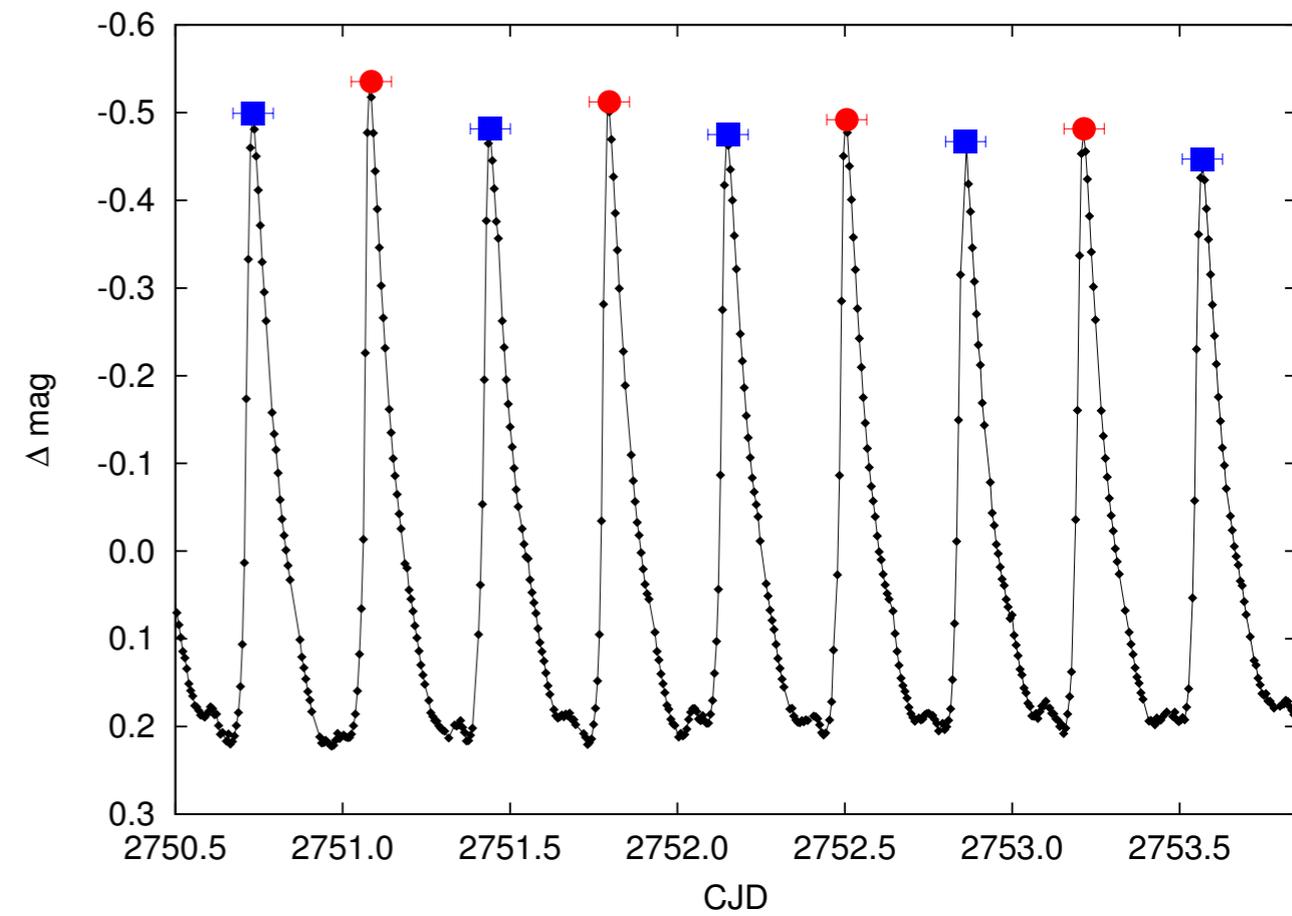
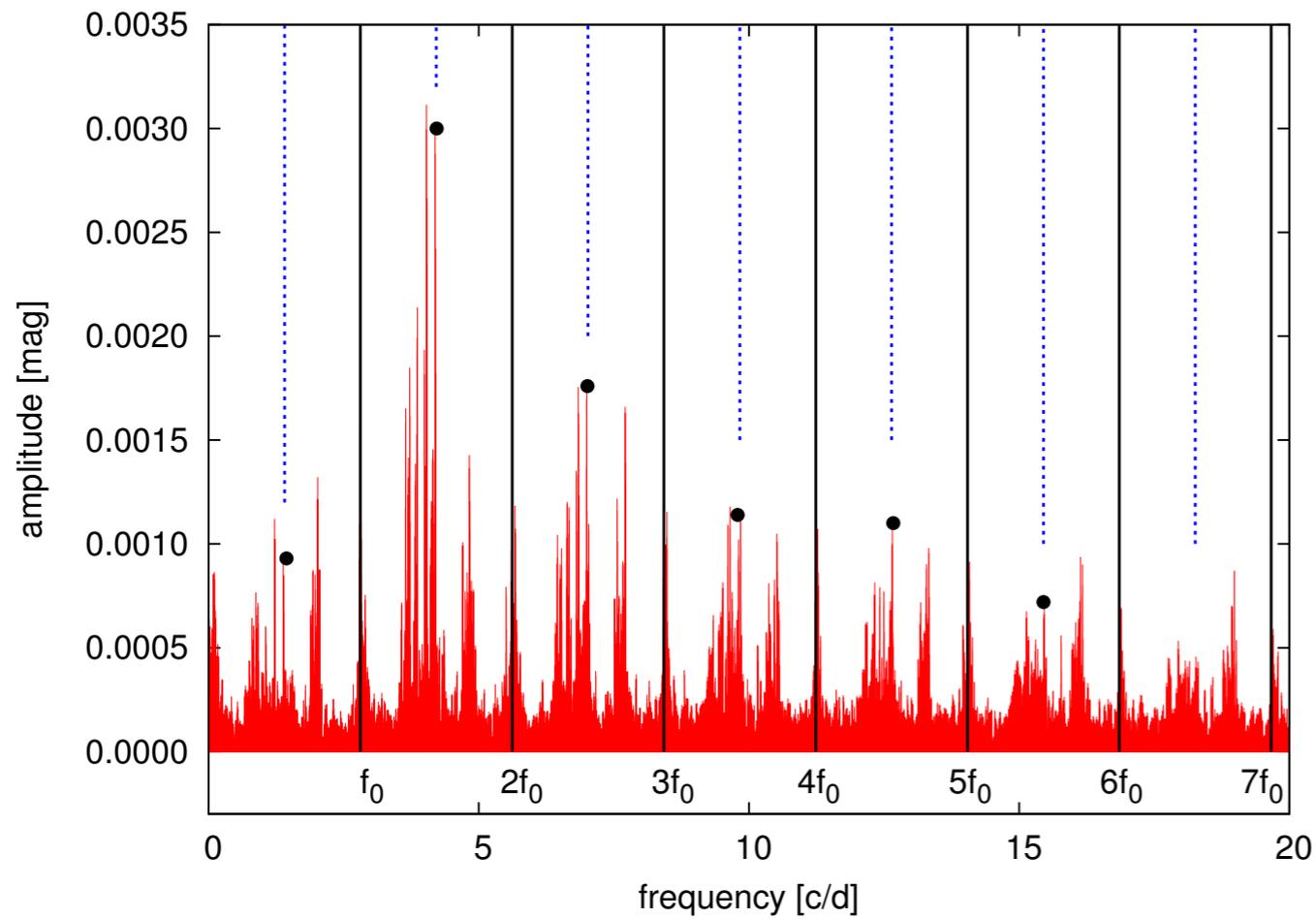
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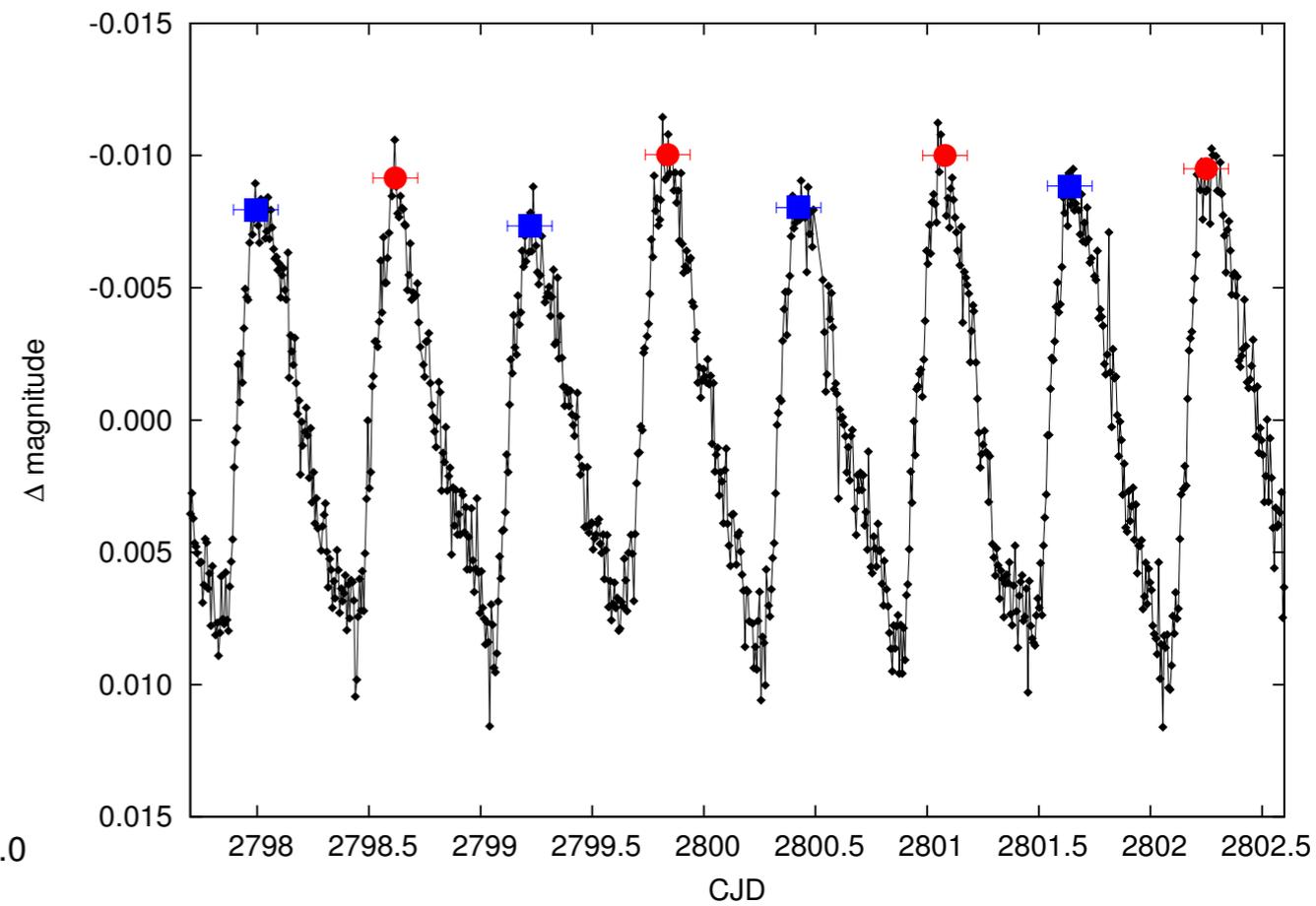
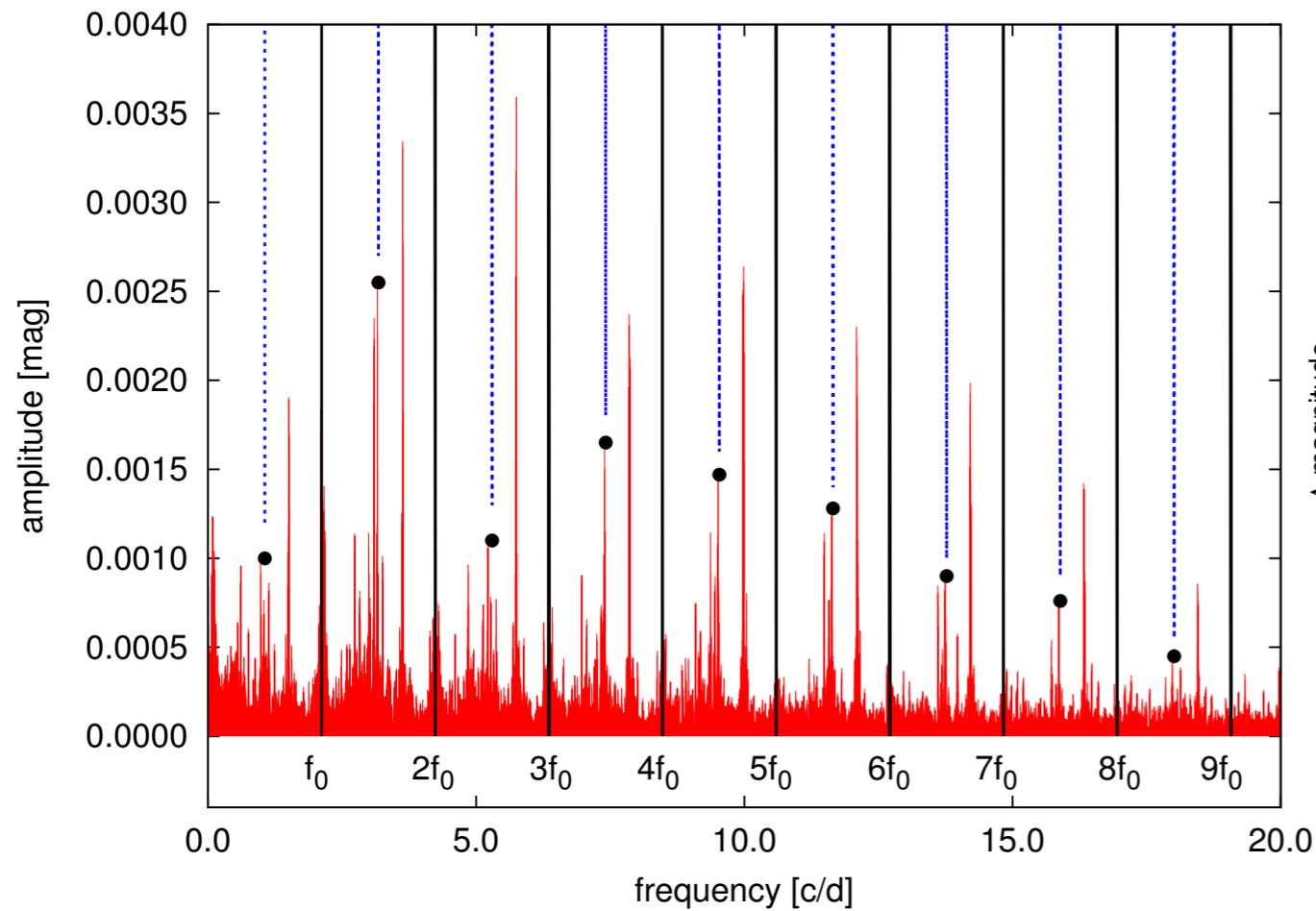
PD bifurcation in
RR Lyrae hydro models
Kolláth et al. 2011



Period doubling



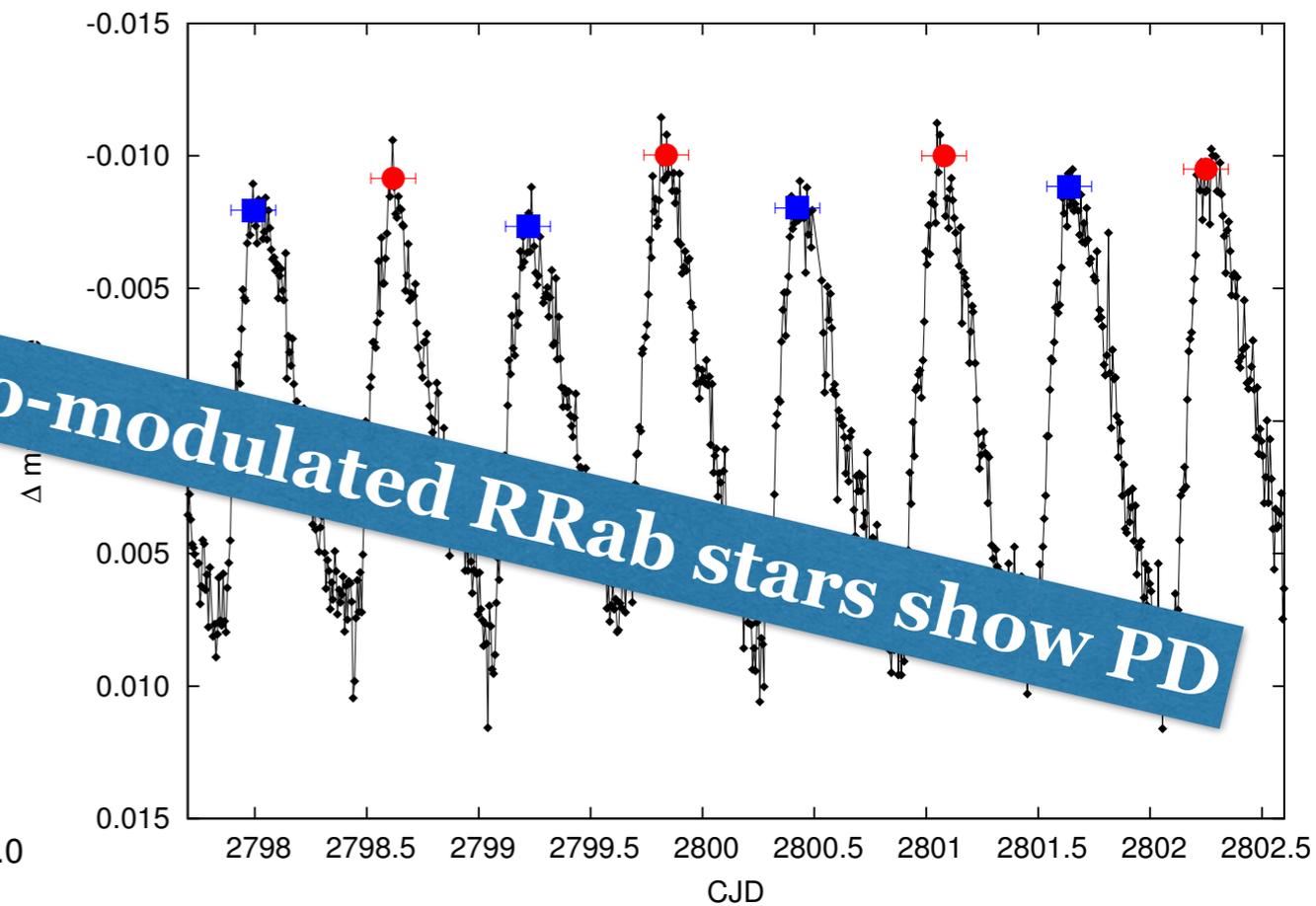
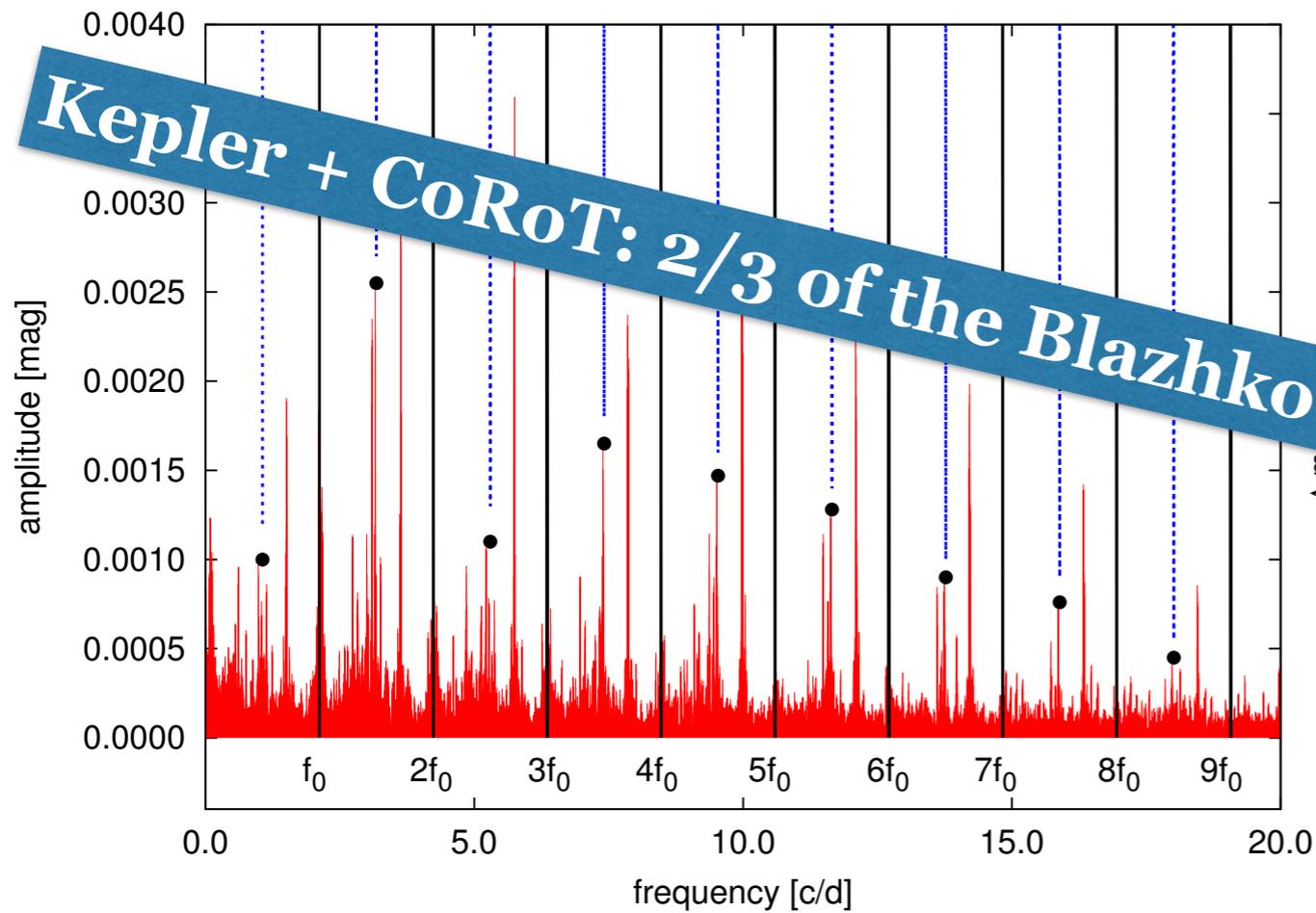
CoRoT 0100689962 - V1127 Aql
 Szabó et al. 2014, submitted



CoRoT 01011289793

CoRoT 0101503544

Szabó et al. 2014, submitted

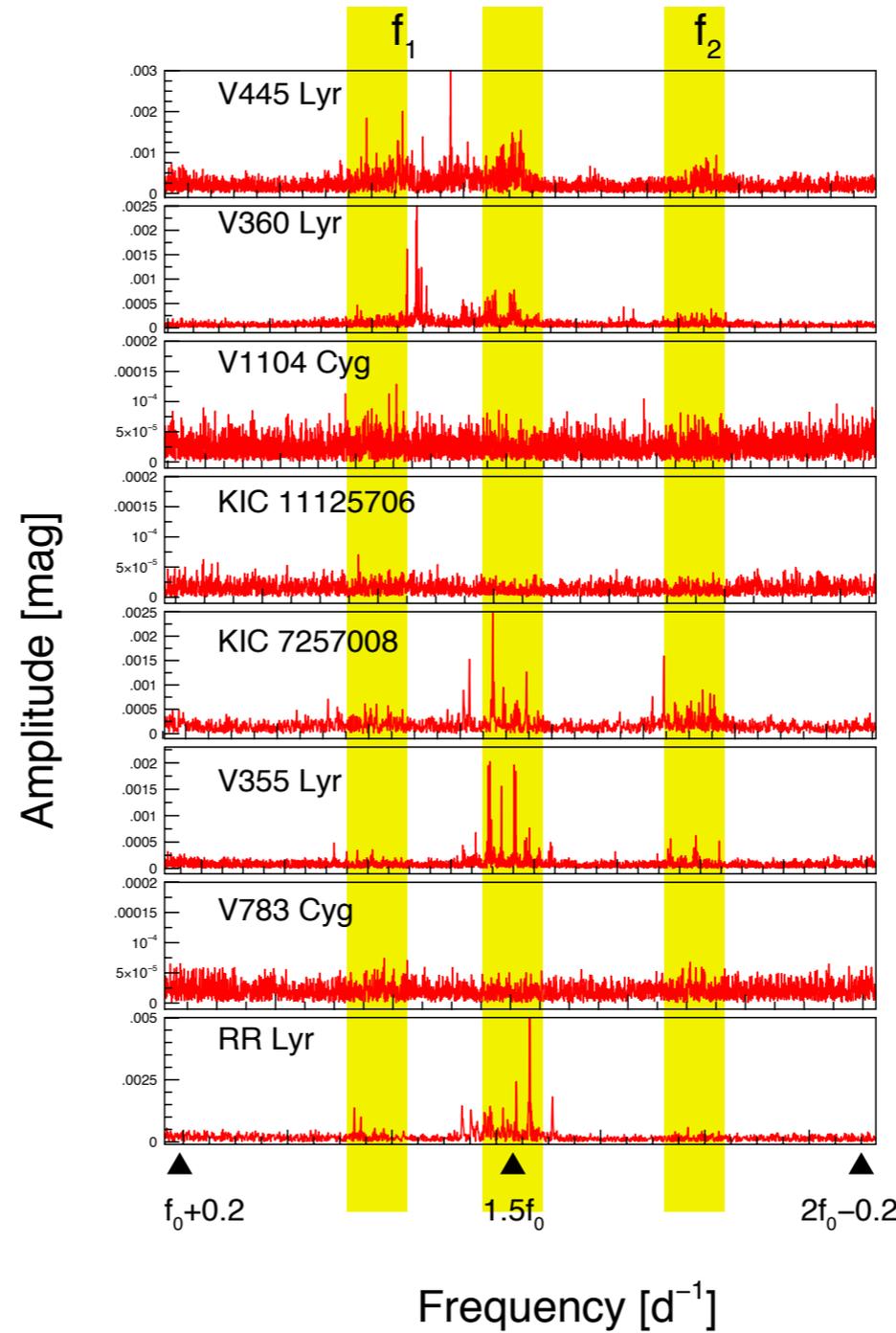
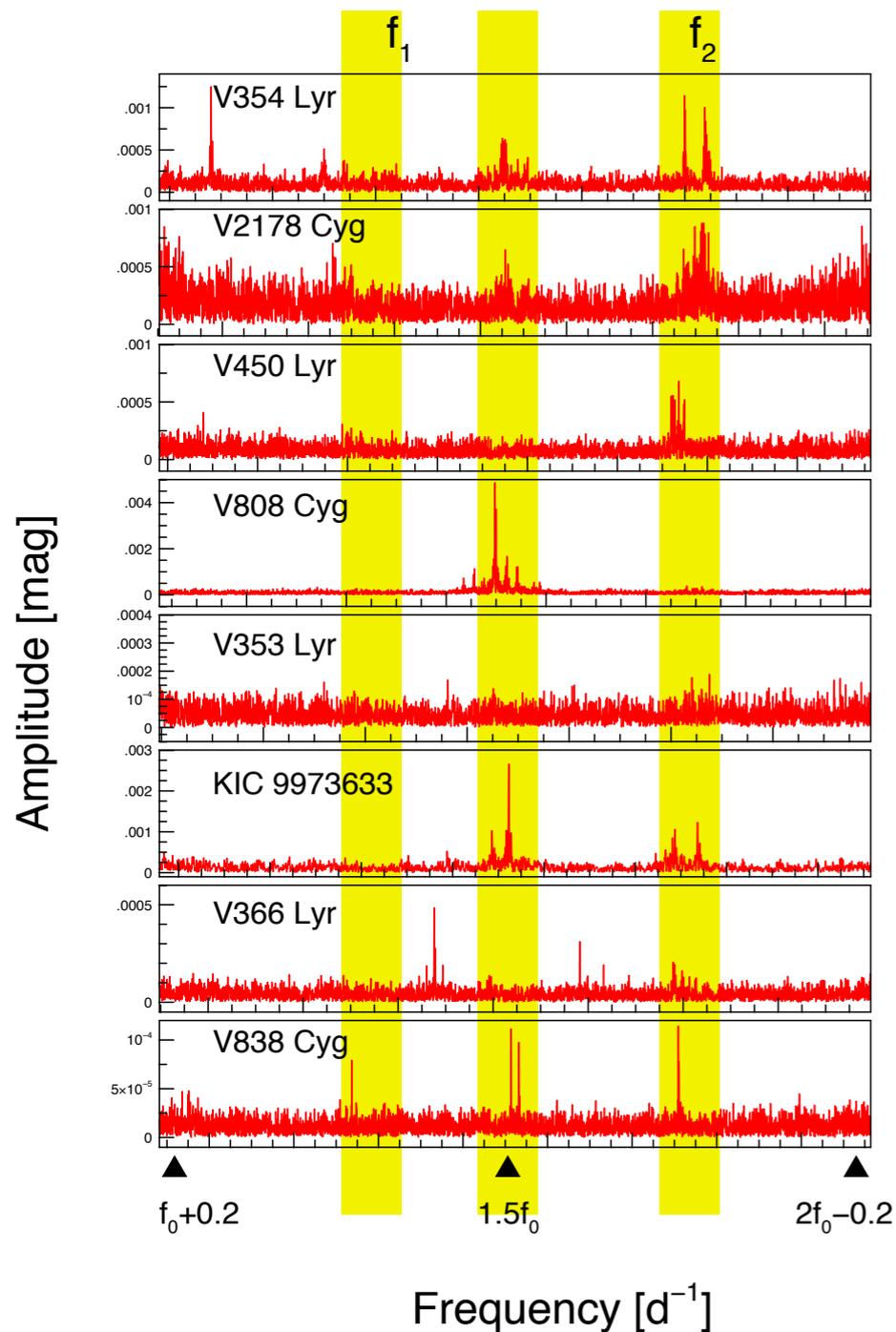


CoRoT 01011289793

CoRoT 0101503544

Szabó et al. 2014, submitted

Radial modes



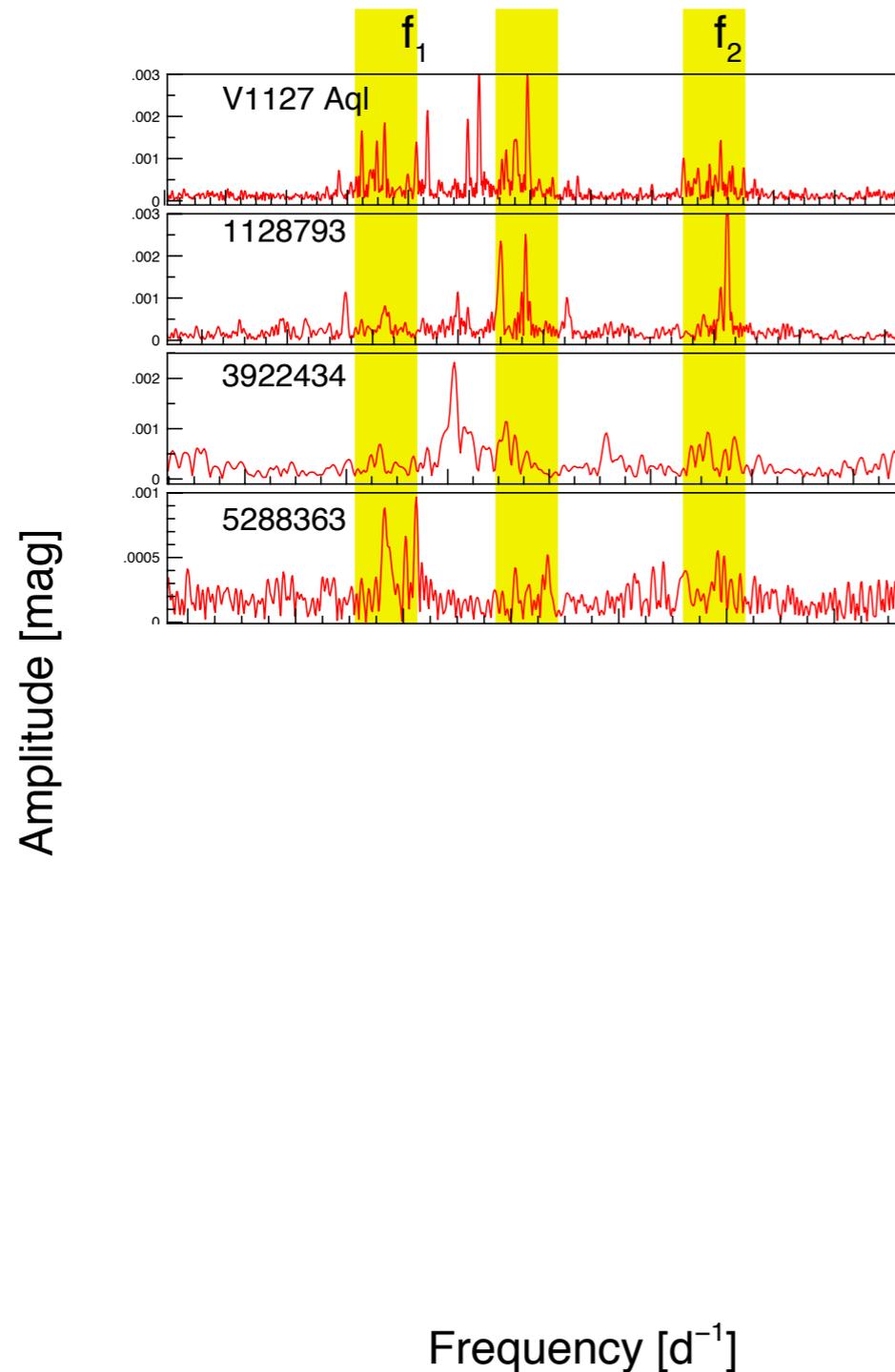
- Frequencies around the radial O1, O2 are common

- **Radial modes**
RR Lyrae: O1
Molnár et al. 2012

or

- **Nonradial modes**
in 1:1 resonance with the radial mode
Dziembowski & Mizerski 2004
van Hoolst et al. 1998

Radial modes

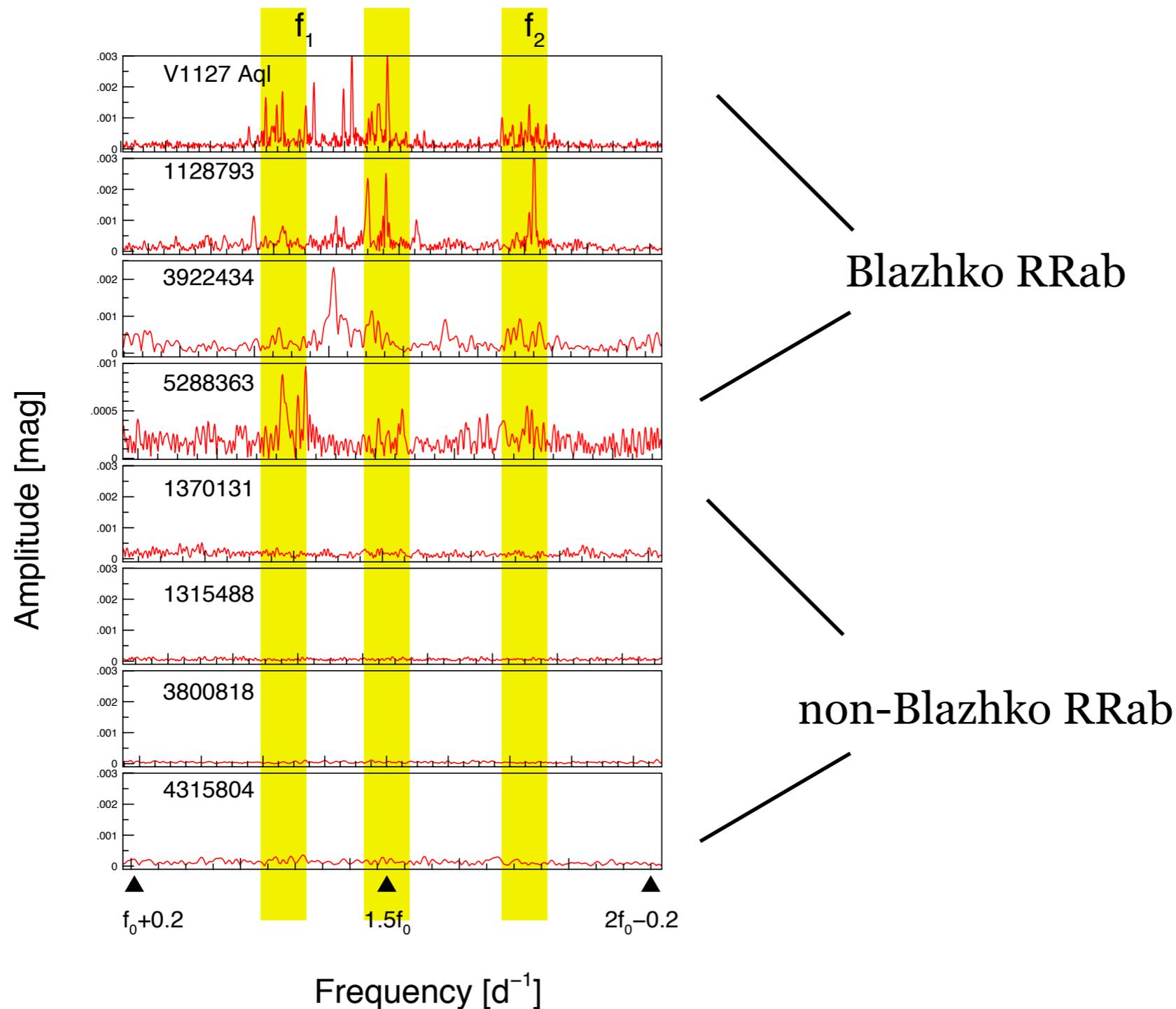


Blazhko RRab

Modulated stars:
frequencies around the
radial O1, O2 are
common.

**Radial modes or
Nonradial modes in
1:1 resonance with the
radial mode**

Radial modes

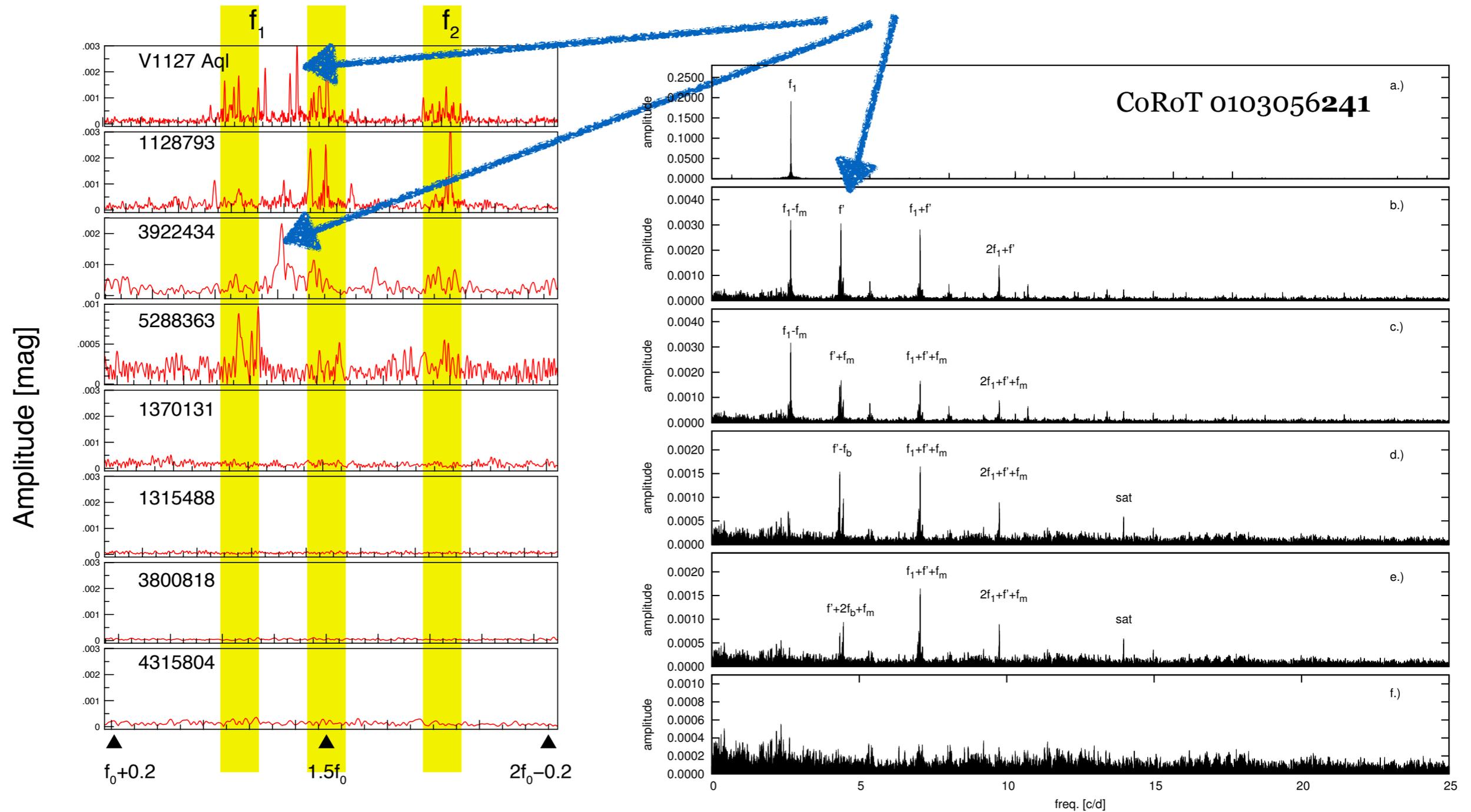


Modulated stars:
frequencies around the radial O1, O2 are common.

Radial modes or Nonradial modes in 1:1 resonance with the radial mode

Non-modulated stars:
no additional frequencies in the spectra down to the Kepler & CoRoT limits.
Nemec et al. 2011 - Kepler
Szabó et al. 2014 - CoRoT

Nonradial modes

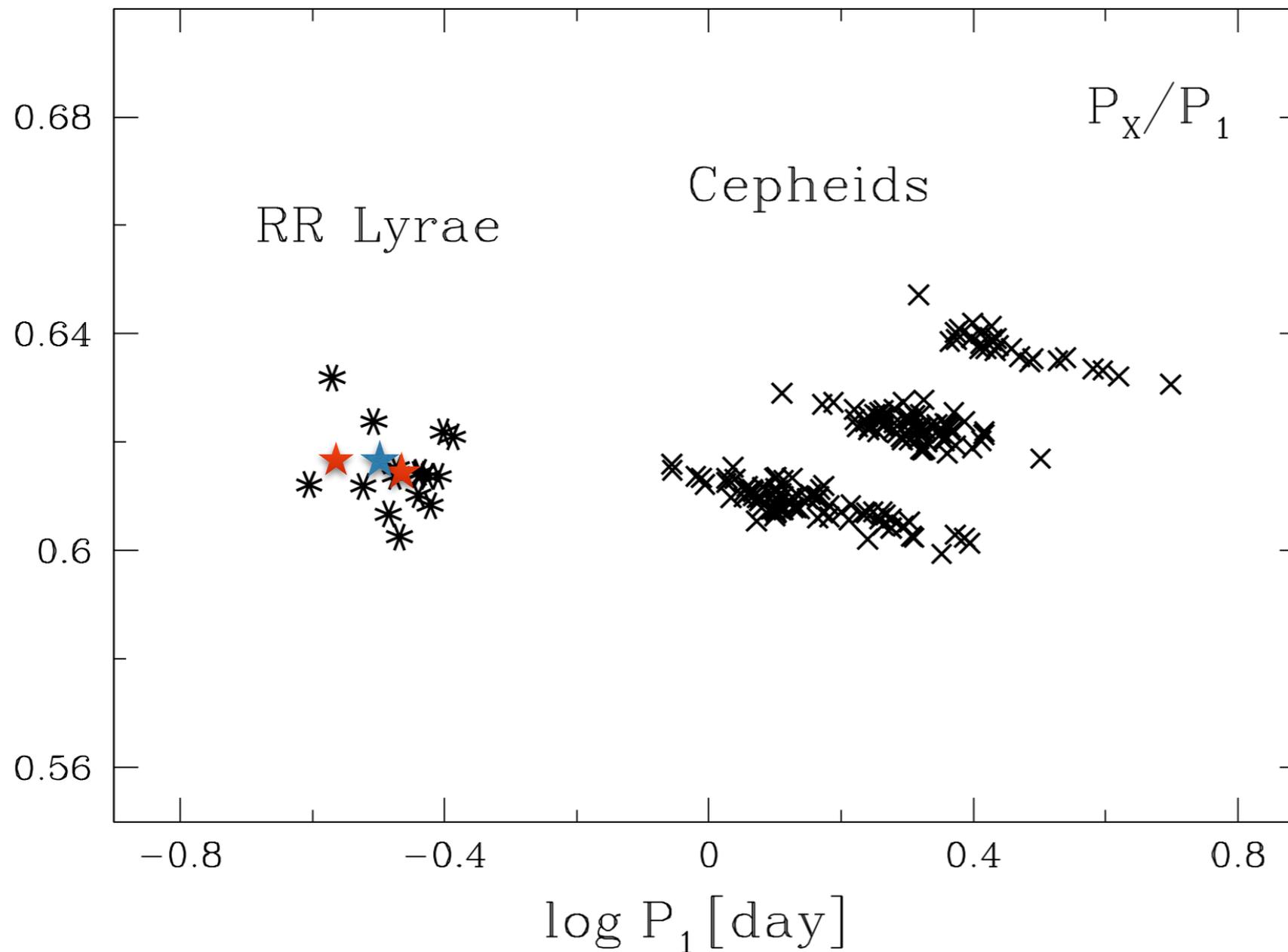


Frequency [d^{-1}]

CoRoT sample Szabó et al. 2014, A&A submitted

Non-radial modes

Moskalik, 2014



- RRc and RRd stars: period ratio of $P_x/P_1 \sim 0.61$ is frequent
- 15 RRc (4 Kepler) + 1 RRd
- CoRoT:
 - 0105036**241** RRc
 - 0105735**652** RRc
 - 0101368**812** RRd (Szabó et al. 2014)

Additional frequencies in space photometry targets

- Blazhko-modulated RRab stars : all
- non-modulated RRab stars: none
- RRC (first overtone pulsators O1): all
- RRd (fundamental mode + O1): all



Additional frequencies in space photometry targets

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- non-modulated RRab stars.



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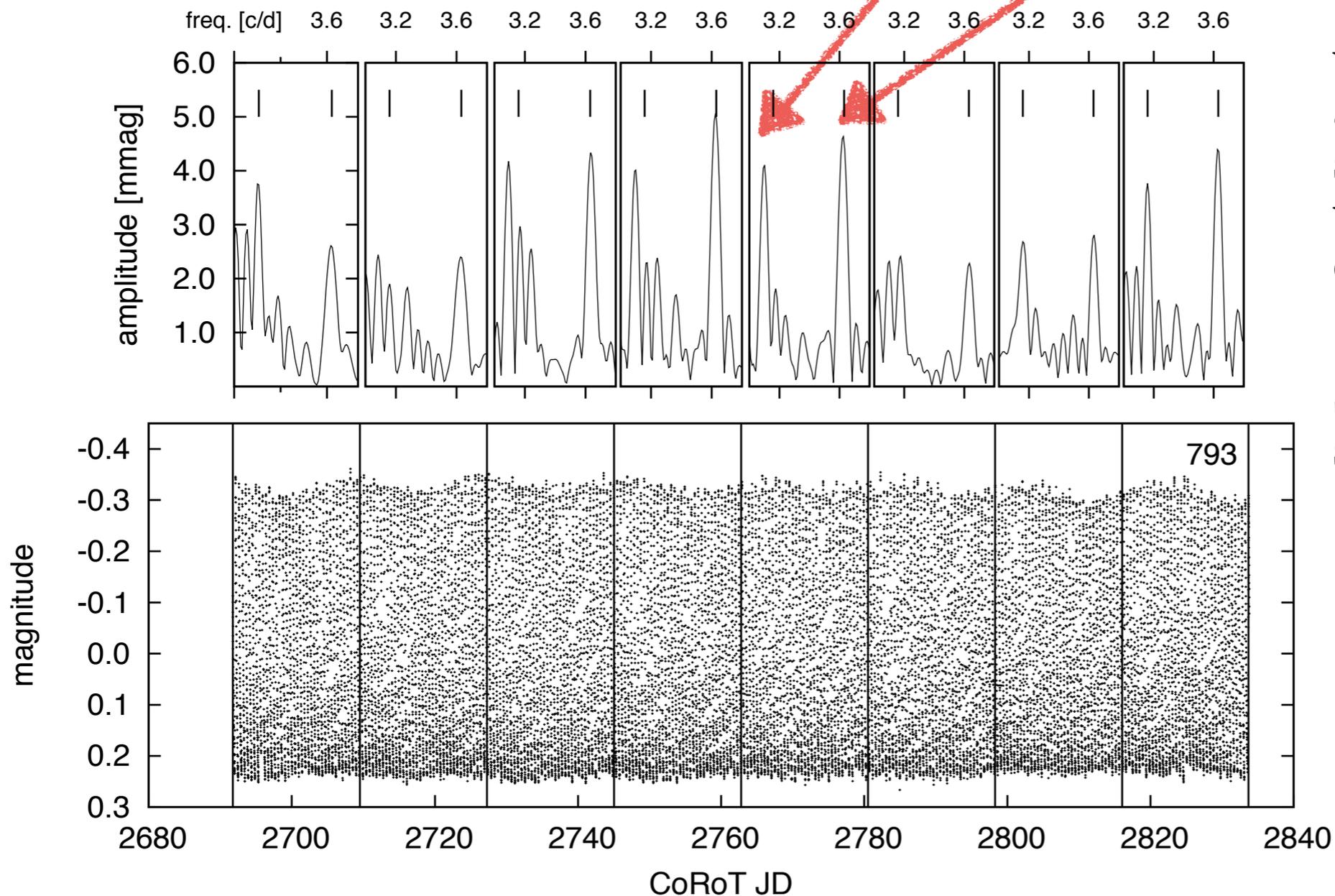


- RRd (fundamental mode + O1): all



Kepler + CoRoT: additional modes are universal!

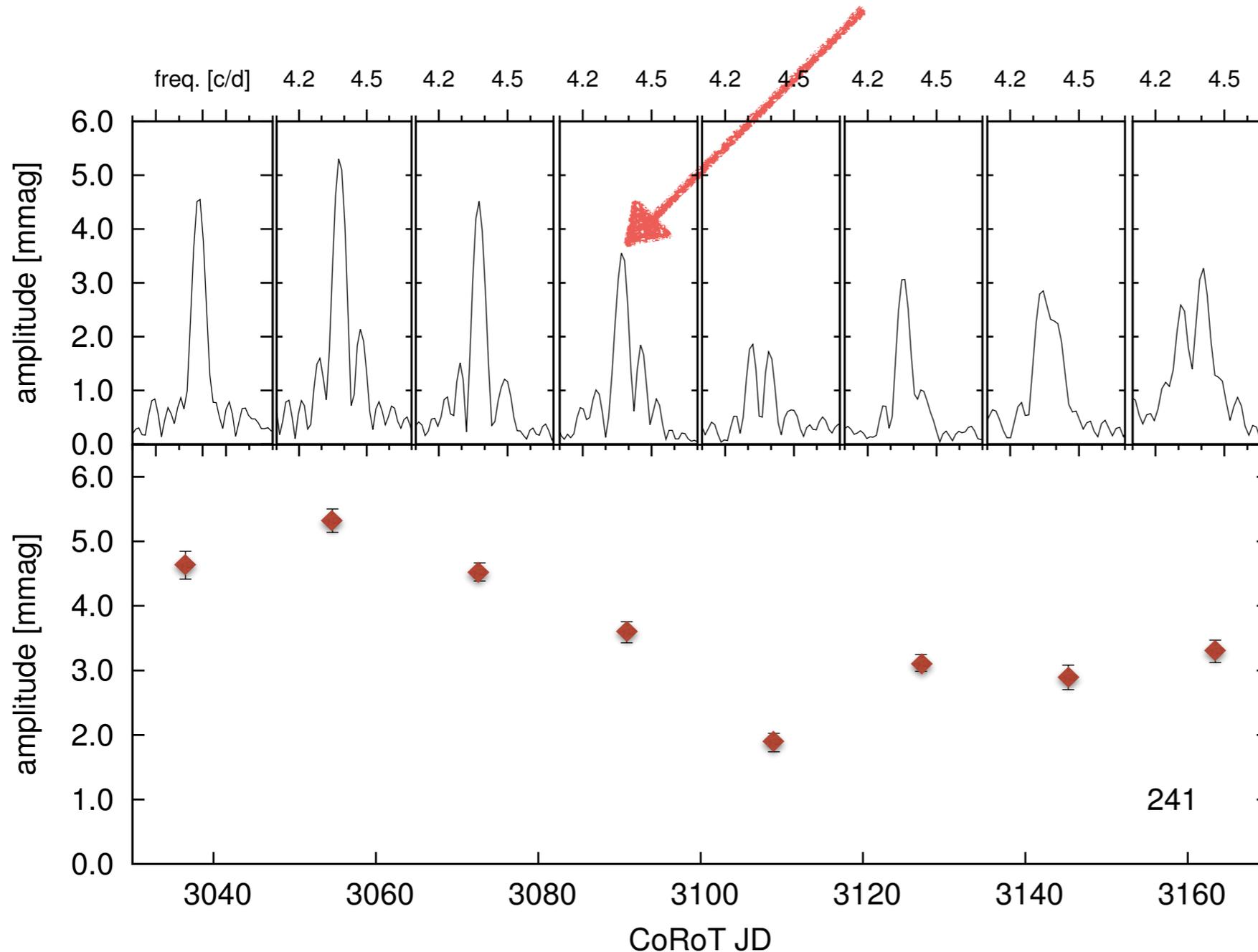
CoRoT 0101128793 Blazhko R Rab $3/2 f_0$ O2



Temporal and structural variability of the additional frequencies seems to be **ubiquitous** whether be HIFs, O2, f_x or other nonradial

Szabó et al. 2014 submitted

CoRoT 0105036241 RRc f_x



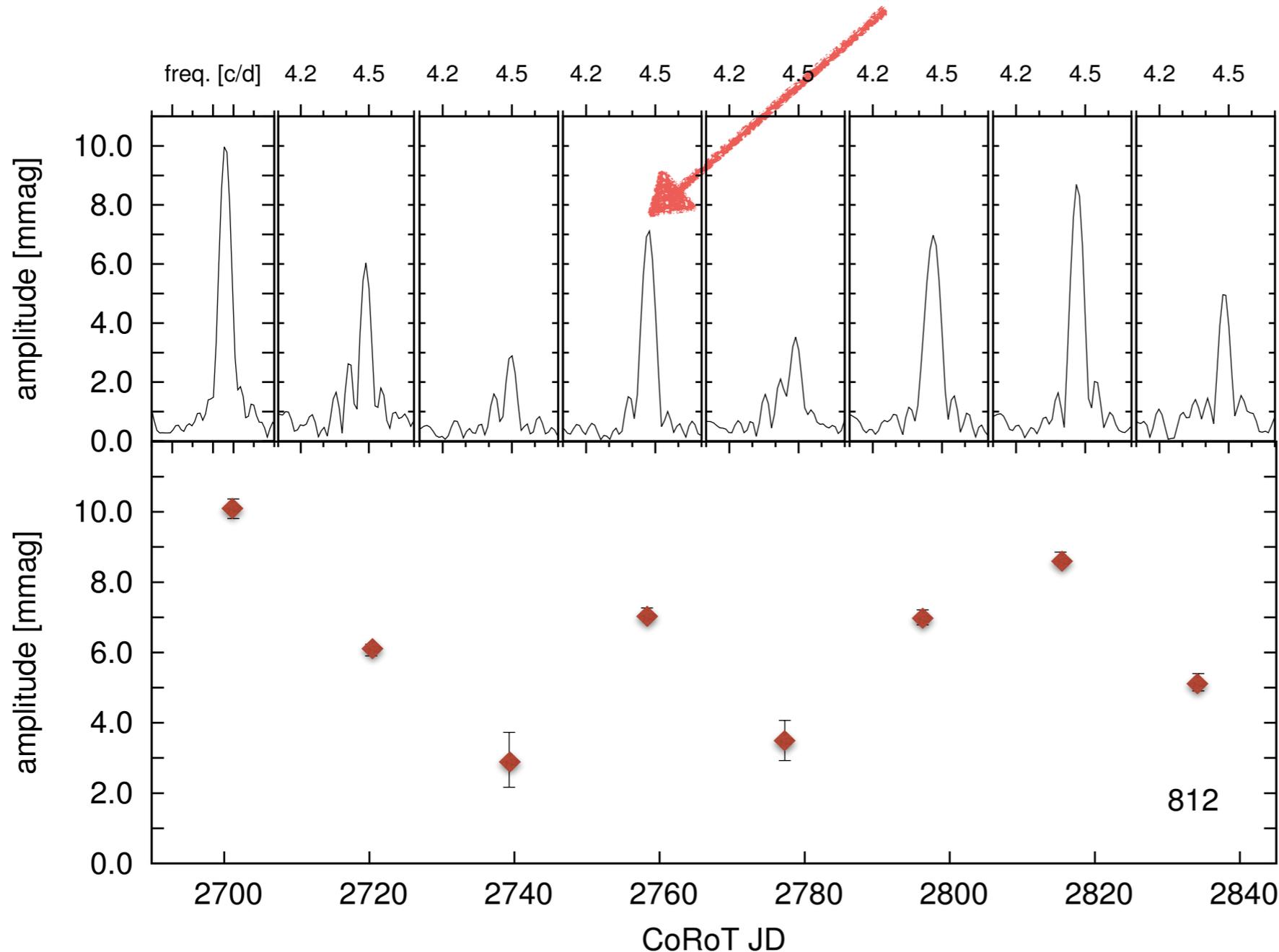
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Szabó et al. 2014 submitted

amplitude change or close-by unresolved frequencies?

CoRoT 0101368812 RRd

f_x



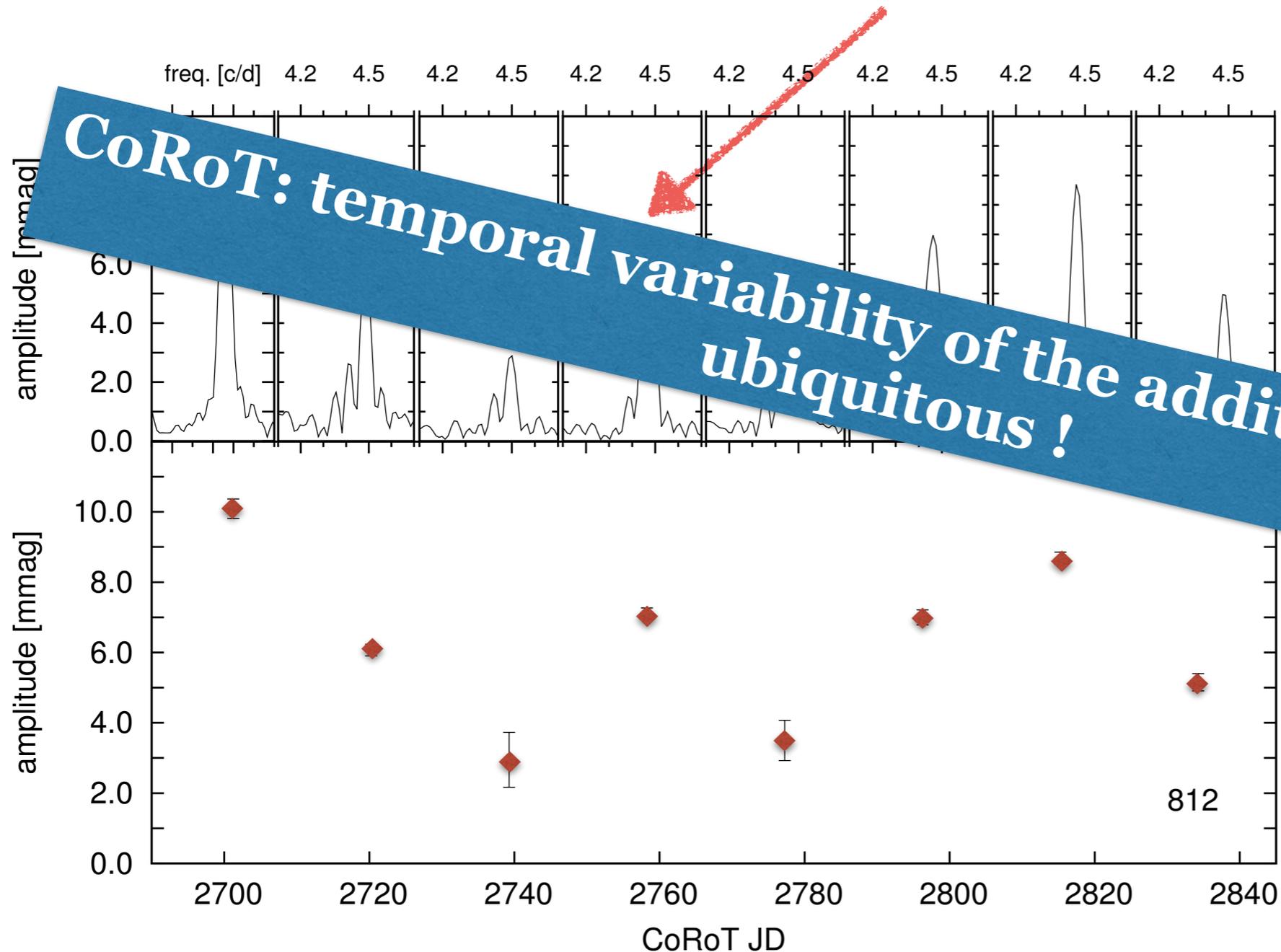
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Szabó et al. 2014 submitted

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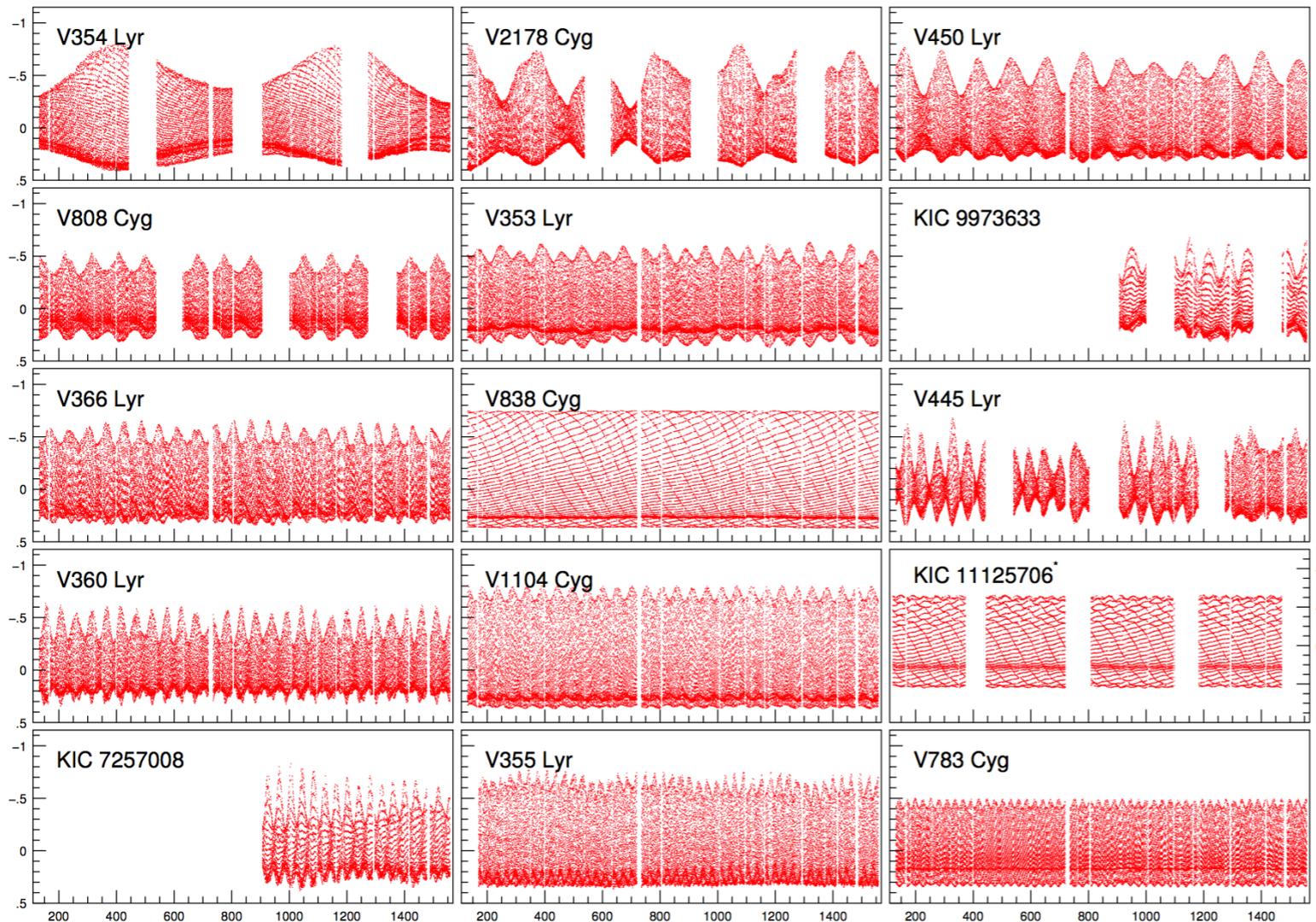
CoRoT 0101368812 RRd

f_x



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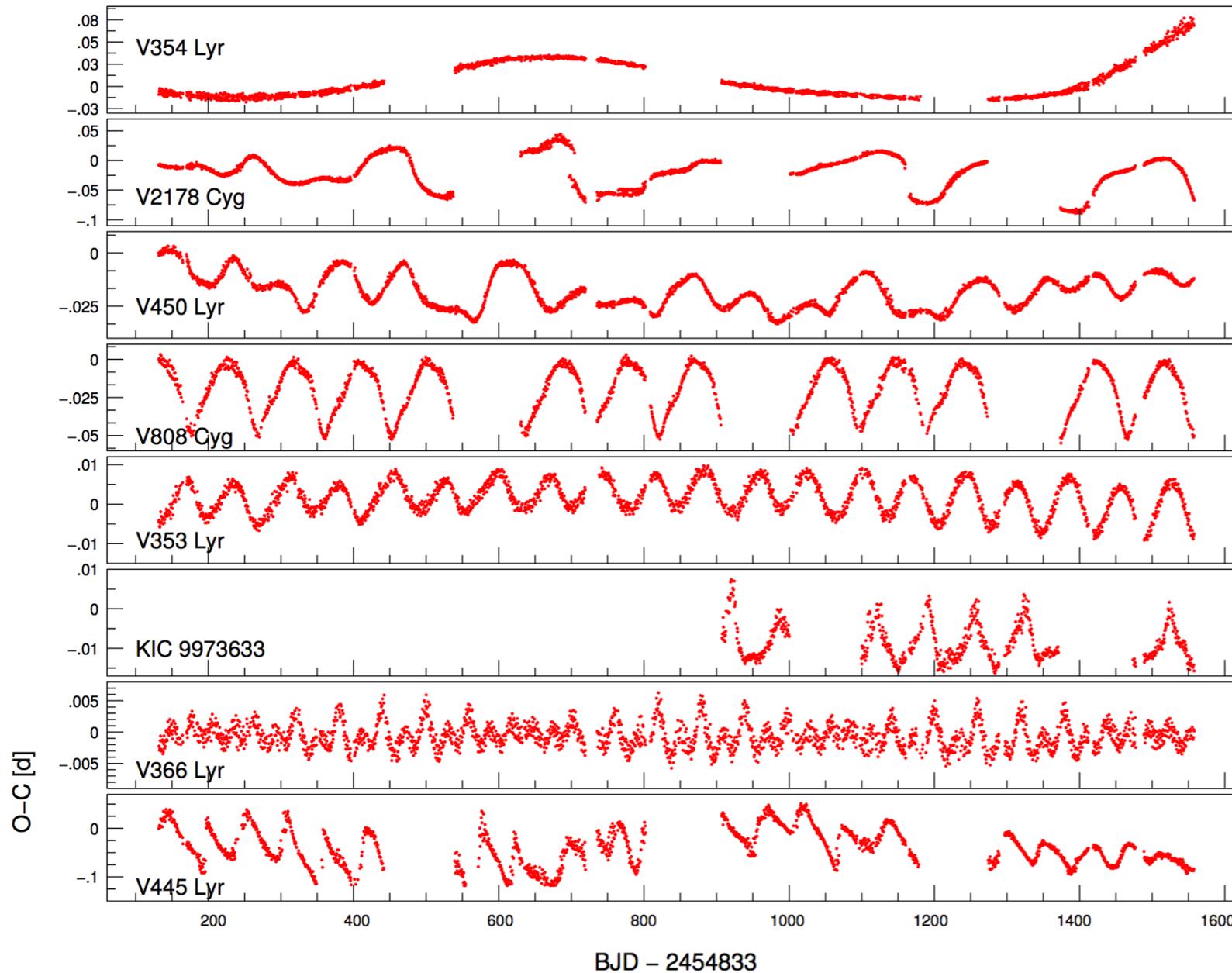
BJD – 2454833

Kepler Blazhko sample Benkő et al. 2014

- Is the Blazhko modulation simple and monopерiodic? in general: **NO**
- **2 modulation periods**

XZ Cyg	LaCluyzé et al.	2004
UZ UMa	Sódor et al.	2006
SU Col	Szczygiel & Fabrycky	2007
LS Her	Wils et al.	2008
CZ Lac	Sódor et al.	2011
- Large sample
 Multiperiodic and irregularly modulated: **12%**
 Skarka 2014, A&A 562, A90
- Kepler sample
 Multiple modulations: **80%**
 Benkő et al. 2014, ApJS, accepted, arXiv: 1406.5864

Final, rectified, normalized, stitched, 4-year-long Kepler Blazhko light curves with tailor-made apertures are available: <http://www.konkoly.hu/KIK/data.html>



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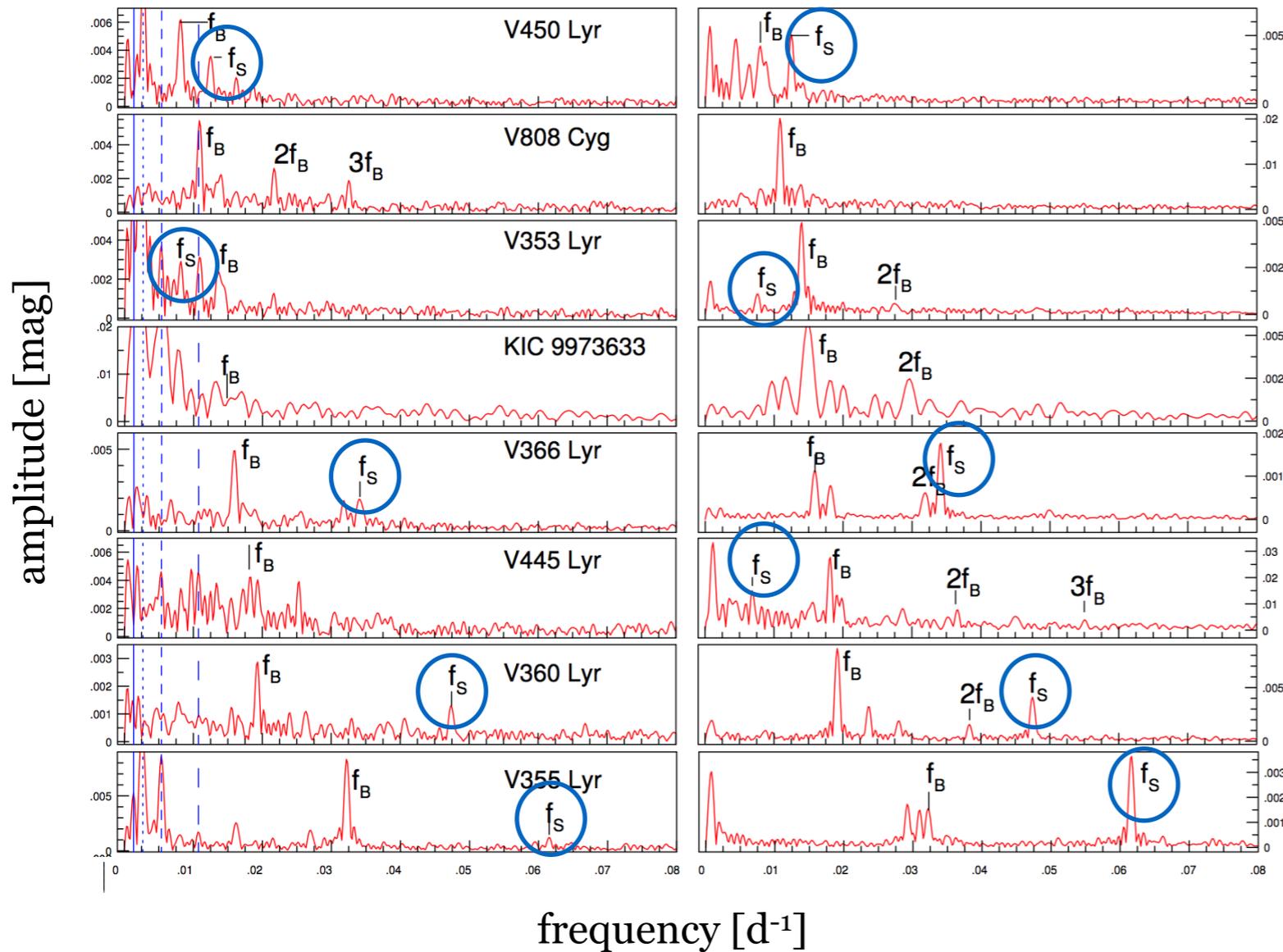
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Light curve

O-C diagram



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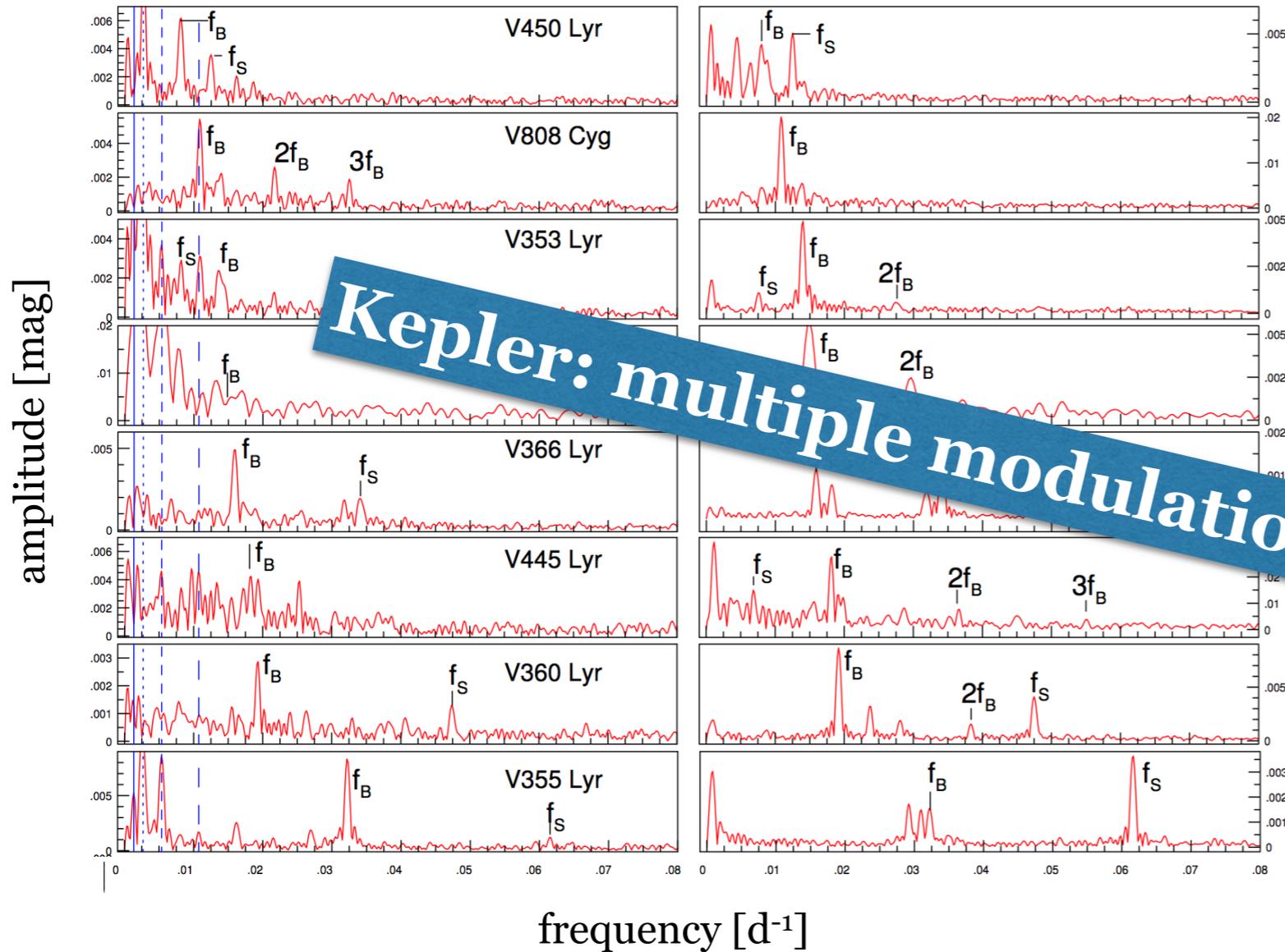
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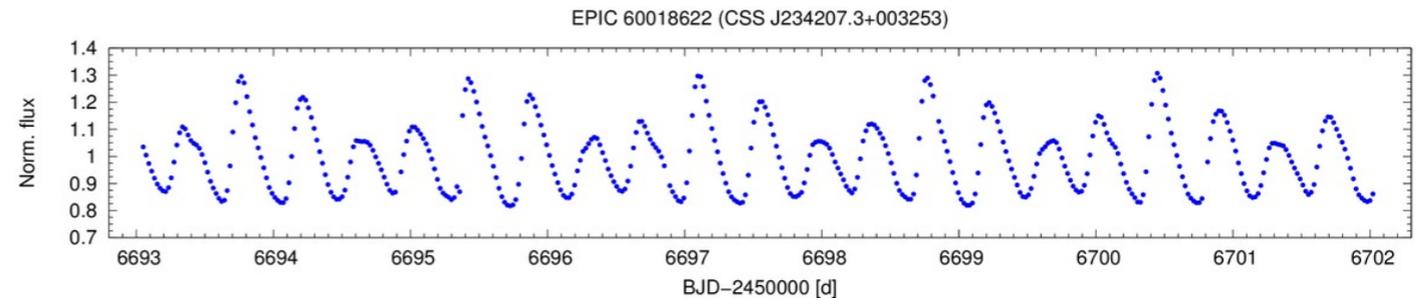
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K2, TESS, PLATO

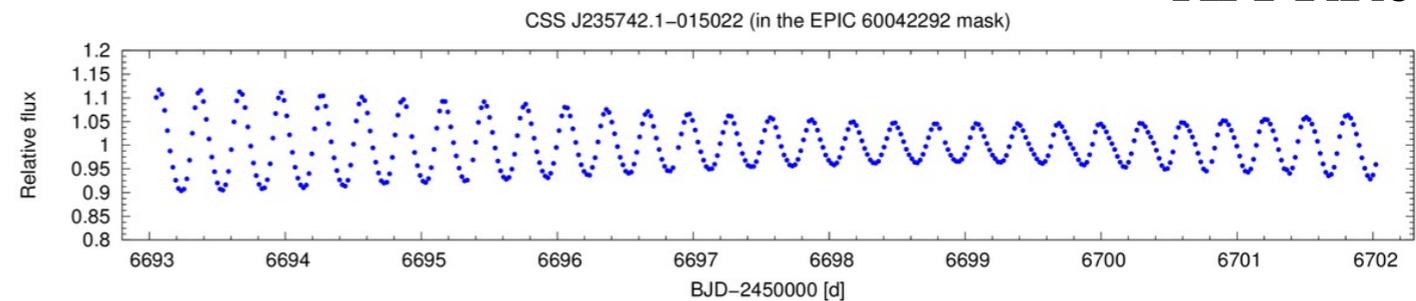
- will see different populations, metallicities
- rare objects (RRd stars, ultra-long period Blazhko modulations, etc.)
- will deliver much better statistics to understand similarities and differences in dynamical phenomena, period doubling, nonradial modes, resonances, ... and occurrence rates

K2 preliminary results

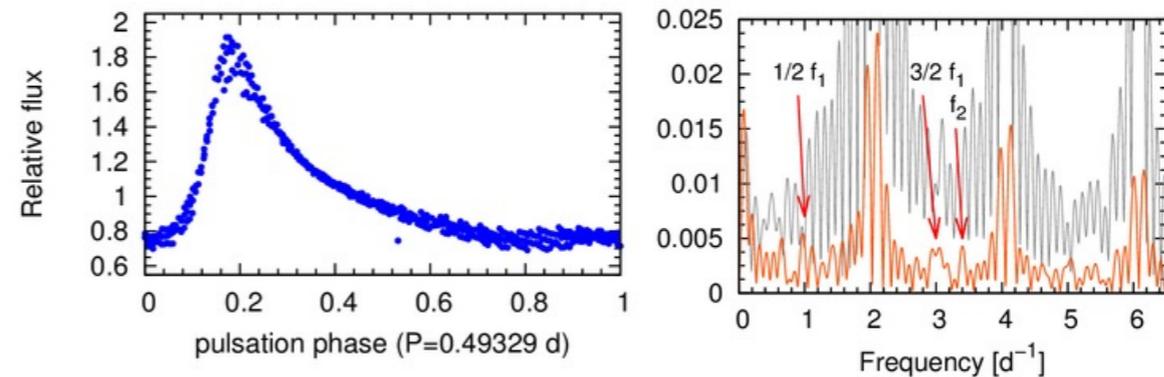
RRd



AM RRc



EPIC 60018743



Blazhko
RRab
with PD

Molnár et al. 2014, IBVS, 6108
See L. Molnár's poster

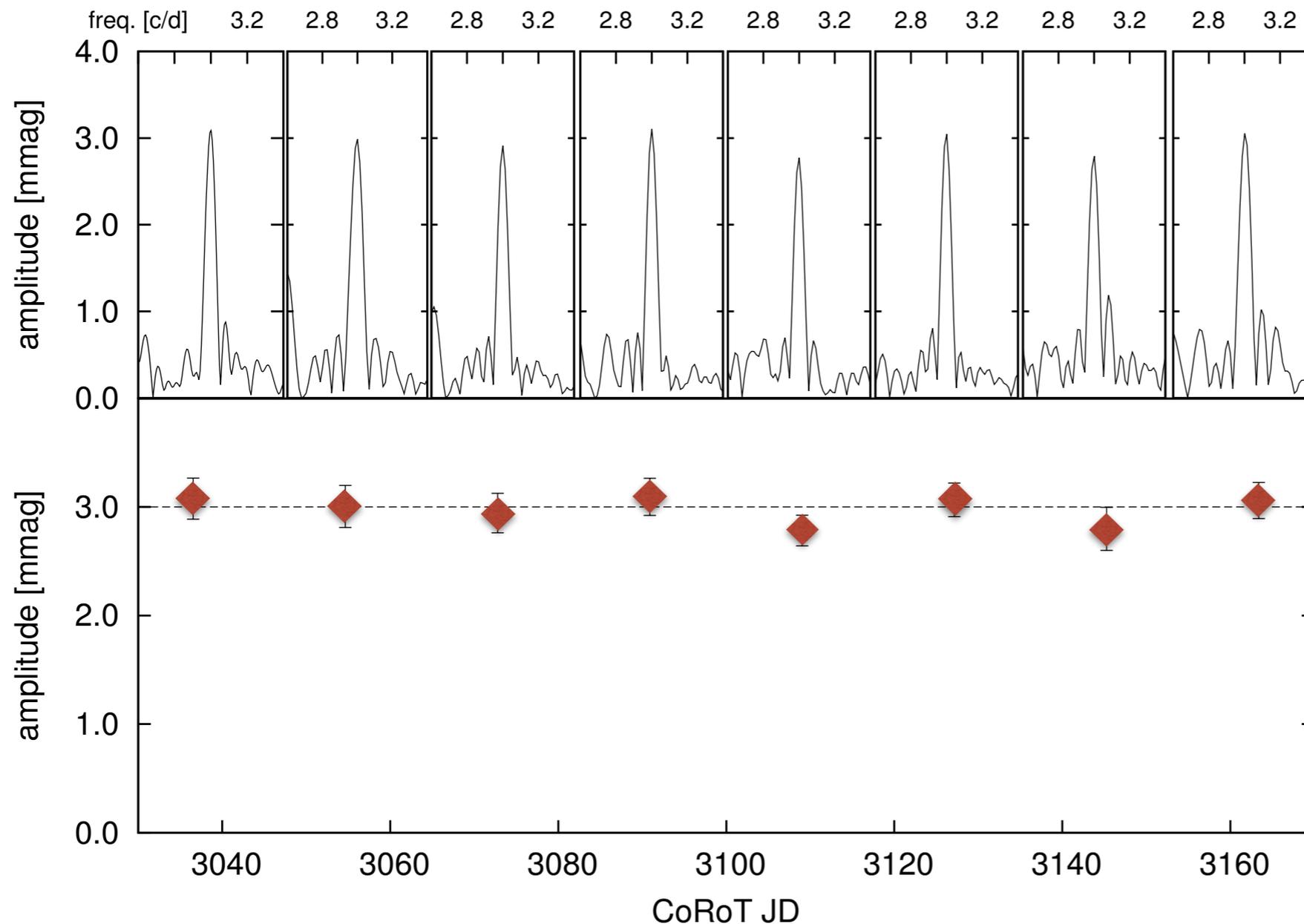
- RR Lyrae stars are interesting!
- **Period doubling** is seen in many Blazhko RRab stars
- Additional modes are **ubiquitous**
 - in **Blazhko RRab, RRc** and **RRd** stars
 - missing from **non-modulated RRab stars**
- Additional frequencies show **temporal variability**
- **Multiple Blazhko-modulations** are very common
- K2, TESS, PLATO: **bright future** for space photometry



Thank you

- ESA PECS project No. 4000103541/11/NL/KML
- Hungarian OTKA grant K83790
- European Community's Seventh Framework Programme (FP7/2007–2013) grants
 - no. 312844 (SPACEINN)
 - no. 269194 (IRSES/ASK)
 - ERC grant agreement no. 338251 (StellarAges)
- János Bolyai Research Scholarship of the Hungarian Acad. of Sciences

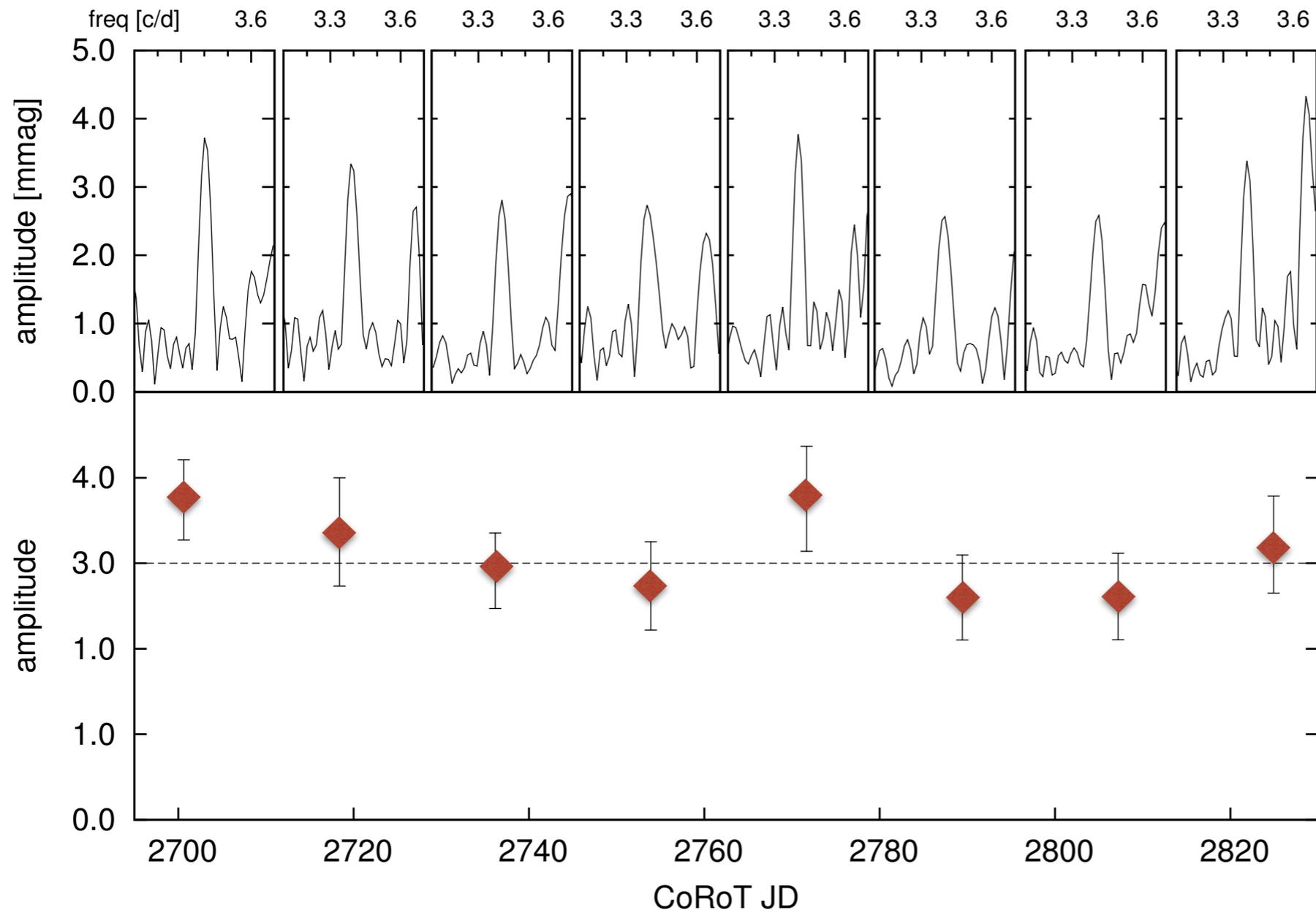
test cases - clean spectrum



Temporal and structural variability of the additional frequencies seems to be **ubiquitous** whether be HIFs, O_2 , f_x or other nonradial

Szabó et al. 2014 submitted

test cases - cluttered spectrum



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Szabó et al. 2014 submitted