

Formation and evolution of planetary systems:

What have we learnt from transit methods?

Clément Baruteau

CNRS / IRAP, Toulouse

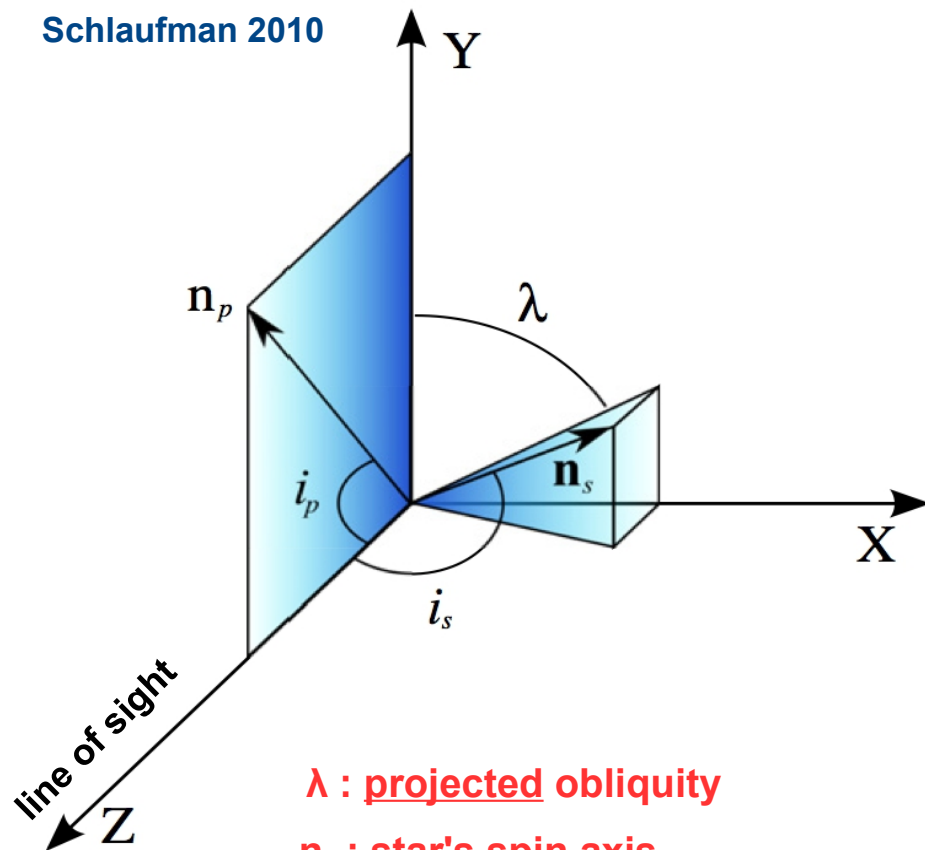
Planet formation & evolution: transits highlights

- **Misaligned hot Jupiters**
→ high-eccentricity migration, tides
- **Close-in, compact multiplanet systems**
→ in situ growth vs. disc migration
- **Circumbinary planets**
→ “extreme” planet formation
- **Planets around post main sequence stars**
→ impact of stellar evolution on close-in planets?

Talk outline

- **Misaligned hot Jupiters**
→ high-eccentricity migration, tides
- **Close-in, compact multiplanet systems**
→ in situ growth vs. disc migration
- **Circumbinary planets**
→ “extreme” planet formation
- **Planets around post main sequence stars**
→ impact of stellar evolution on close-in planets?

Planets obliquities: constraints on migration scenarios?



λ : projected obliquity

\mathbf{n}_s : star's spin axis

\mathbf{n}_p : planet's orbital axis

$\psi = \text{acos}(\mathbf{n}_s \cdot \mathbf{n}_p)$: true obliquity

□ Rossiter-MacLaughlin effect

e.g., Winn+ 2005

□ Planet-starspot crossings

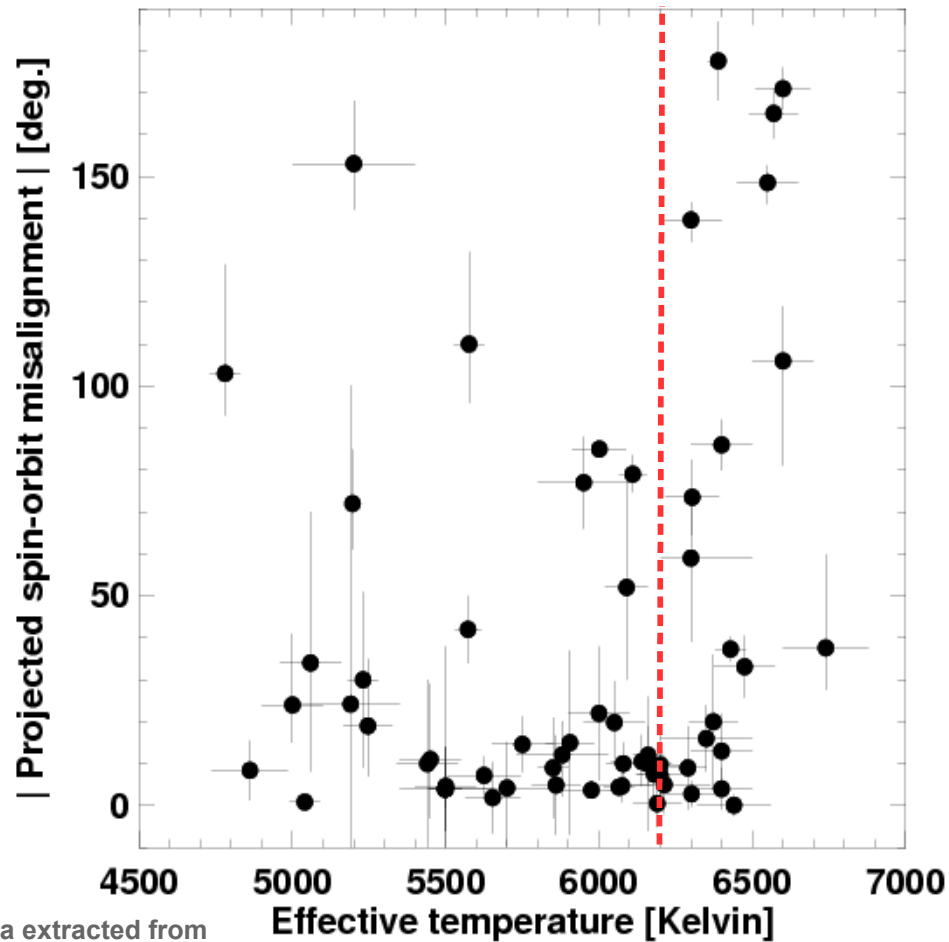
e.g., Sanchis-Ojeda+ 2011

□ Constraining true obliquity (ψ)
with stellar spin axis angle (i_s)

e.g., Huber+ 2013, Chaplin+ 2013

Planets obliquities: constraints on migration scenarios?

OBSERVATIONS



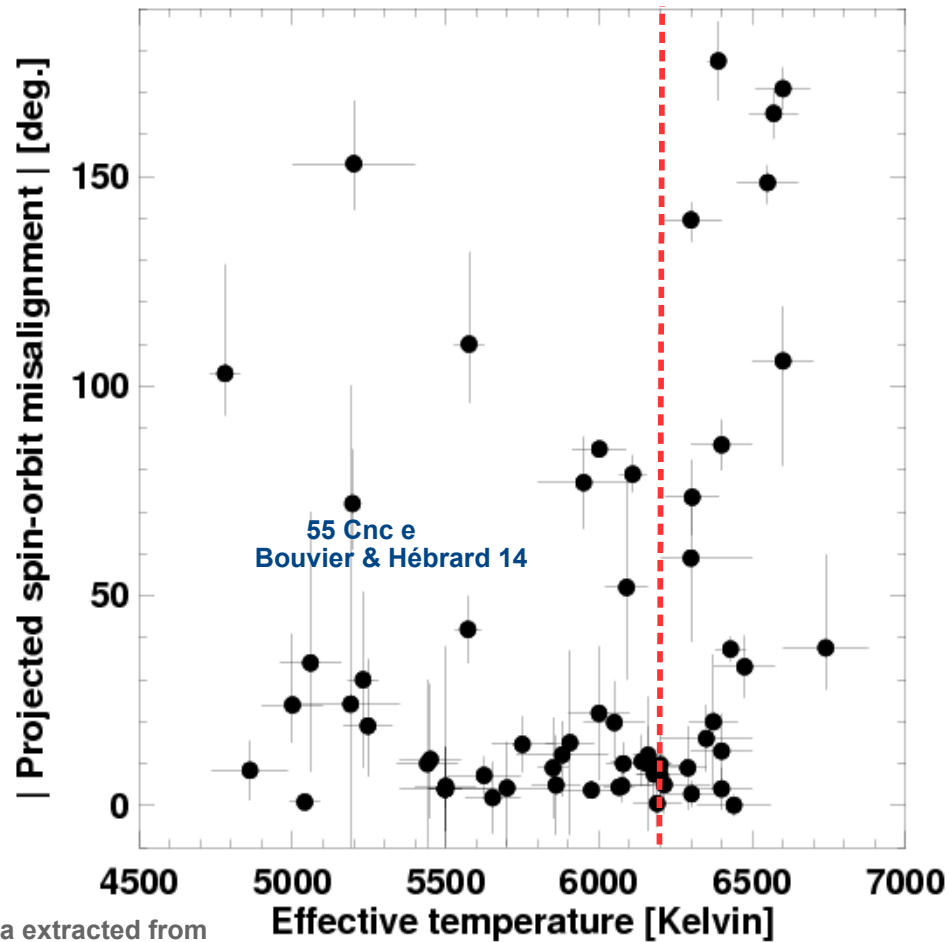
data extracted from
exoplanets.org (06/2014)

□ (Why) do hot Jupiters around hot stars tend to have high obliquities?

Planets obliquities: constraints on migration scenarios?

OBSERVATIONS

MODELS



data extracted from
exoplanets.org (06/2014)

□ **Disc migration** is a natural source of **aligned** hot Jupiters

→ disc misaligned by nearby stars?
Bate+ 2010, Batygin 2012

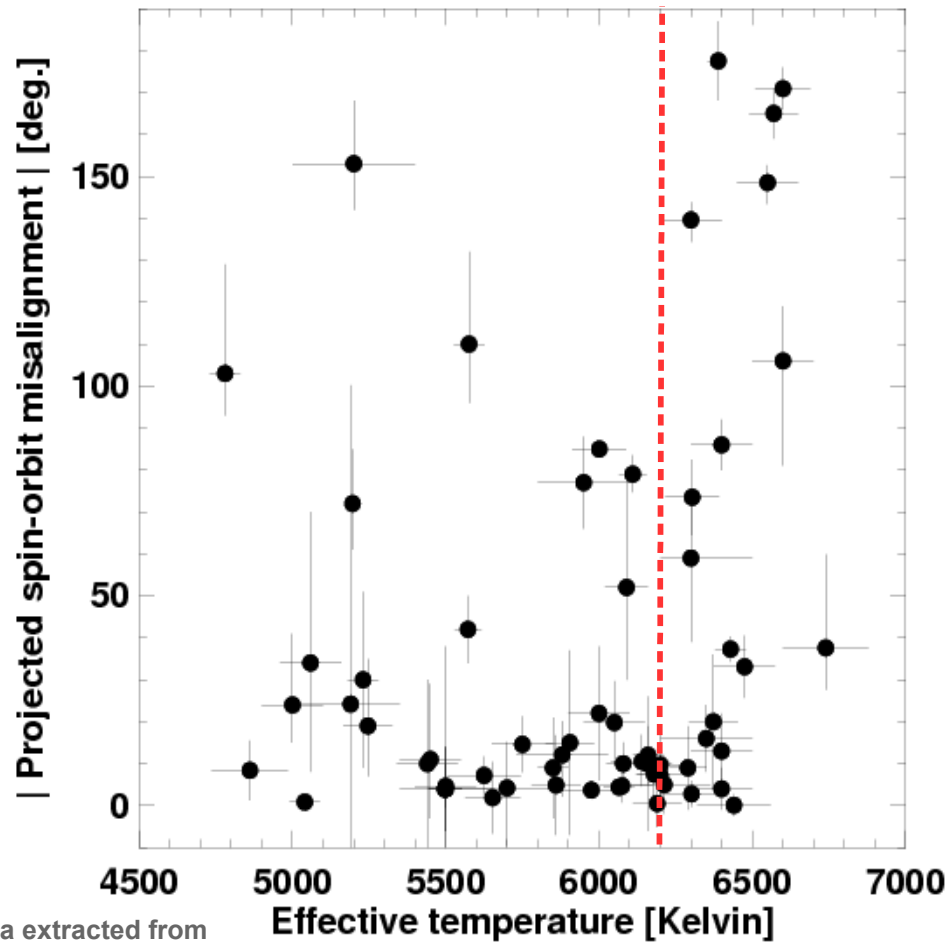
→ tidal flip of stellar axis?
Cebren+ 2013, Barker & Lithwick 2014

□ (Why) do **hot Jupiters** around **hot stars** tend to have **high obliquities**?

Planets obliquities: constraints on migration scenarios?

OBSERVATIONS

MODELS



□ **Disc migration** is a natural source of **aligned** hot Jupiters

→ disc misaligned by nearby stars?

Bate+ 2010, Batygin 2012

→ tidal flip of stellar axis?

Cebtron+ 2013, Barker & Lithwick 2014

□ **High-eccentricity migration** followed by **tidal circularization** is a natural source of **misaligned** hot Jupiters

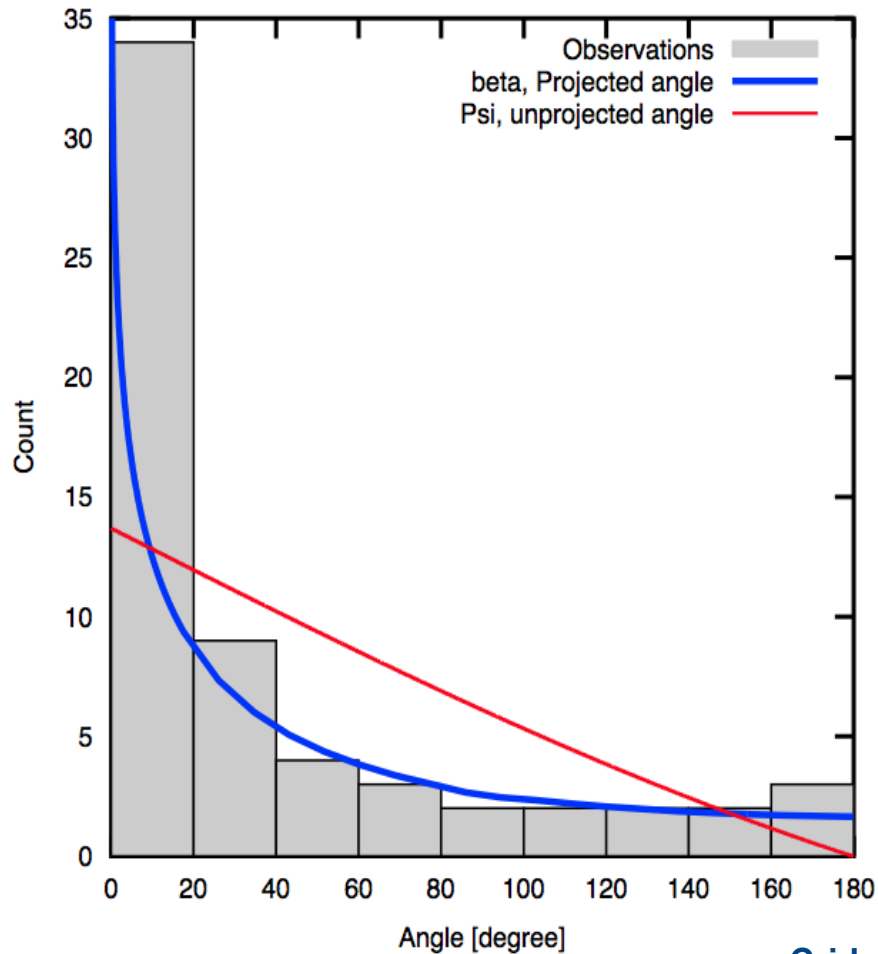
→ can **all** hot Jupiters form this way?

Triaud+ 2010, Albrecht + 2012, but see
Rogers & Lin 2013, Lai 2012

□ (Why) do **hot Jupiters** around **hot stars** tend to have **high obliquities**?

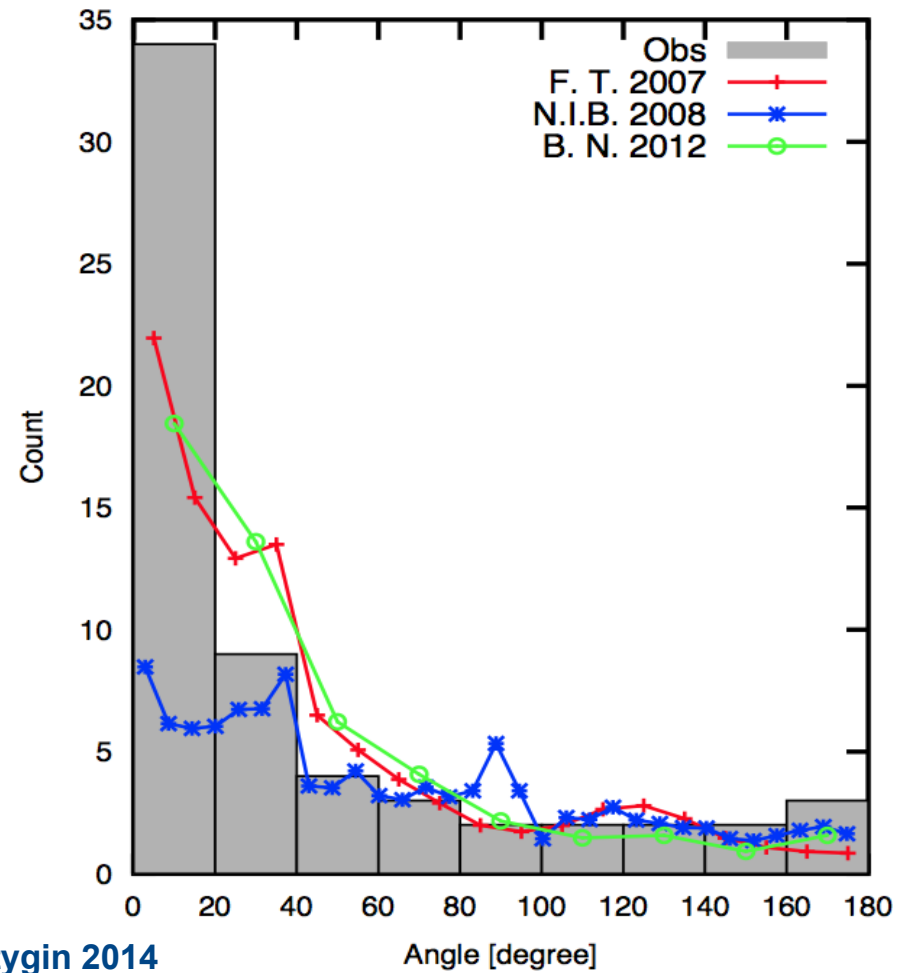
Planets obliquities: constraints on migration scenarios?

□ Disc misaligned by nearby stars



Crida & Batygin 2014

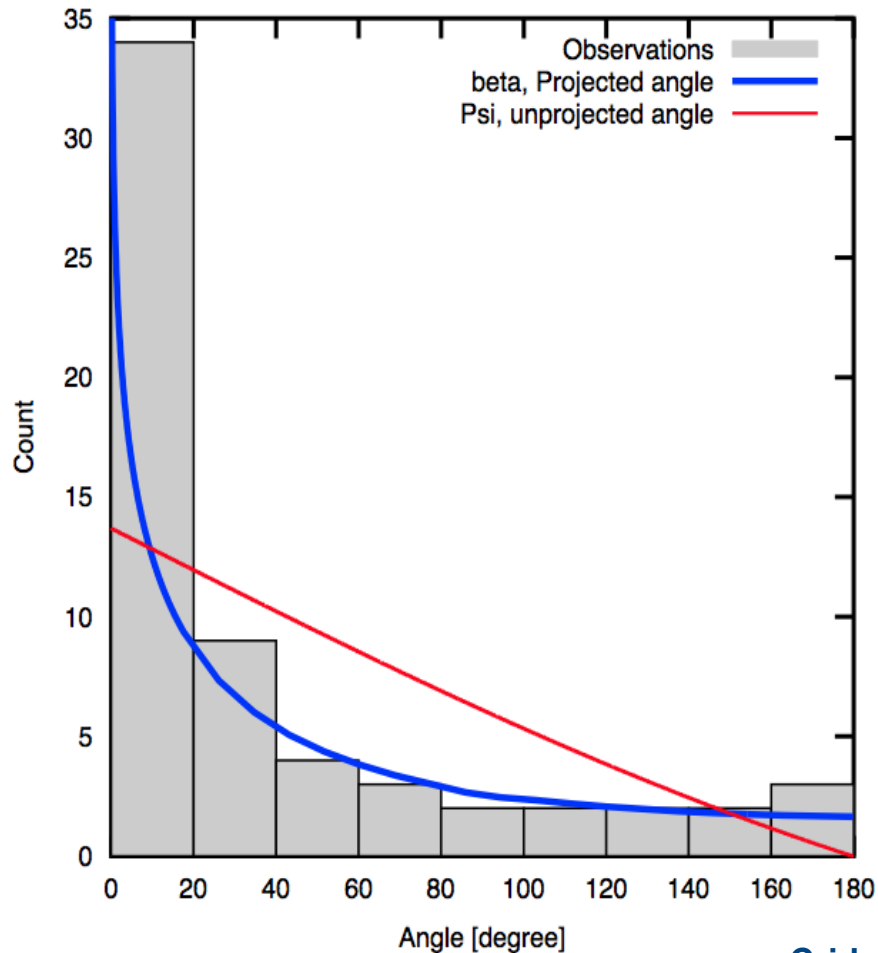
□ High-eccentricity migration + tides



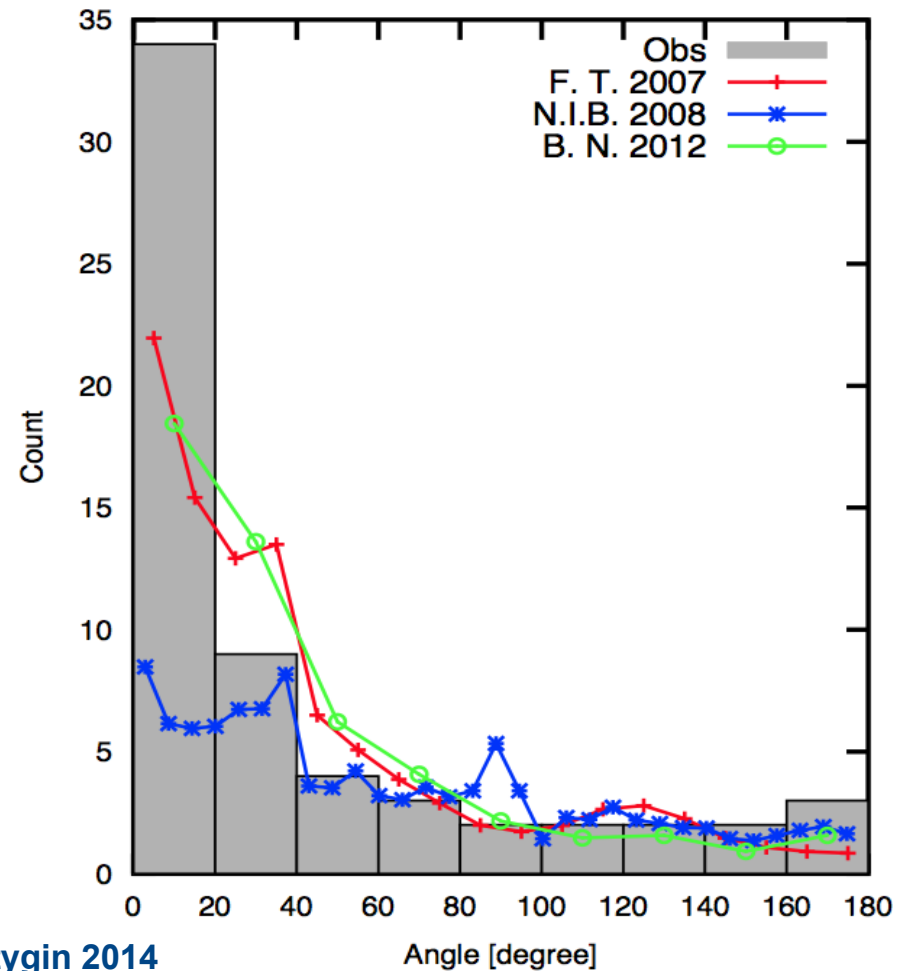
Angle [degree]

Planets obliquities: constraints on migration scenarios?

□ Disc misaligned by nearby stars



□ High-eccentricity migration + tides



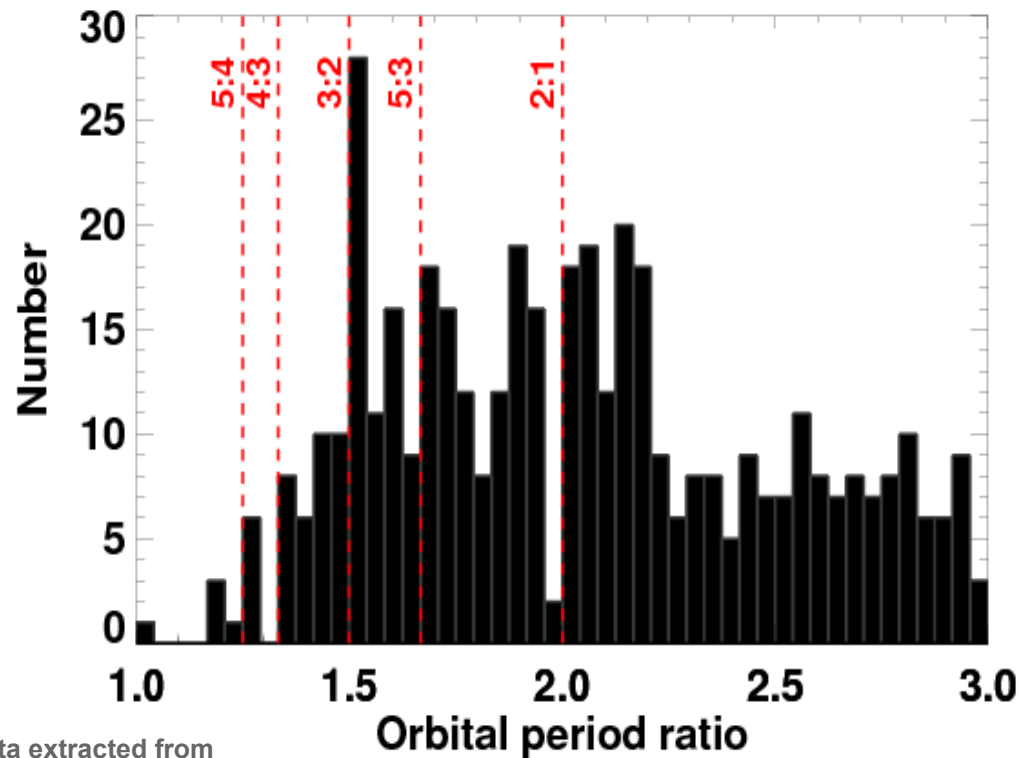
→ disc migration needed to account for number of aligned hot Jupiters.

→ observations cannot (yet) distinguish between misalignment mechanisms

Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

→ 836 planets in 335 systems
66% of 2 planets, 22% of 3, 12% of 4 and more



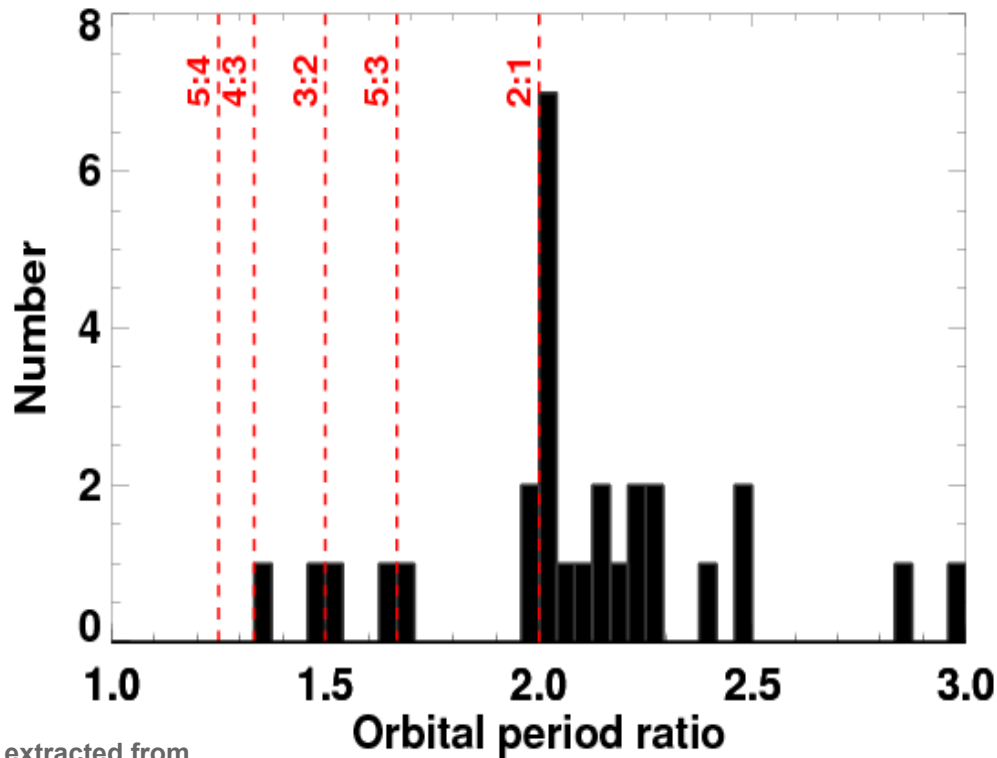
data extracted from
exoplanets.org (06/2014)

- Many planet pairs are *not* in resonance, but those near resonances tend to have period ratios slightly **greater** than resonant

Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

→ 171 planets in 68 systems
66% of 2 planets, 23% of 3, 11% of 4 and more



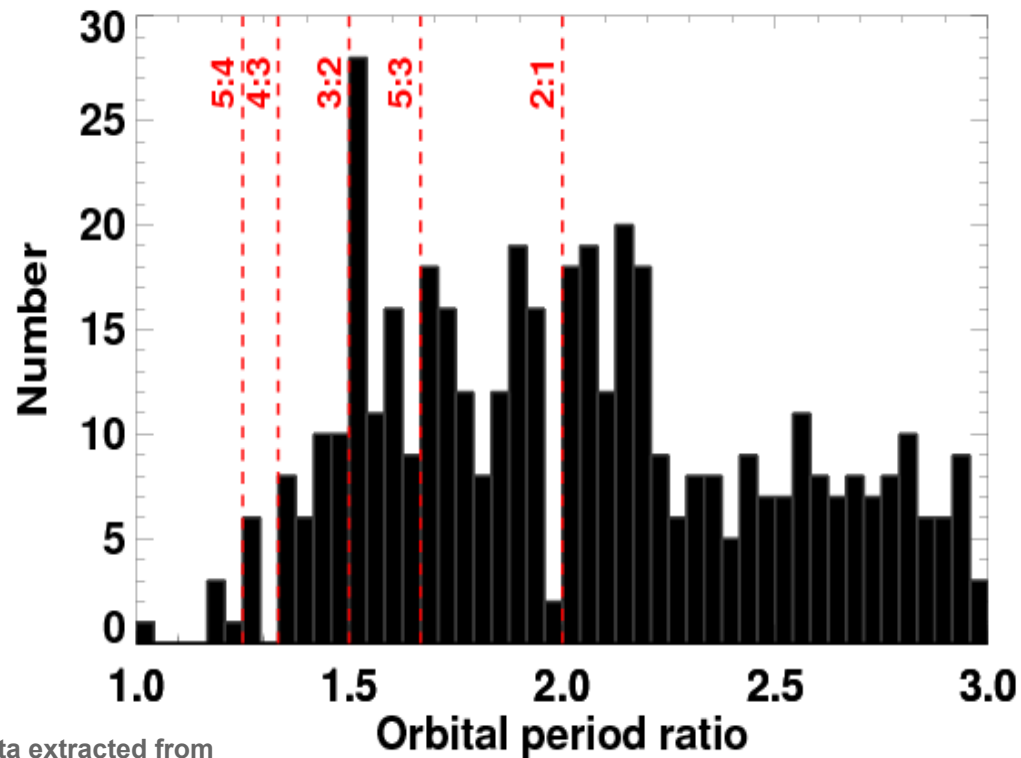
data extracted from
exoplanets.org (06/2014)

□ Same trend for RV-detected multiple systems?

Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

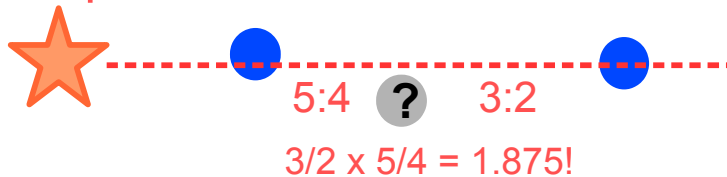
→ 836 planets in 335 systems
66% of 2 planets, 22% of 3, 12% of 4 and more



data extracted from
exoplanets.org (06/2014)

□ Unseen companions?

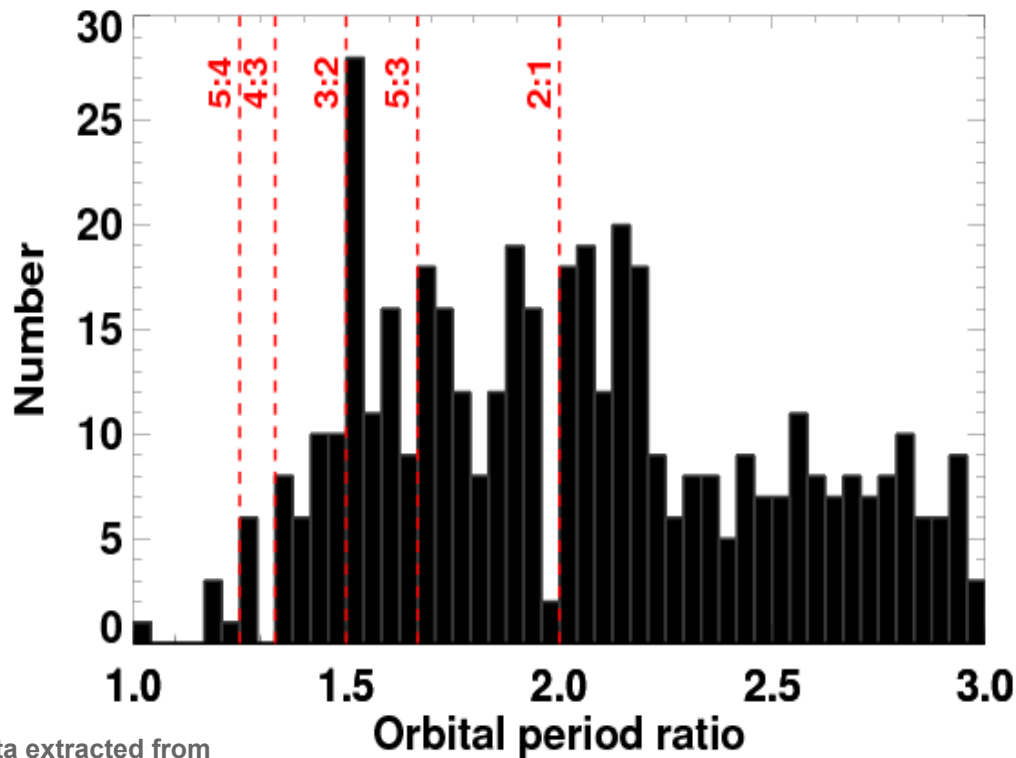
Steffen 2013



Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

→ 836 planets in 335 systems
66% of 2 planets, 22% of 3, 12% of 4 and more



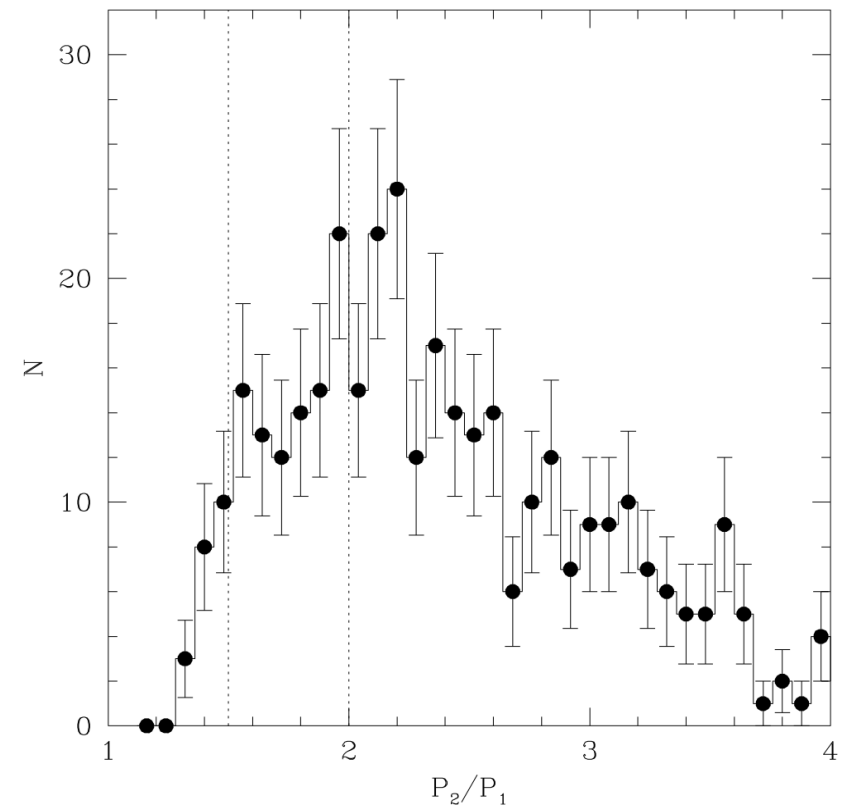
data extracted from
exoplanets.org (06/2014)

□ Many planet pairs are *not* in resonance, but those near resonances tend to have period ratios slightly **greater** than resonant

MODELS

□ In-situ growth of planet embryos

Hansen & Murray 2013; see also Raymond & Cossou 2014 and talk by Elisa Quintana

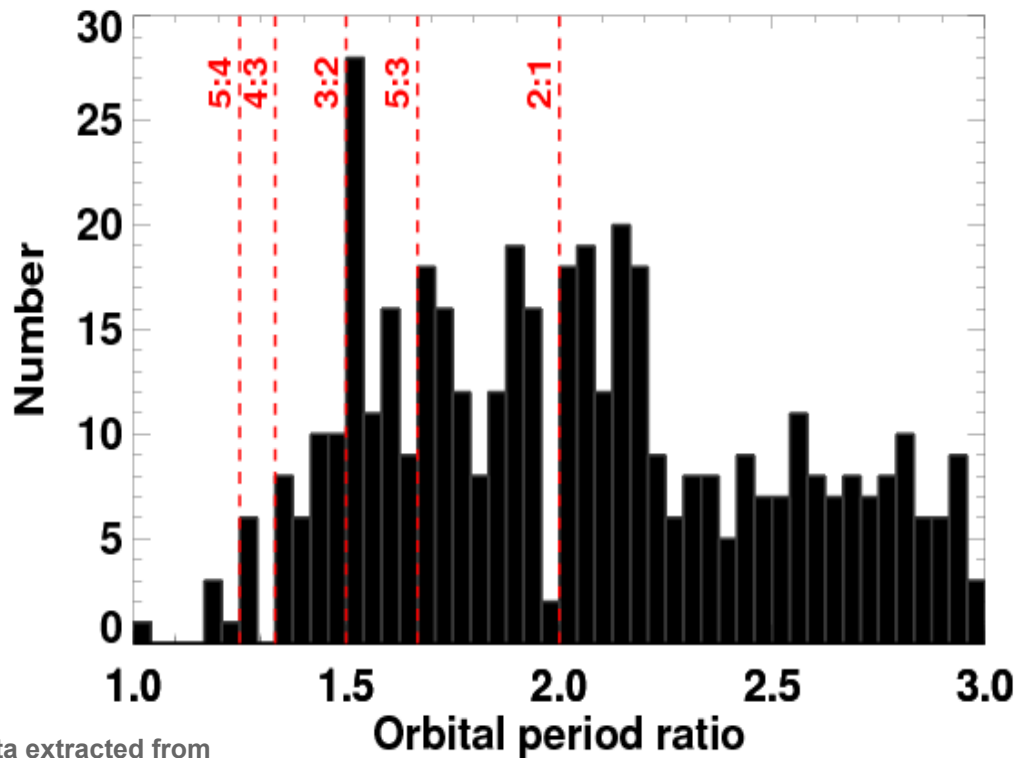


Hansen & Murray 2013

Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

→ 836 planets in 335 systems
66% of 2 planets, 22% of 3, 12% of 4 and more



data extracted from
exoplanets.org (06/2014)

□ Many planet pairs are *not* in resonance, but those near resonances tend to have period ratios slightly **greater** than resonant

MODELS

□ In-situ growth of planet embryos

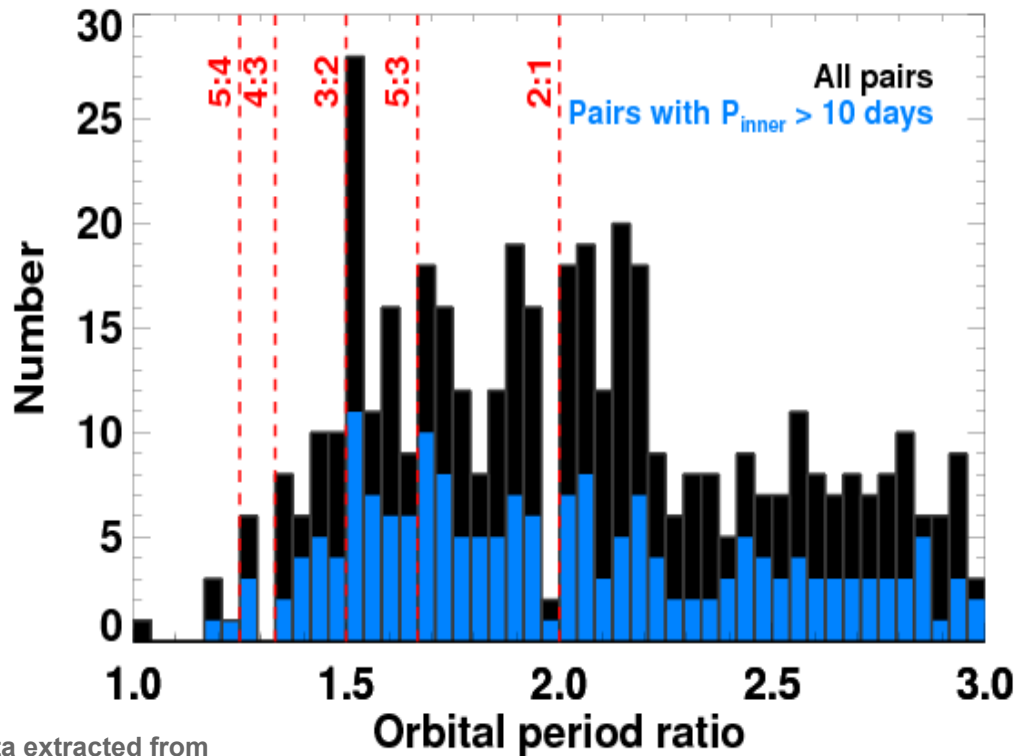
□ Tidal dissipation of close-in resonant planetary systems

Papaloizou 2011, Lithwick & Wu 2012, ...

Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

→ 836 planets in 335 systems
66% of 2 planets, 22% of 3, 12% of 4 and more



data extracted from
exoplanets.org (06/2014)

□ Many planet pairs are *not* in resonance, but those near resonances tend to have period ratios slightly **greater** than resonant

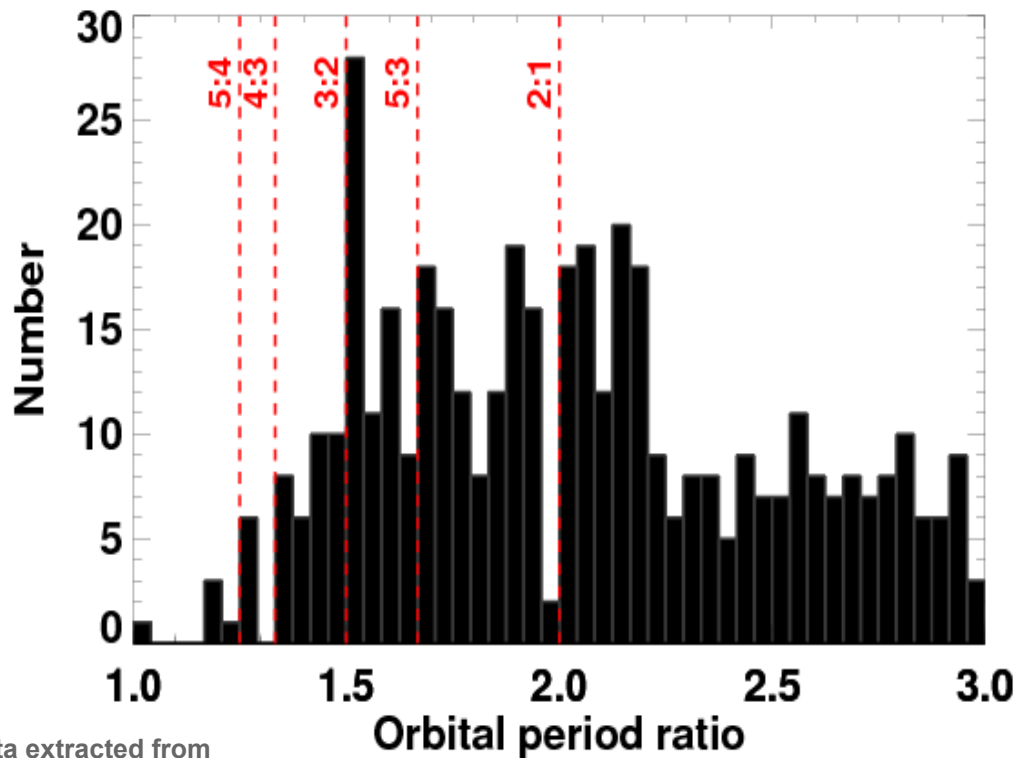
MODELS

- In-situ growth of planet embryos
- Tidal dissipation of close-in resonant planetary systems
Papaloizou 2011, Lithwick & Wu 2012, ...

Architecture of *Kepler's* confirmed multiplanet systems

OBSERVATIONS

→ 836 planets in 335 systems
66% of 2 planets, 22% of 3, 12% of 4 and more

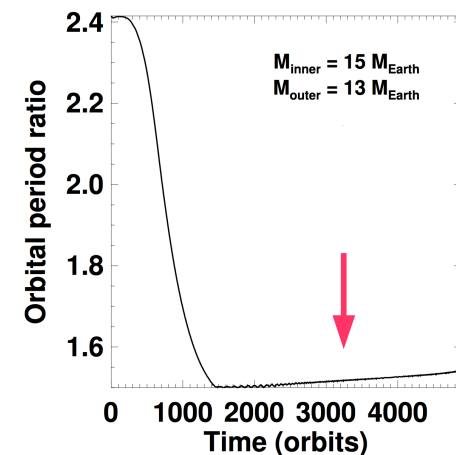
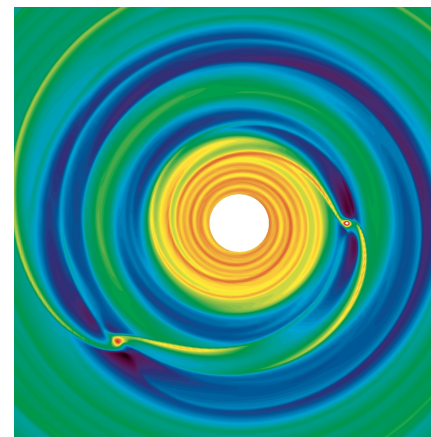


data extracted from
exoplanets.org (06/2014)

□ Many planet pairs are *not* in resonance, but those near resonances tend to have period ratios slightly **greater** than resonant

MODELS

- In-situ growth of planet embryos
- Tidal dissipation of close-in resonant planetary systems
- Disc-migration of partial gap-opening planets
Baruteau & Papaloizou 2013



Formation and evolution of planetary systems: what have we learnt from transit methods?

- Many mechanisms contribute to the orbital evolution of planetary systems

 - disc and high-eccentricity migrations, interactions with host and nearby stars all play some role.

- Importance of stellar evolution in planetary evolution

- More observations to constrain evolution models

 - CHEOPS (2017), TESS (2017), PLATO (2024)...