



GD1212

~~KIC 11911480:~~

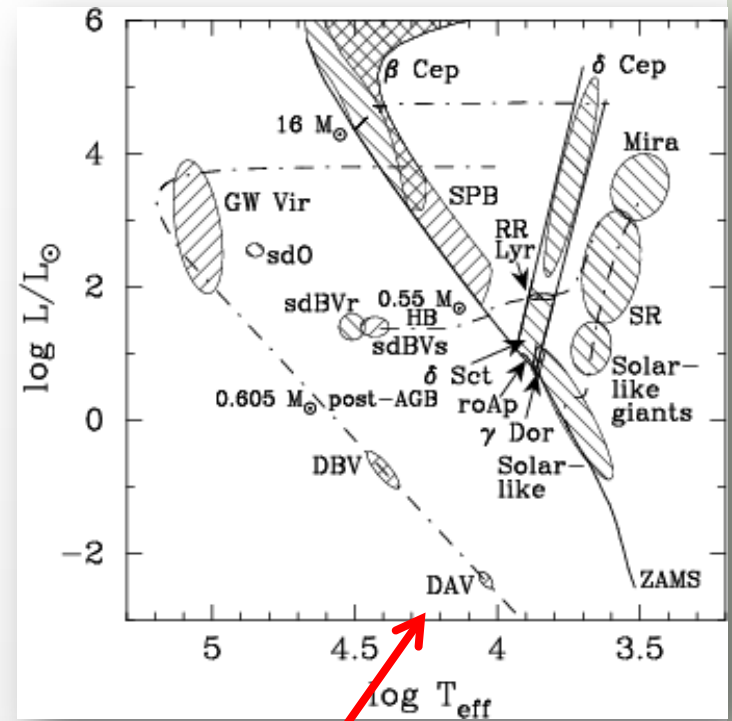
**PROBING DEEP INTO THE INTERIOR  
OF A PULSATING WHITE DWARF  
STAR**

Giammichele N., Fontaine G., Brassard P.  
*Université de Montréal*

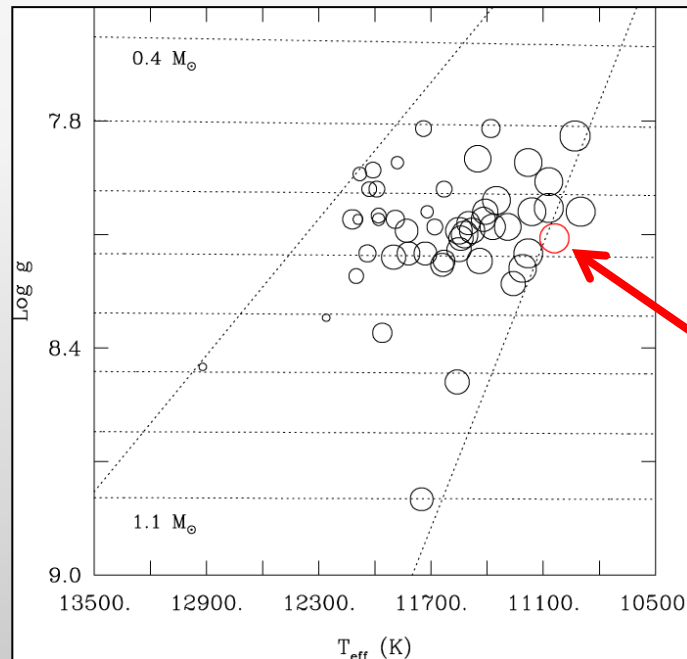
Charpinet S. *Université Toulouse 3, IRAP*  
Greiss S. *Warwick university*

# ZZCETI INSTABILITY STRIP

- Why GD1212?
  - We have data
  - Simplest of the “complicated” DAVs



Handler et al. (2008)

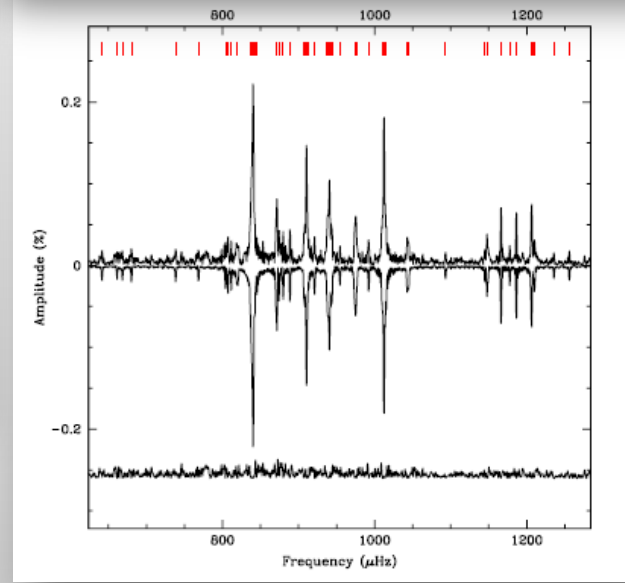
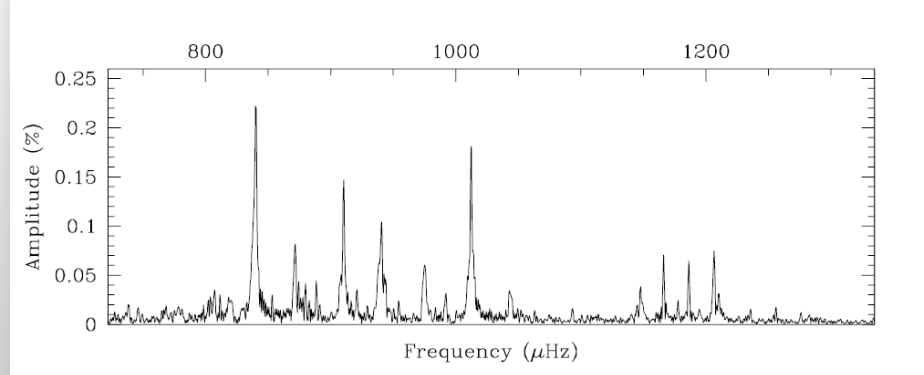
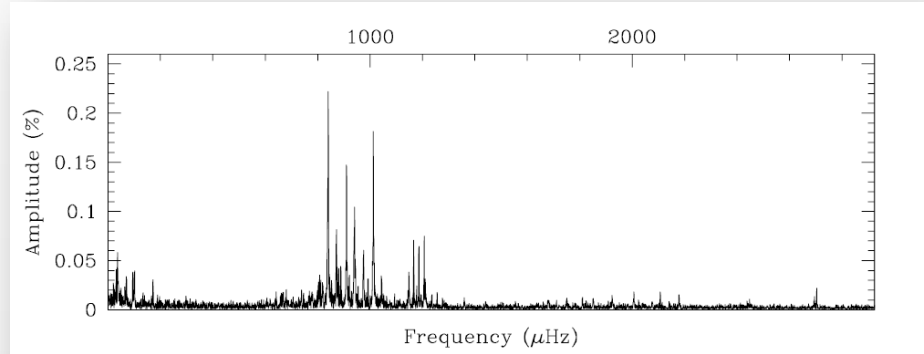
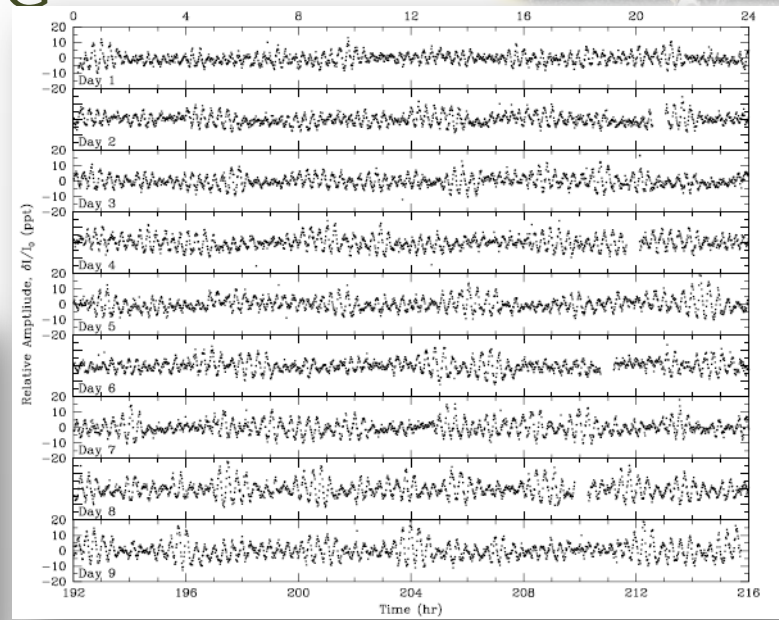


Fontaine et al. (2008)

# LIGHT CURVE PROCESSING



- Kepler K2 test run
- 9 days



See also *Hermes et al. 2014*  
and poster 65



Table 3: Final list of significant frequencies above  $4\sigma$  for the 9d run. Bracketted frequencies may be unsure.

Id.	Frequency ( $\mu\text{Hz}$ )	$\sigma_f$ ( $\mu\text{Hz}$ )	Period (s)	$\sigma_P$ (s)	Amplitude (%)	$\sigma_A$ (%)	Phase (s)	$\sigma_{Ph}$ (s)	S/N	Comments
$f_{048}$	20.6931	0.1402	48325.2773	327.3270	0.0301	0.0059	0.2202	0.0270	5.1	f1-f21
$f_{054}$	25.1616	0.1624	39743.1002	256.5863	0.0259	0.0059	0.5367	0.0313	4.4	f42-f21
$f_{032}$	32.4511	0.1046	30815.5820	99.2811	0.0402	0.0059	0.5069	0.0200	6.8	f7-f3
$f_{015}$	36.4081	0.0654	27466.4281	49.3249	0.0641	0.0059	0.5414	0.0063	10.9	
$[f_{027}]$	[38.1695]	[0.0965]	[26198.9033]	[66.2164]	[0.0433]	[0.0059]	[0.9670]	[0.0095]	[7.4]	f2-f10?; blend with f15
$[f_{046}]$	[40.1084]	[0.1342]	[24932.4336]	[83.4495]	[0.0310]	[0.0059]	[0.1985]	[0.0259]	[5.3]	f4-f7?; blend with f15
$f_{057}$	64.7091	0.1666	15453.7724	39.7892	0.0238	0.0056	0.4011	0.0335	4.3	f4-f6 ?
$f_{049}$	69.8447	0.1361	14317.4751	27.8906	0.0292	0.0056	0.6677	0.0139	5.2	f2-f5 ?
$f_{053}$	71.4853	0.1494	13988.8825	29.2276	0.0265	0.0056	0.6006	0.0152	4.8	f4-f1 ?
$f_{036}$	94.5555	0.1028	10575.8013	11.5015	0.0371	0.0054	0.5161	0.0213	6.9	f2-f31 ?
$[f_{047}]$	[100.2178]	[0.1265]	[9978.2649]	[12.5987]	[0.0303]	[0.0054]	[0.4270]	[0.0130]	[5.6]	f5-f1?; blend with f33
$f_{033}$	101.6019	0.1006	9842.3383	9.7465	0.0382	0.0054	0.5418	0.0103	7.1	f31-f21 ?
$f_{044}$	171.8924	0.1091	5817.5941	3.6936	0.0325	0.0050	0.7269	0.0241	6.5	f2-f1 ?
$f_{089}$	641.6091	0.1575	1558.5814	0.3826	0.0181	0.0040	0.0048	0.0433	4.5	
$[f_{073}]$	[661.4700]	[0.1769]	[1511.7843]	[0.4043]	[0.0172]	[0.0043]	[0.4423]	[0.0460]	[4.0]	several weak peaks around
$[f_{072}]$	[668.5275]	[0.1724]	[1495.8247]	[0.3858]	[0.0179]	[0.0043]	[0.3825]	[0.0442]	[4.1]	*; several weak peaks around
$f_{085}$	680.2633	0.1554	1470.0190	0.3358	0.0199	0.0044	0.0097	0.0395	4.6	
$f_{064}$	738.4715	0.1696	1354.1484	0.3110	0.0202	0.0048	0.8138	0.0390	4.2	
$f_{061}$	768.4389	0.1692	1301.3396	0.2866	0.0222	0.0053	0.9265	0.0354	4.2	*
$f_{051}$	803.9796	0.1447	1243.8127	0.2239	0.0272	0.0055	0.8632	0.0311	4.9	*
$f_{035}$	807.1185	0.1049	1238.9754	0.1611	0.0375	0.0055	0.1074	0.0233	6.8	*
$f_{082}$	811.5749	0.1788	1232.1722	0.2714	0.0220	0.0055	0.9186	0.0373	4.0	*
$f_{021}$	819.9788	...	1219.5437	...	0.0536	...	0.6406	...	9.6	
$f_{001}$	840.2108	...	1190.1774	...	0.2477	...	0.9957	...	43.4	
$[f_{042}]$	[844.8180]	[0.1201]	[1183.6868]	[0.1683]	[0.0333]	[0.0056]	[0.1009]	[0.0243]	[5.9]	*; may be linked to f1
$f_{007}$	871.6296	...	1147.2763	...	0.0866	...	0.7558	...	15.3	
$f_{030}$	874.0884	0.0992	1144.0491	0.1298	0.0406	0.0057	0.8393	0.0209	7.2	*
$f_{023}$	879.8627	0.0827	1136.5410	0.1068	0.0486	0.0057	0.0691	0.0166	8.6	
$f_{029}$	888.5756	0.0946	1125.3967	0.1198	0.0416	0.0055	0.7081	0.0191	7.5	
$f_{004}$	910.4287	...	1098.3837	...	0.1513	...	0.2721	...	28.7	
$f_{031}$	921.1198	0.0936	1085.6351	0.1103	0.0402	0.0053	0.6656	0.0197	7.6	
$f_{005}$	940.7330	...	1063.0009	...	0.1099	...	0.1926	...	21.1	
$f_{080}$	954.4381	0.1667	1047.7369	0.1829	0.0223	0.0052	0.4139	0.0354	4.3	*
$f_{010}$	975.3833	...	1025.2379	...	0.0740	...	0.5623	...	14.1	
$f_{039}$	992.3267	0.1021	1007.7326	0.1037	0.0358	0.0052	0.1987	0.0220	7.0	
$f_{002}$	1012.3867	...	987.7648	...	0.1846	...	0.5738	...	37.0	
$f_{040}$	1043.7635	...	958.0714	...	0.0348	...	0.6009	...	7.1	*
$f_{076}$	1093.3840	0.1962	914.5918	0.1641	0.0162	0.0045	0.2470	0.0485	3.6	probably higher S/N
$[f_{071}]$	[1145.0845]	[0.1522]	[873.2980]	[0.1161]	[0.0180]	[0.0039]	[0.9228]	[0.0453]	[4.7]	*; may be linked to f34
$f_{034}$	1147.7071	...	871.3025	...	0.0380	...	0.1785	...	9.9	
$f_{011}$	1166.2026	0.0367	857.4839	0.0270	0.0722	0.0037	0.7412	0.0109	19.3	
$f_{059}$	1177.7694	0.1146	849.0626	0.0826	0.0226	0.0036	0.2073	0.0350	6.2	*
$f_{014}$	1186.3756	0.0394	842.9034	0.0280	0.0643	0.0036	0.9827	0.0123	18.0	
$f_{012}$	1206.6583	...	828.7351	...	0.0715	...	0.7307	...	21.5	
$f_{008}$	1210.3198	...	826.2279	...	0.0329	...	0.4581	...	9.9	

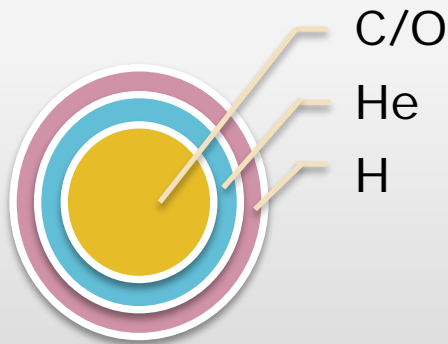
13 independent modes extracted (for structure)

19 components of rotationally split multiplets (for rotation)

$f_{084}$	2447.9299	0.1386	408.5084	0.0231	0.0111	0.0022	0.1256	0.0706	5.1	
$f_{079}$	2695.3338	0.1034	371.0116	0.0142	0.0136	0.0020	0.0507	0.0577	6.9	
$f_{083}$	2703.8441	0.0647	369.8438	0.0089	0.0217	0.0020	0.9544	0.0363	11.0	

# THE FORWARD METHOD

- Comparison:
  - observed frequencies
  - Theoretical frequencies from static, parameterized models

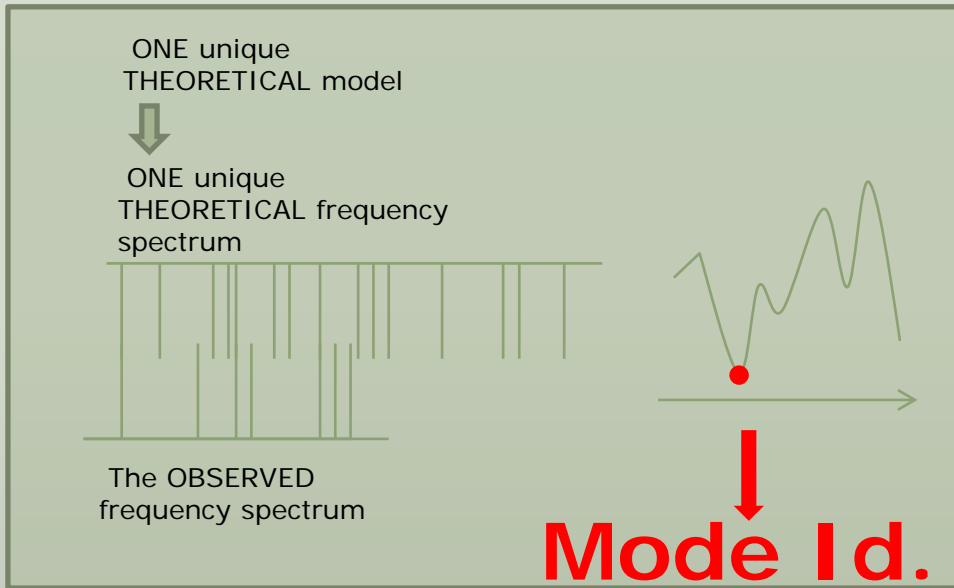


$T_{eff}$	Effective temp.
$\log g$	Surface gravity
$q_H$	H layer mass
$q_{He}$	He layer mass
<i>Core comp.</i>	Homogeneous C-O mix
$Pf_1$	Parameterized param. for the chemical profile at H/He transition
$Pf_2$	Parameterized param. for the chemical profile at He/C-O transition
<i>Conv.</i>	Convective efficiency (MLT)

# TECHNIQUE: DOUBLE-SCHEME OPTIMISATION

$T_{eff}, \log g, q_H, q_{He},$  core  
comp.,  $Pf_1, Pf_2, \dots$

Using a multimodal optimizer based on a genetic algorithm

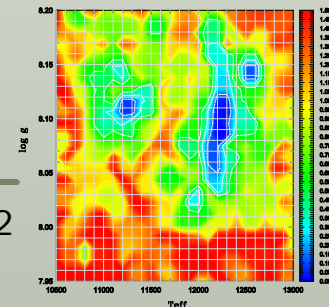


**An optimal model:**

$T_{eff}, \log g, q_H, q_{He},$  core  
comp.,  $Pf_1,$   
 $Pf_2$



Merit function  $S^2$

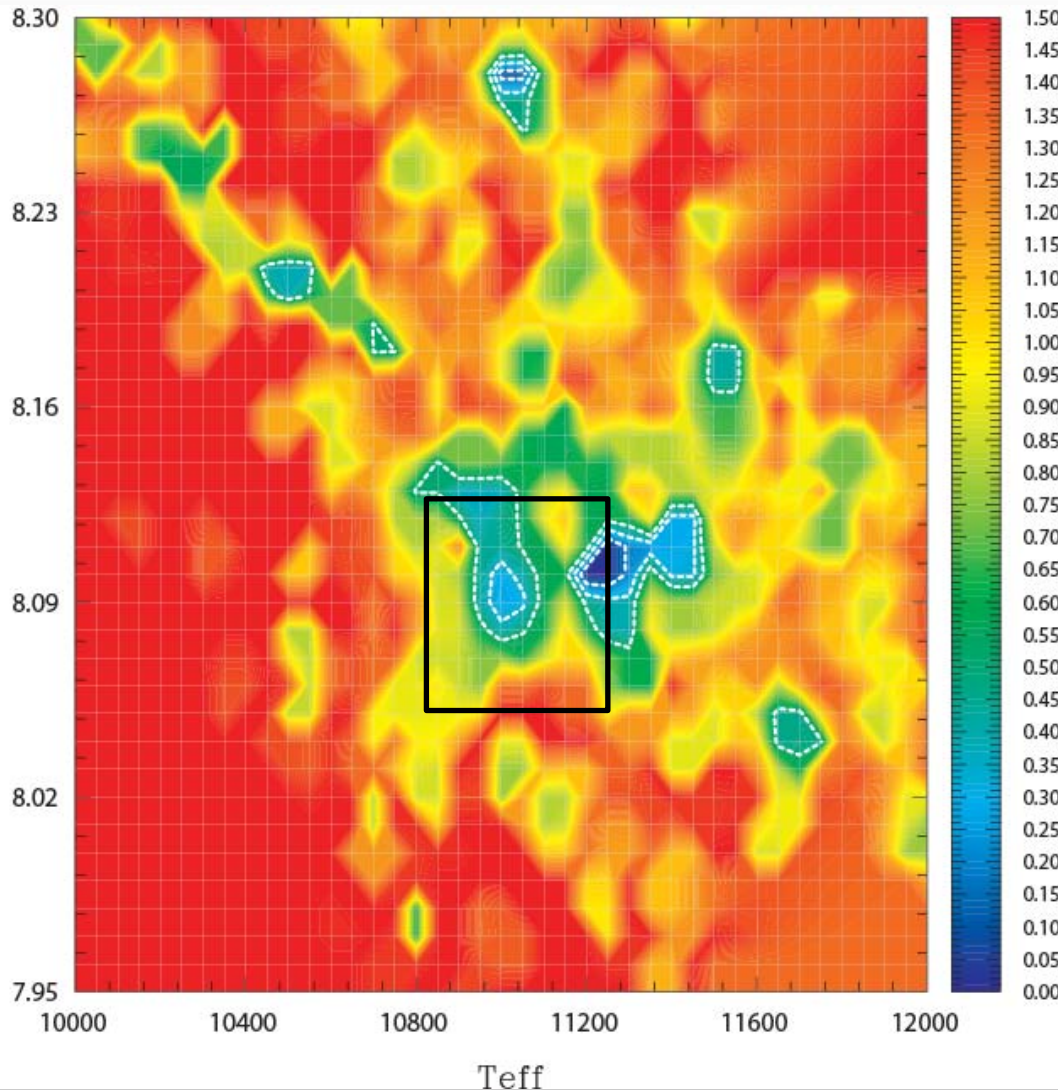


# WHAT CAN WE INFER?

- Independent measure of:
  - Surface gravity (Mass),
  - Temperature,
  - Layering: D(H), D(He)
  
- Extra:
  - Bulk core composition
  - Internal rotation profile



# SOLUTION: PROJECTION LOG G - $T_{\text{EFF}}$ PLANE



**Spectroscopic solution:**  
 $T_{\text{eff}} = 11,035 \pm 194$  K  
 $\text{Log } g = 8.08 \pm 0.05$

**Sismic solution:**  
 $T_{\text{eff}} = 11,244$  K  
 $\text{Log } g = 8.10$  ( $0.65 M_{\odot}$ )

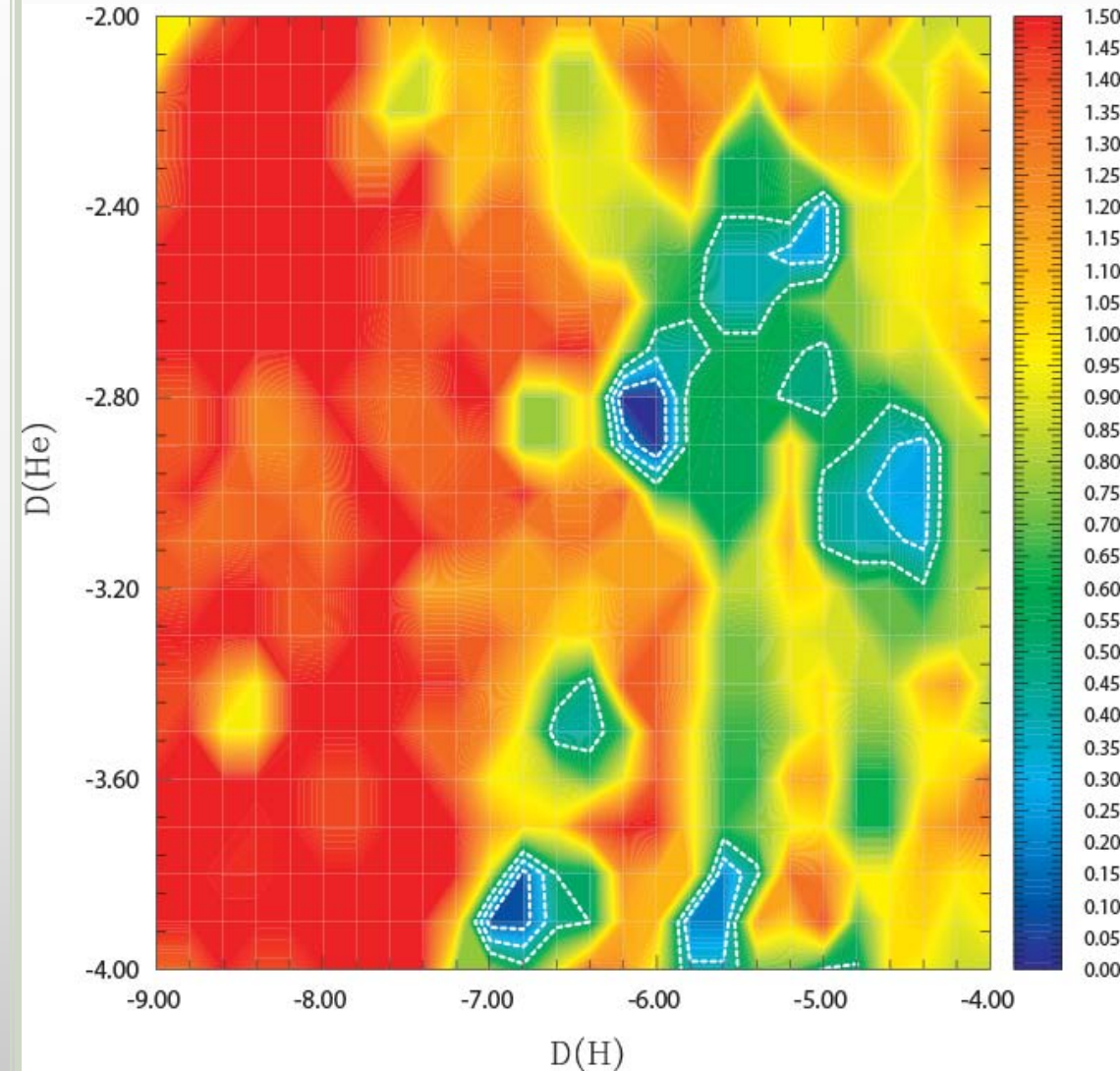
Error estimate:  
200 K  
0.01 dex

Color scale ->  $\log S^2$





# SOLUTION: PROJECTION $D(H) - D(He)$ PLANE



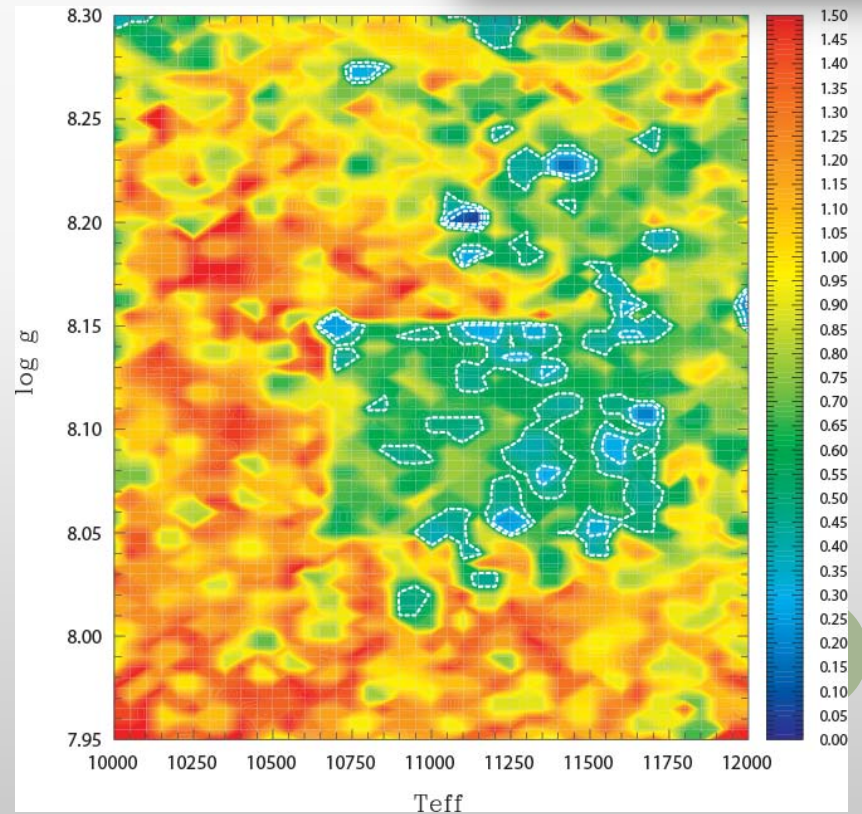
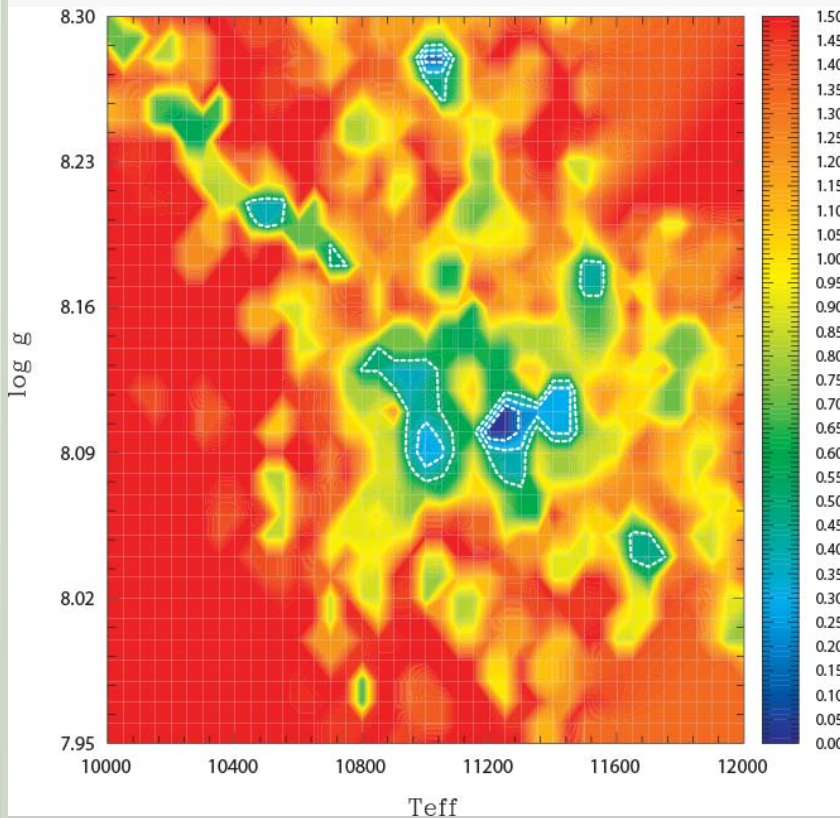
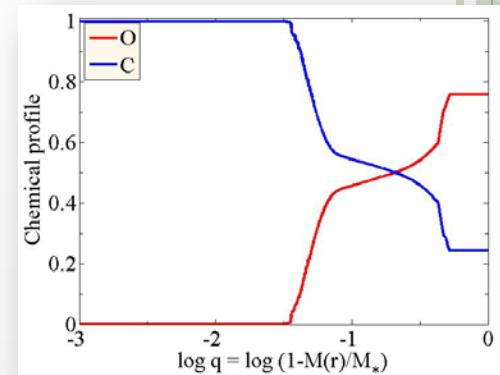
**Sismic solution:**  
 $D(H) = -6.04$   
 $D(He) = -2.82$

Color scale  $\rightarrow \log S^2$

# HOMOGENEOUS VS NON-HOMOGENEOUS CORE

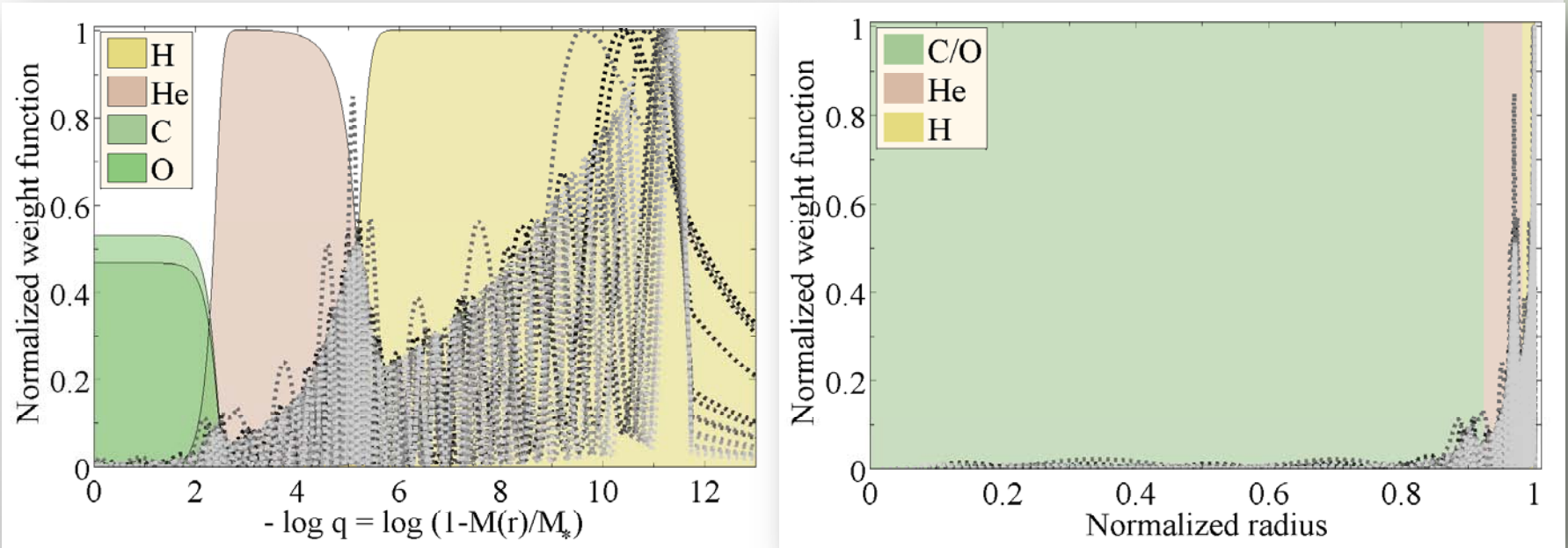
○  $S^2 = 28$

$S^2 = 252$



# WHICH PARTS OF THE STAR CONTRIBUTE THE MOST TO THE PULSATIONS?

- Weight functions:
  - Determination of the bulk core composition

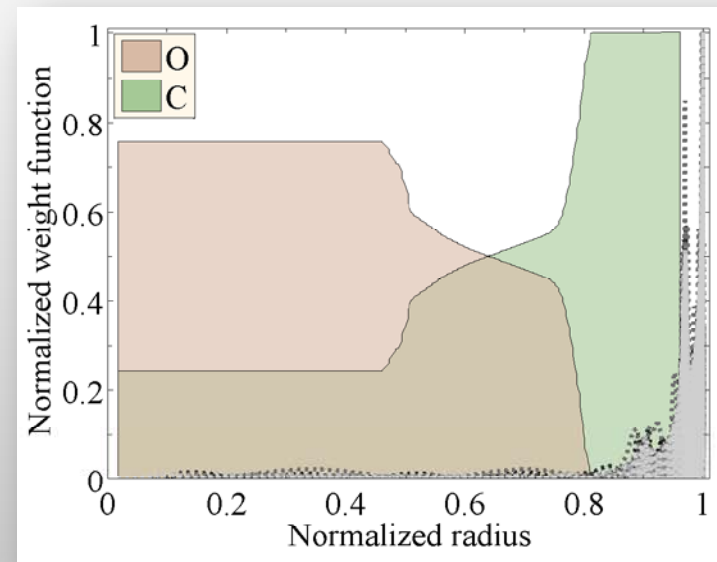
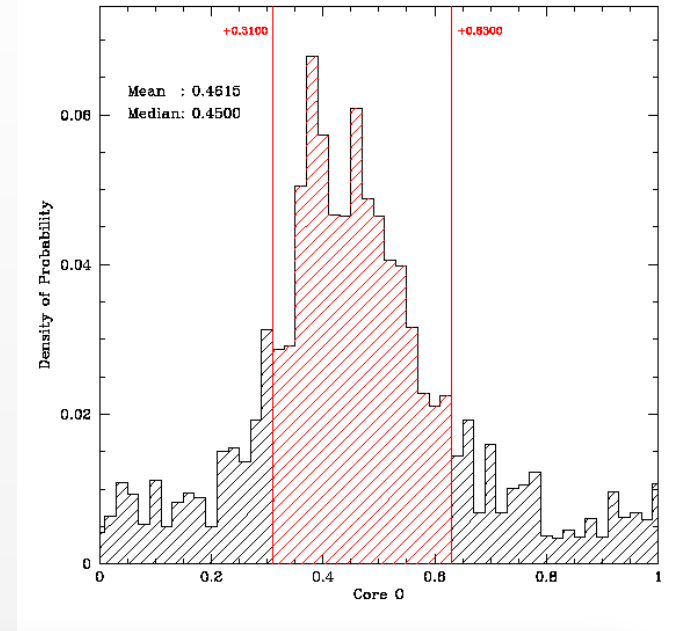


- C/O = 47-53%



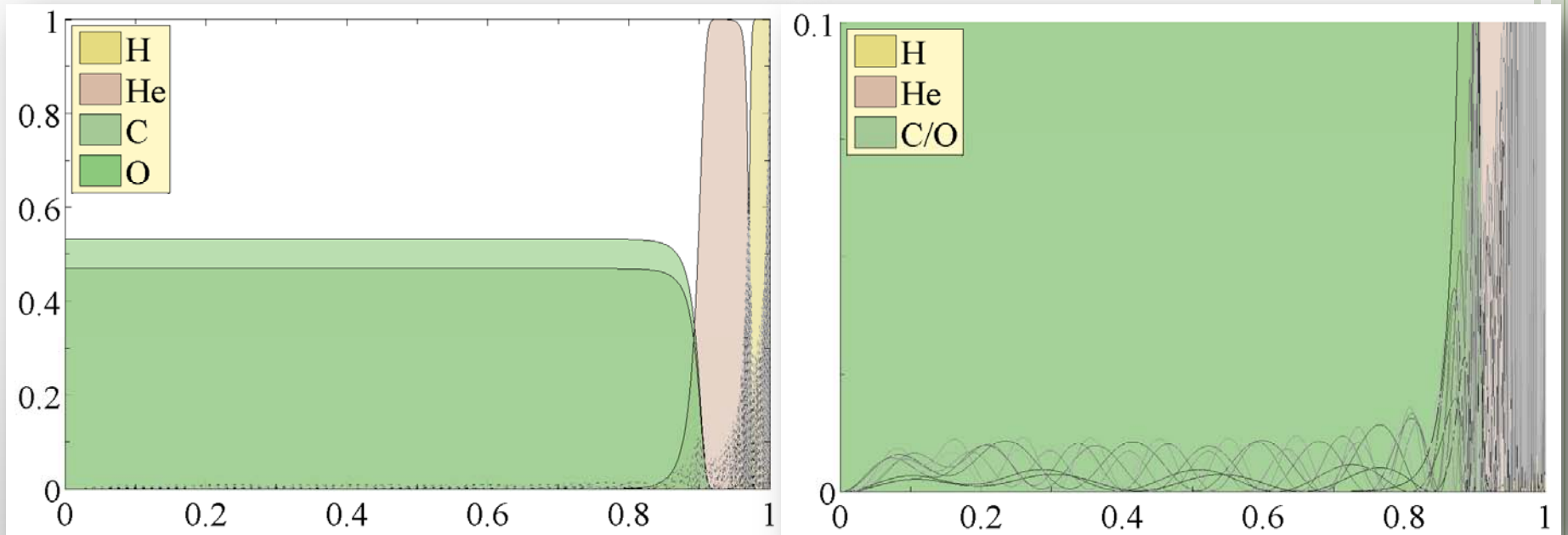
# COMPARED TO EVOLUTIONARY MODEL, WHERE DO WE STAND?

- What we find:
  - $C/O = 47-53\%$
- Compared to profiles from Salaris et al. (2010)
  - $C/O = 43-58\%$

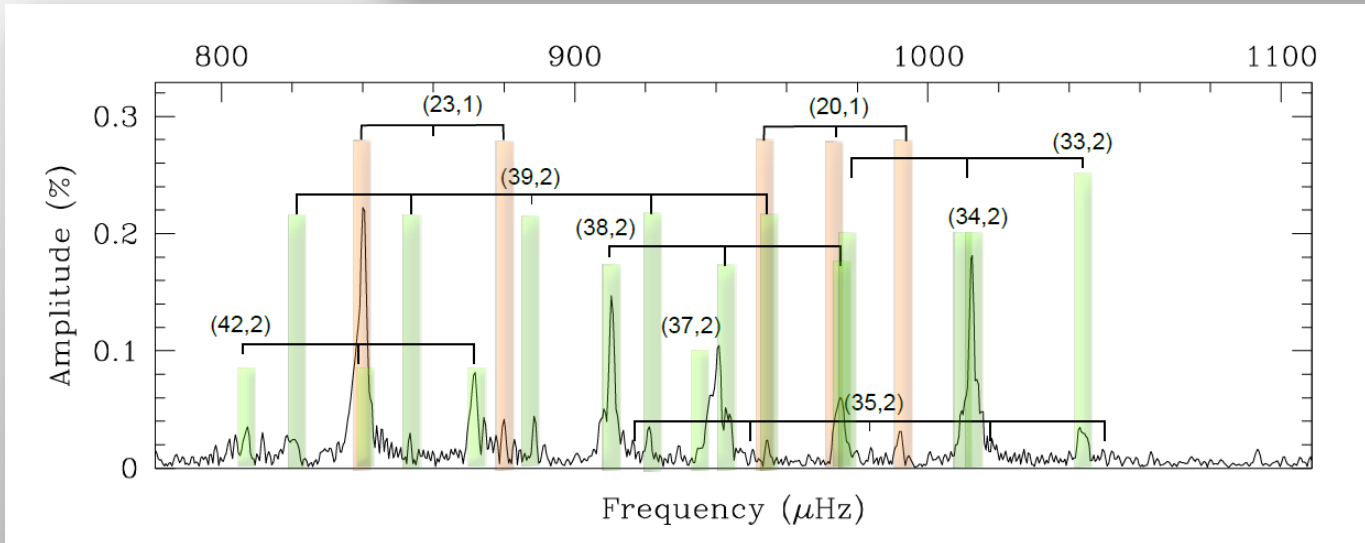
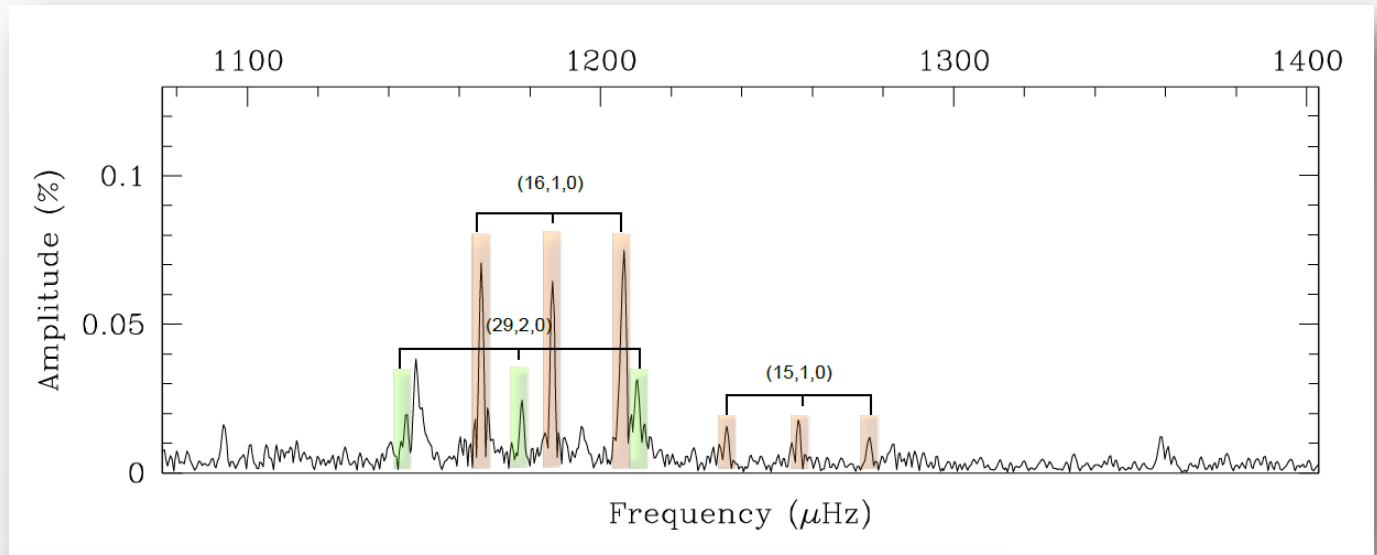


# GD1212: ROTATION KERNELS

- Going deep into the interior

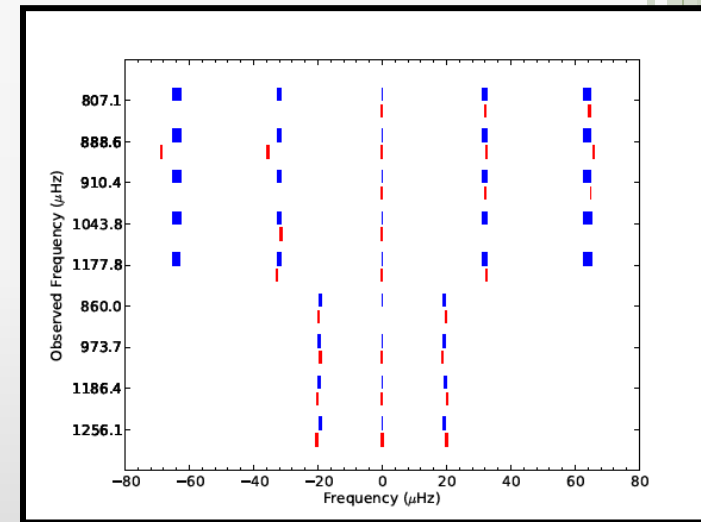
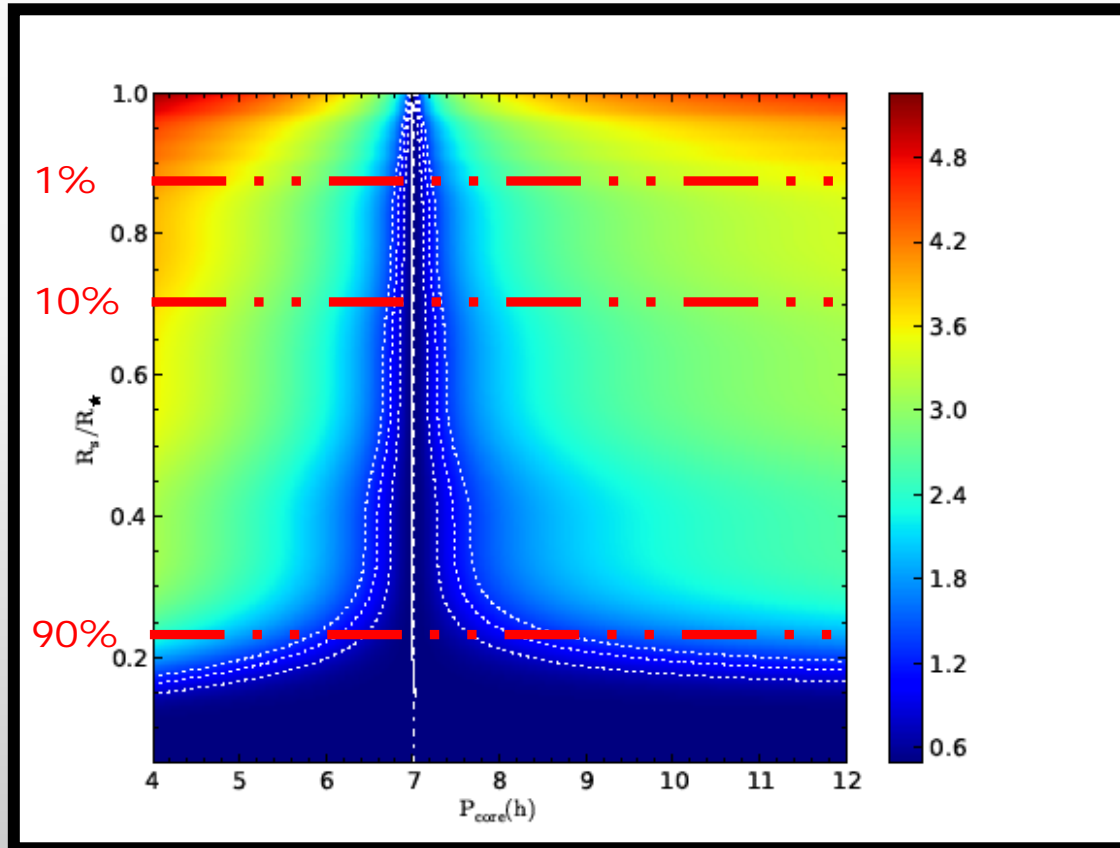
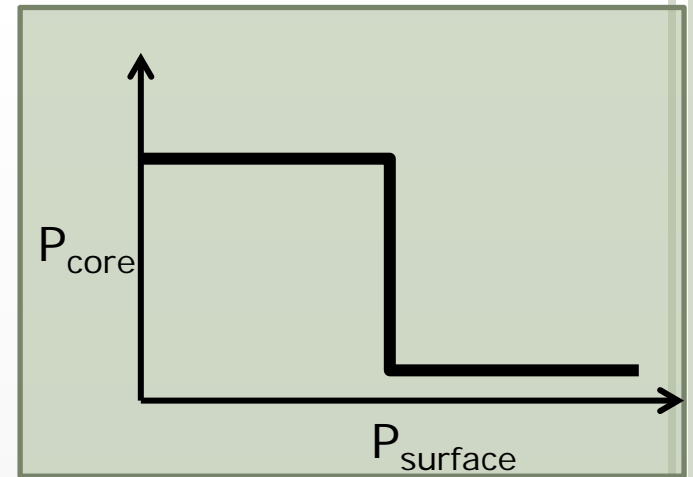


# GD1212: ROTATIONAL SPLITTINGS



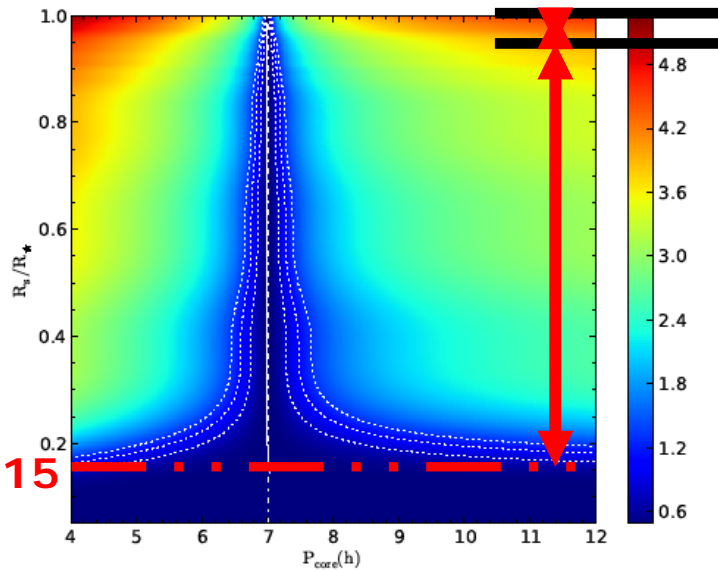
# INTERNAL ROTATION PROFILE

- Solid-body rotation test

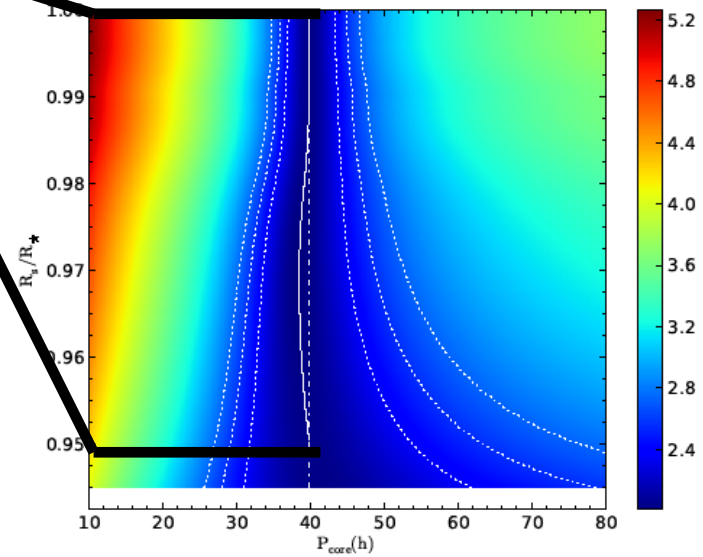


# HOW DOES IT COMPARE TO OTHER ZZCETI?

GD1212



R548

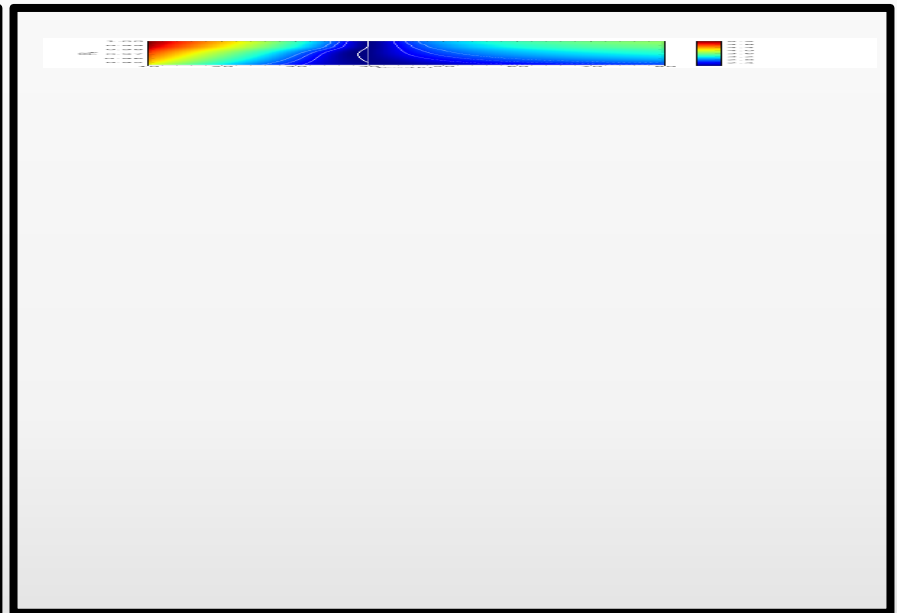
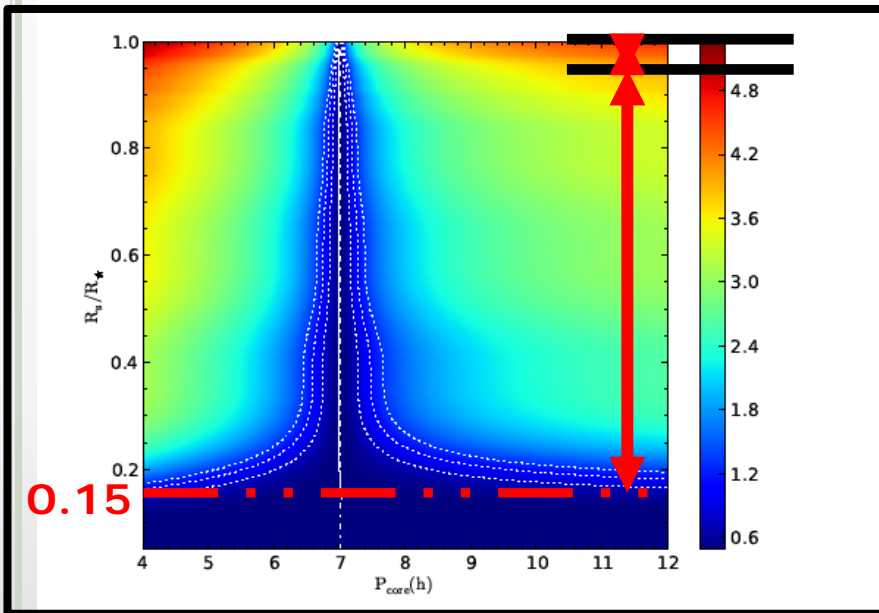




# HOW DOES IT COMPARE TO OTHER ZZCETI?

GD1212

R548



# CONCLUSION

- IT WORKS!!
- We have been able to determine for GD1212:
  - Mass (Surface gravity)
  - Temperature
  - Chemical stratification: H/He layering
- Extra:
  - Bulk core composition
  - Internal rotation profile
- Future prospects:
  - Analysis of other Kepler 1-2 WDs

*(See Greiss et al. Poster 6)*

